

NUCLEAR PHYSICS 16

Tensor forces



One boson exchange model



RECAP

Previous lesson \rightarrow T-matrix (Review)

(*) $f(\omega) = -\frac{\mu}{2\pi} \langle \vec{k} | T(E) | \vec{k} \rangle, \frac{d\sigma}{d\omega} = |f(\omega)|^2$

$$T(E) = V + V G_0(E) T(E) \quad \text{Lippmann}$$

- Schwinger
equation

- (*) T-matrix is easy to solve
for a particular type of potentials
(separable potentials)

$$\langle \vec{p}' | V | \vec{p} \rangle = \lambda g(p)g(p')$$

$$\downarrow T = U + V \delta_0 T$$

$$\langle \vec{p}' | T(E) | \vec{p} \rangle = Z(E) g(p)g(p')$$

(definition of
a separable
potential)

$$Z(E) = \frac{1}{1/\lambda - J(E)}$$

$$J(E) = \int \frac{d^3 q}{(2\pi)^3} \frac{g^2(q)}{E - \frac{q^2}{2\mu}}$$

Lippmann-Schwinger
equation
(for T-matrix)



Separable
potential



Simple
T-matrix

→ This is also useful

for contact-range theories

universal low-energy theories

Contid-range |

$$\langle \vec{r}'' | V | \vec{r} \rangle = C_0(\Lambda) f(\frac{\vec{r}'}{\Lambda}) f(\frac{\vec{r}}{\Lambda})$$

$$\rightarrow \left\{ \begin{array}{l} f(x) \rightarrow 1, x \rightarrow 0 \\ f(x) \rightarrow 0, x \rightarrow \infty \end{array} \right. |$$

Kill the unwanted
high-energy modes

V depends on Λ

$$C_0(\Lambda) \rightarrow Z(\Lambda)$$

$$\boxed{\frac{dZ(\Lambda)}{d\Lambda} \leq 0}$$

Example \rightarrow $\sigma \rightarrow 4\pi |a_0|^2$ } $E \rightarrow 0$ } $\Rightarrow \frac{dE}{dn} + \frac{d}{dn}(\sigma E) = 0$

$$\Rightarrow \frac{1}{C_0(n)} = \frac{\mu}{2\pi} \left(\frac{1}{a_0} - \beta n \right)$$

Family of
solutions
that we will get



$$P(L(x)) = \Theta(1-x) \Rightarrow \beta = \frac{N}{F}$$

(sharp cutoff)

* Bound states w/ T-matrix

$$T(E) \rightarrow \frac{R_{\text{est}}(E_B)}{E - E_B}, \quad R_{\text{est}} = V|B\rangle\langle B|V$$

$$= G_0^{-1}(E_B)|B\rangle\langle B|G_0^{-1}(E_B)$$

$$\langle \vec{p}' | V | \vec{p} \rangle = \phi_0(\vec{p}) \phi_0^*(\vec{p}') \quad S \quad A \rightarrow S \quad \text{and fixed}$$

$$f_B(\vec{p}) = \langle \vec{p} | B \rangle = \frac{\sqrt{8\pi\gamma}}{p^2 + \gamma}, \quad \gamma = \frac{1}{\alpha_0}$$

→ These were the most important
points to remember
about the Tumultus 1

Tensor Force | What is γ ?

3) Yukawa's idea to explain nuclear forces

$$V_y(\vec{r}) = -g_y^2 \frac{e^{-mr}}{4\pi r} \quad \begin{matrix} \text{meson} \\ \cdots \cdots \end{matrix} \quad (J\Gamma = \sigma^+)$$

But \exists a problem \rightarrow deuteron quadrupole moment

Deuteron quadrupole moment $\rightarrow Q_d \approx 0.23 \frac{\text{fm}^2 e}{\hbar}$

$$Q_d \neq 0$$

\Rightarrow Deuteron contains

a D-wave component

$$Q_d \propto (3z^2 - r^2) \Rightarrow$$

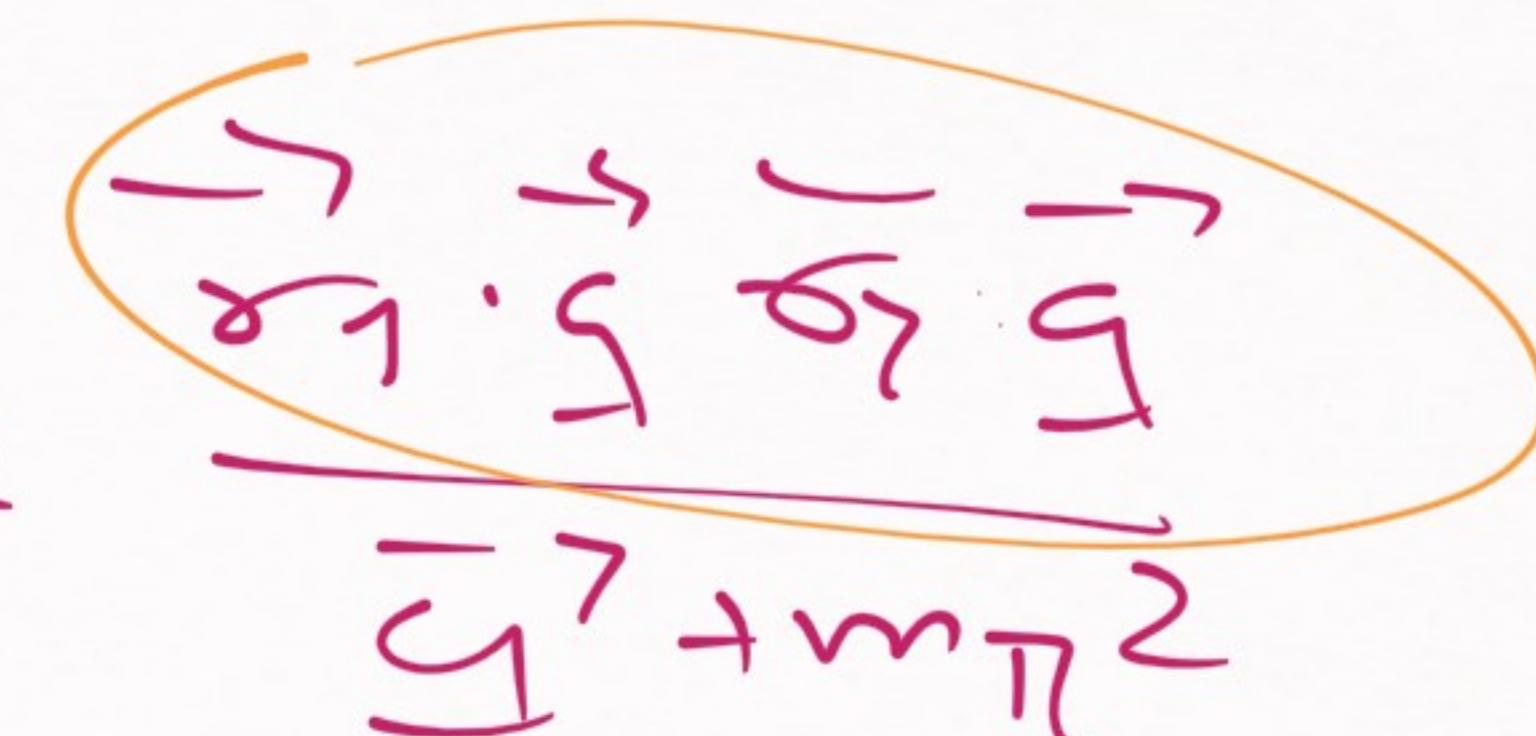
\vec{J}_d / S-wave ($\psi(\vec{r}) = \psi(r)$)

$$\left[\Rightarrow \text{D-wave component} \right] \Leftrightarrow \langle \psi | Q_d | \psi \rangle = 0$$
$$\left(\langle z^2 \rangle - \frac{1}{3} \langle r^2 \rangle \right)$$

2) Usual explanation is $\alpha J^P = \sigma$ meson exchange

$OPE \rightarrow$ One pion exchange (pion)

$$V_{\sigma \rightarrow \pi}(\vec{q}) = - \frac{g_\sigma \vec{\gamma}}{4f_\pi^2} \vec{\gamma}_1 \cdot \vec{\gamma}_2$$



$$\frac{\sigma_1 \cdot \vec{\gamma} \sigma_2 \cdot \vec{\gamma}}{\vec{\gamma}^2 + m_\pi^2} eD \odot$$

$$[g_\sigma \approx 1.26, f_\pi \approx 92.4 \text{ MeV}, m_\pi \approx 133 \text{ MeV}]$$

$$\Theta \Rightarrow \vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q} = \sigma_{1i} \sigma_{2j} q_i q_j \Rightarrow \underline{\underline{q_i q_j}}$$

$$\Theta = \sigma_{1i} \sigma_{2j} \left[\underbrace{q^2 \delta_{ij}}_{\text{O5}} \right] + \left(q_i q_j - \frac{1}{3} q^2 \delta_{ij} \right) = \Theta$$

$$1S1 = O_5 \oplus 1_A \oplus 2_S \quad \left. \begin{array}{l} q_i q_j = q_j q_i \rightarrow S \\ = \end{array} \right\}$$

$$\Theta = \frac{1}{3} \vec{\sigma}_1 \cdot \vec{\sigma}_2 q^2 + \left(3 \vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q} - \frac{1}{3} \vec{\sigma}_1 \cdot \vec{\sigma}_2 q^2 \right)$$

$$\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q} = \frac{1}{3} \vec{\sigma}_1 \cdot \vec{\sigma}_2 \vec{I} + (\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \vec{q} - \frac{1}{3} \vec{q} \vec{\sigma}_2 \vec{q})$$

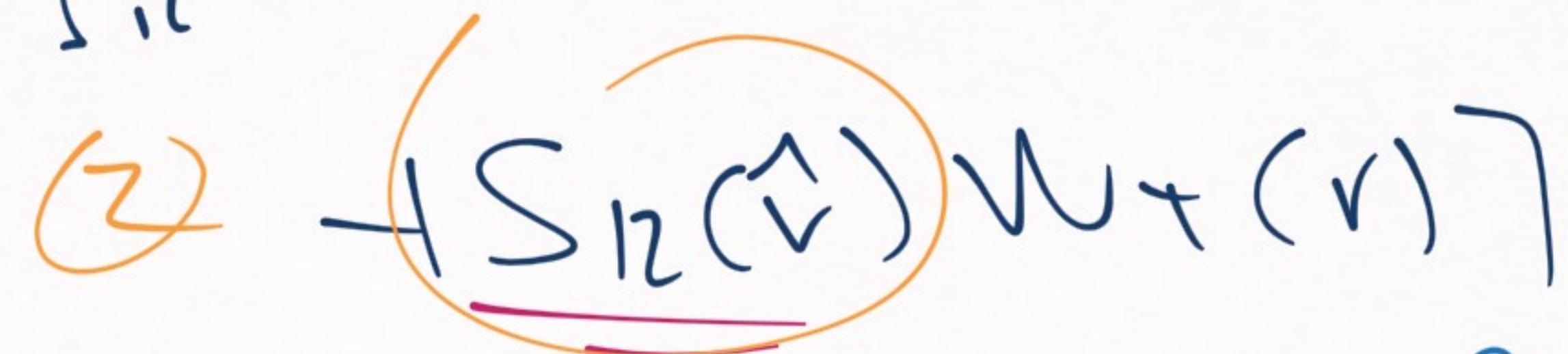
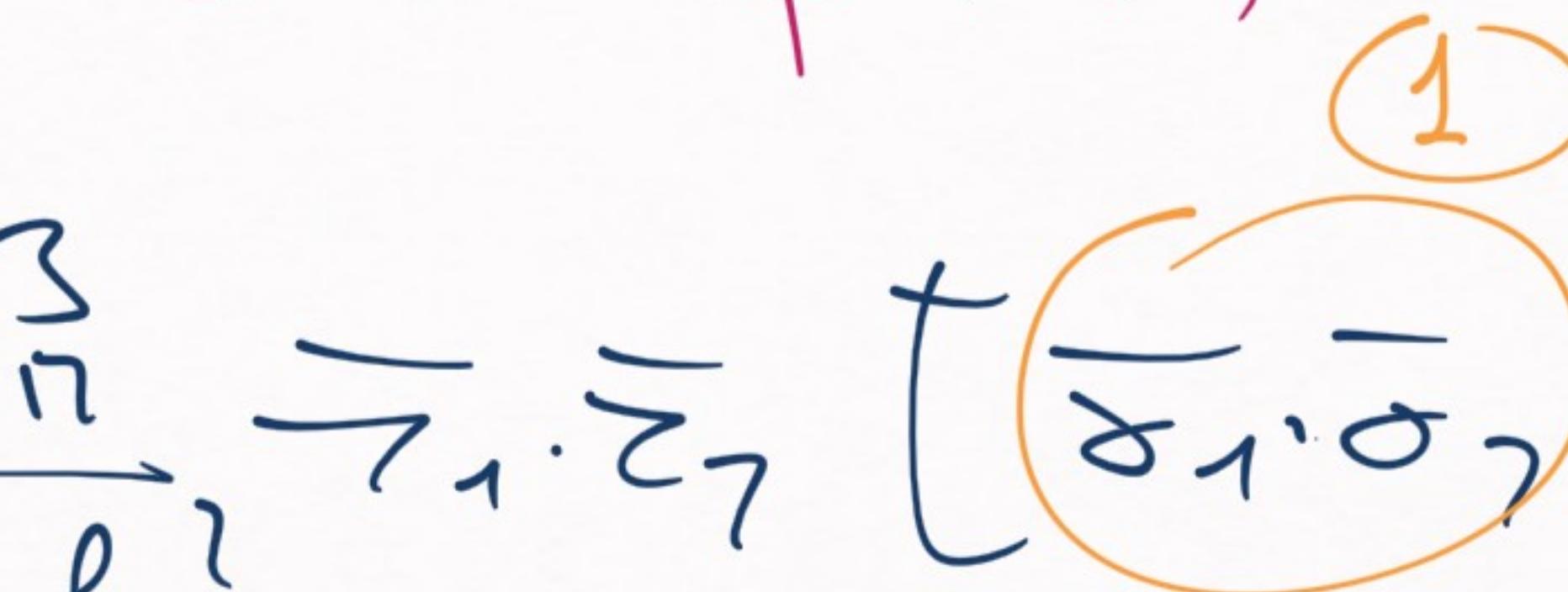
spin-spin
operator
 $(L=0)$

tensor operator
 $(L=2)$

\Downarrow
Detector will have
an $L=2$ component

(configuration space] (r-space)

$$V_{OBE}(\vec{r}) = \frac{g_A^2 m_n^3}{48\pi f_\pi^2} \bar{\tau}_1 \cdot \bar{\tau}_2 [+ \underbrace{(S_{12}(\hat{v}))}_{(2)} W_+(r)]$$



$$[S_{12}(\hat{v}) = 3 \bar{\sigma}_1 \cdot \hat{r} \bar{\sigma}_2 \cdot \hat{v} - \bar{\sigma}_1 \cdot \bar{\sigma}_2] \text{ usual definition of tensor operator}$$

① (Spin-Spin) $\vec{\sigma}_1 \cdot \vec{\sigma}_2 = \begin{cases} 1 & S=1 \\ -3 & S<0 \end{cases}$

② (tensor) $S_{12}(\vec{r}) \rightsquigarrow \underline{[S_{12}, \vec{r}^2]} \neq 0$

a) $\Rightarrow S_{12}$ mixes different values of L

$$[S_{12}, \vec{r}^2] = 0^{(*)} \quad \left. \begin{array}{l} b, d \\ c \end{array} \right\}$$

b, c) $\Rightarrow S_{12}$ conserves S_{12}

$$[S_{12}, \vec{j}^2] = 0$$

(*) only happens if $|S| < 2$ Always happens in NN

$S_{12}(\vec{r}) \rightarrow$ [matrix in \mathcal{C}]
scalar in s_{ij}

→ (when computing its matrix elements)

$$\langle (s'e')j'm' | S_{12}(\vec{r}) | (se)jm \rangle = \sum_{\ell\ell'}^{\infty} \delta_{jj'} \delta_{mm'} \delta_{ss'}$$

We can drop this, because $s=1$ for $\langle S_{12} \rangle \neq 0$

Additional simplification:

$$\underline{S=0} \Rightarrow \langle S_{12} \rangle = 0 \quad \text{Why?}$$

$$S_{12}(r) = 3 \vec{\sigma}_1 \cdot \hat{r} \vec{\sigma}_2 \cdot \hat{r} - \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

If we exchange spins: $S_{12} \rightarrow S_{17}$

$$S=0 \rightarrow |100\rangle = \frac{1}{\sqrt{2}}(|+-\rangle - |-+\rangle) \quad |100\rangle \rightarrow -|100\rangle$$

$$\langle S_{12} | 100 \rangle = -\langle S_{12} | 000 \rangle \Rightarrow \langle S_{12} | 100 \rangle = 0$$

(this explanation is a bit wacky)

. .

$$(\zeta=1) \Rightarrow \langle S_{12} \rangle \neq 0 \quad (\zeta \neq \zeta' \neq \delta)$$

$$\int d^2 \vec{r} \left(3 \vec{\sigma}_1 \cdot \vec{r} \vec{\sigma}_2 \cdot \vec{r} - \overline{\vec{\sigma}_1} \overline{\vec{\sigma}_2} \right) = 0$$

$\downarrow \overline{\vec{\sigma}_1} \overline{\vec{\sigma}_2}$

$\zeta = 0$
 $\zeta' = 0$

$$\langle s=0 | S_{12} | s=0 \rangle = 0$$

$$\langle \rho=0 | S_{12} | \rho=0 \rangle = 0$$

— . —

How to calculate $S_{12}(\vec{r})$?

$[S_{12}(\vec{r}), \vec{j}] = 0 \rightarrow$ Define states w/
good j

$|s\rangle_j m\rangle$

no/tensor force \rightarrow

$$|E_{\text{me}}(s m_s)\rangle = \underbrace{\sum_{\ell m_\ell} Y_{\ell m_\ell}(\hat{r})}_{\text{new part}} |s m_s\rangle$$

with tensor force $\rightarrow |(\ell s)_{jm}\rangle$

$$|(\ell s)_{jm}\rangle = \sum_{\substack{s m_s \\ \ell m_\ell}} \underbrace{Y_{\ell m_\ell}(\hat{r})}_{\text{Clebsch-Gordan coefficient}} |s m_s\rangle \times \langle \ell m_\ell, s m_s | j m \rangle$$

Clebsch-Gordan
coefficient

(new part)

$$\langle (s' e') j' m' | \zeta_{17}(r) | (s e) j m \rangle$$

$$= \langle \bar{s} e' | \delta_{ss'} \zeta_w | s' n m' \rangle$$



assumes $s < 2$



(which is true for
the NN system)

How does this look for the deuteron?

w/o tensor force

$$\psi_d(\vec{r}) = \frac{u(r)}{\sqrt{v}} \Sigma_{\alpha} \langle \vec{r} | \downarrow \downarrow m_d \rangle$$

$$|\downarrow \downarrow\rangle = |\leftarrow \leftarrow\rangle$$

(very simple
wave function)

$$|\downarrow \uparrow\rangle = \frac{1}{\sqrt{2}}(|\leftarrow \rightarrow\rangle + |\rightarrow \leftarrow\rangle)$$

$$|\uparrow \downarrow\rangle = |\rightarrow \leftarrow\rangle$$

w/ tensor wave

$$\psi_d(r\vec{r}) = \frac{U(r)}{r} \sum_{dd}(\vec{r}) |1smd\rangle$$

$$|1s m_s\rangle \langle s m_s 001 \{m_d\}| = |1s m_s\rangle$$

$$+ \left[\frac{w(r)}{\sqrt{2}} \sum_{ms1m1} \sum_{ms}(\hat{v}') |1s m_s\rangle \right]$$

$$\langle z_{m_1} 1_{m_1} | 1_{m_d} \rangle$$

D-Wave component

$$\psi_d(\vec{r}) = \frac{u(r)}{\sqrt{r}} |S(1md)\rangle + \frac{w(r)}{\sqrt{r}} |D(1md)\rangle$$

$$= D \langle S_{12}(\vec{r}) \rangle = \begin{pmatrix} \langle S_1 S_{12} S_2 \rangle & \langle S_1 S_{12} D_2 \rangle \\ \langle D_1 S_{12} S_2 \rangle & \langle D_1 S_{12} D_2 \rangle \end{pmatrix}$$

$$= \begin{pmatrix} 0 & 2\sqrt{2} \\ 2\sqrt{2} & -2 \end{pmatrix}$$

For the deuterium:

$$S_{12} = \begin{pmatrix} 0 & z\sqrt{z} \\ z\sqrt{z} & -2 \end{pmatrix}$$

Knowing this, we can write a Schrödinger equation for $u(r), w(r)$

w/o tensor : $V(r) = V_C(r)$ (might include spin-spin)

$$[-V''(r) + 2\mu V(r) V(r) = -\gamma^2 V(r)]$$

It was a very simple differential equation

w/tensor $V(\vec{r}) = V_C(r) + S_{12}(\vec{r}) V_T(r)$

$$-U''(r) + 2\mu [V_C(r) U + 2\sqrt{2} V_T(r) W] = -\gamma^2 U(r)$$

$$-W''(r) + \frac{6}{r} W(r) +$$

$$2\mu [2\sqrt{2} V_T(r) U + (V_C - 2\sqrt{2}) W(r)] = -\gamma^2 W(r)$$

$$(0, 2\sqrt{2}, 2\sqrt{2}, -2)$$

More compact in matrix form.

$$-\begin{pmatrix} v \\ w \end{pmatrix}'' + 2\mu \begin{pmatrix} v_c & 2\sum V_i \\ 2\sum V_i & (v_c - 2V_1) \end{pmatrix} \begin{pmatrix} v \\ w \end{pmatrix}$$

$$+ \begin{pmatrix} 0 & 0 \\ 0 & \ell(\ell+1)/r^2 \end{pmatrix} \begin{pmatrix} v \\ w \end{pmatrix} = -\gamma \begin{pmatrix} v \\ w \end{pmatrix}$$



$\frac{\ell(\ell+1)}{r^2} \rightarrow$ centrifugal barrier

this also changes the asymptotic behavior

$$v(r) \rightarrow A_s e^{-\gamma r} \quad] \rightarrow \text{dominant comp.}$$

$$w(r) \rightarrow A_D e^{-\gamma r} \left(1 + \frac{3}{\gamma r} + \frac{3}{(\gamma r)^2} \right)$$

$$r \rightarrow \infty$$

$$\frac{A_D}{A_s} \underset{=} \approx 0.0256$$

Normalization \rightarrow

$$\int_0^D v^2(r) dr = 1 \Rightarrow \int_0^D (v^2(r) + w^2(r)) dr = 1$$

Observables \rightarrow

$$\begin{aligned} \langle r_m^2 \rangle &= \frac{1}{\zeta} \langle r^2 \rangle \\ &= \frac{1}{\zeta} \int_0^\infty r^2 v^2(r) dr \quad \Rightarrow \quad \langle r_m^2 \rangle = \frac{1}{\zeta} \times \\ &\quad \times \int_0^\infty r^2 (v^2(r) + w^2(r)) dr \end{aligned}$$

Quadrupole moment:

$$\langle Q_d \rangle = 0 \Rightarrow \langle Q_d \rangle = \frac{1}{20} \int_0^S (r^2 w(r) G \times (2\sqrt{2}U(r) - w(r))) dr$$



$$-TQ \quad w(r) = 0 \Rightarrow Q_d = 0$$

D-wave probability:

$$P_D = \int_0^{\infty} w^2(r) dr$$

↗ D

$$P_D \sim (3-S)\gamma_0$$

(only an observable
until we include
quantum corrections
to e.m.
observables)

W

Review of deuteron observables;

$$r_m = 1.97 \text{ fm}$$

$$Q_d = 0.2859(3) \text{ fm}$$

$$P_D \sim (3.5)^\circ$$

$$\eta = \frac{\Delta b}{\Delta S} = 0.6256(4)$$

≡

[TENSOR FORCE]

- More complicated deuteron
(now it contains a D-wave)
- also changes scattering amplitudes & phase shifts
(greatly complicate)

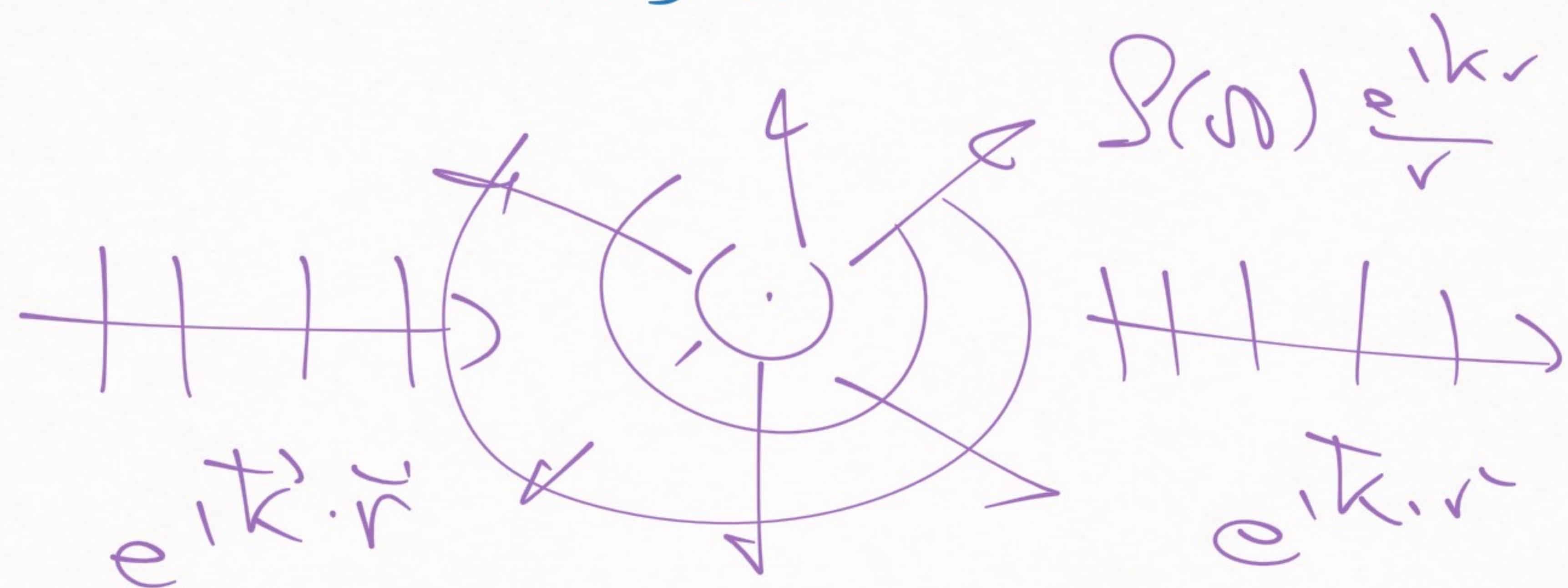
tensor force + scattering

→ very simplified overview

Effect of including spin in scattering

$$\text{no spin} \rightarrow \psi(\vec{r}) \rightarrow e^{i\vec{k} \cdot \vec{r}} + S(\sigma) \frac{e^{ikr}}{\sqrt{r}}$$

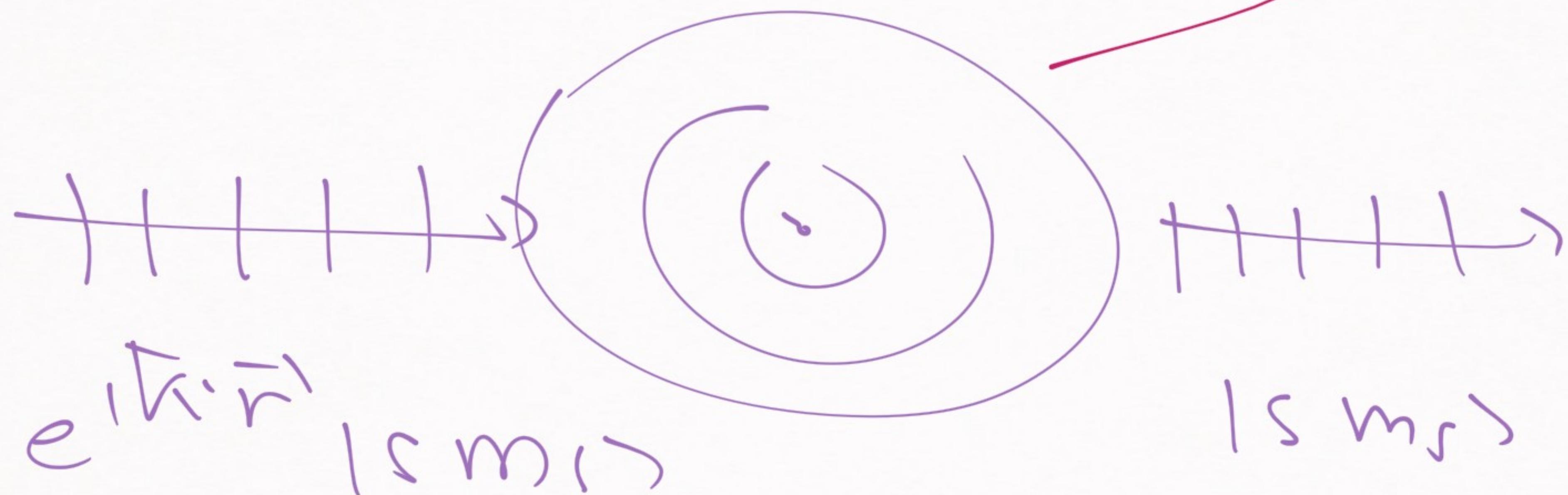
$r \rightarrow 0$



w/spin

$$\psi(\vec{r}) \rightarrow e^{i\vec{k} \cdot \vec{r}} |s m_s\rangle + P_{ms'm'}(\theta) \frac{e^{i\vec{k} \cdot \vec{r}}}{r} |s' m'_s\rangle$$

(assumes $s' = s$)



$$\text{w/o spin} \rightarrow P(N) = \sum_e (z e_+)^{\beta_0} P_0(\cos\theta)$$

$$P_0(k) = \frac{e^{2i\beta_0 - 1}}{z k} \xrightarrow{\text{phase shift}}$$

$$\text{w/ spin} \rightarrow P_{\text{wsm}}(N) = \sum_{je} (\dots) P_{se}^j(k) =$$

w_js_im_j
 e_{i,j}

$$f_{es}^j(\omega) = e^{\frac{z_i \zeta J_{sc}(k)}{z_k} - 1}$$

↳ this assumes $j=j', l'=l, s'=s$

If I break any of these assumptions

→ ρ will become a matrix

Tensor Force \rightarrow $e \neq e'$

$$\rho_{es(k)} = \frac{e^{2.5\sigma_{es}(k)} - 1}{z_{ik}} \rightarrow \rho_{ee}^j =$$

$$\frac{1}{z_{ik}} (S_{ee}^j(k) - S_{ee'}^j)$$

Ignore s because $s = s' = 1$

$$\rho_{ee'}^J = \frac{1}{Z(x)} (S_{ee'}^J(x) - \delta_{ee'})$$

$$\rho_e(u) = \frac{1}{Z_e(u)} (\zeta_e(u) - 1)$$

$$S_e S_e^\dagger = I$$

$$\Rightarrow \zeta_e(u) = e^{Z_e(u)}$$

$\zeta(u)^+ = 1$

$S_{op}^J(x)$
unitary
in L^2 -space

$$S^j(\kappa) = \left(\begin{array}{c} \text{ } \\ \text{ } \\ \text{ } \\ \text{ } \\ \text{ } \end{array} \right)$$

$\xrightarrow{\text{2x2}}$

$L = j \pm 1$

$S = 1, L = j \pm 1, j$

$\xrightarrow{\text{Sx1}}$

$L = j$

different parities

→ exactly identical to standard scattering

$$L = J \pm 1$$

→ diagonal matrix (unitary)

$$S(J, \kappa) = R D R^{-1} \quad \rightarrow \text{rotation matrix}$$

$$= \begin{pmatrix} \cos \kappa & -\sin \kappa \\ \sin \kappa & \cos \kappa \end{pmatrix} \begin{pmatrix} e^{2i\delta_1} & 0 \\ 0 & e^{2i\delta_2} \end{pmatrix} \begin{pmatrix} \cos \kappa + i \sin \kappa & -\sin \kappa \\ -\sin \kappa & \cos \kappa \end{pmatrix}$$

→ Eigen phase shifts

$$S(\kappa) = R(\kappa) D \ R^T(\kappa)$$

(see previous page)

→ intuitive parametrization

→ Nuclear bar

$$\underline{S}(\underline{k}) = \begin{pmatrix} e^{i\bar{\delta}_1} & 0 \\ 0 & e^{i\bar{\delta}_2} \end{pmatrix} \begin{pmatrix} \cos 2\bar{\epsilon}_j & \sin 2\bar{\epsilon}_j \\ \sin 2\bar{\epsilon}_j & -\cos 2\bar{\epsilon}_j \end{pmatrix}$$

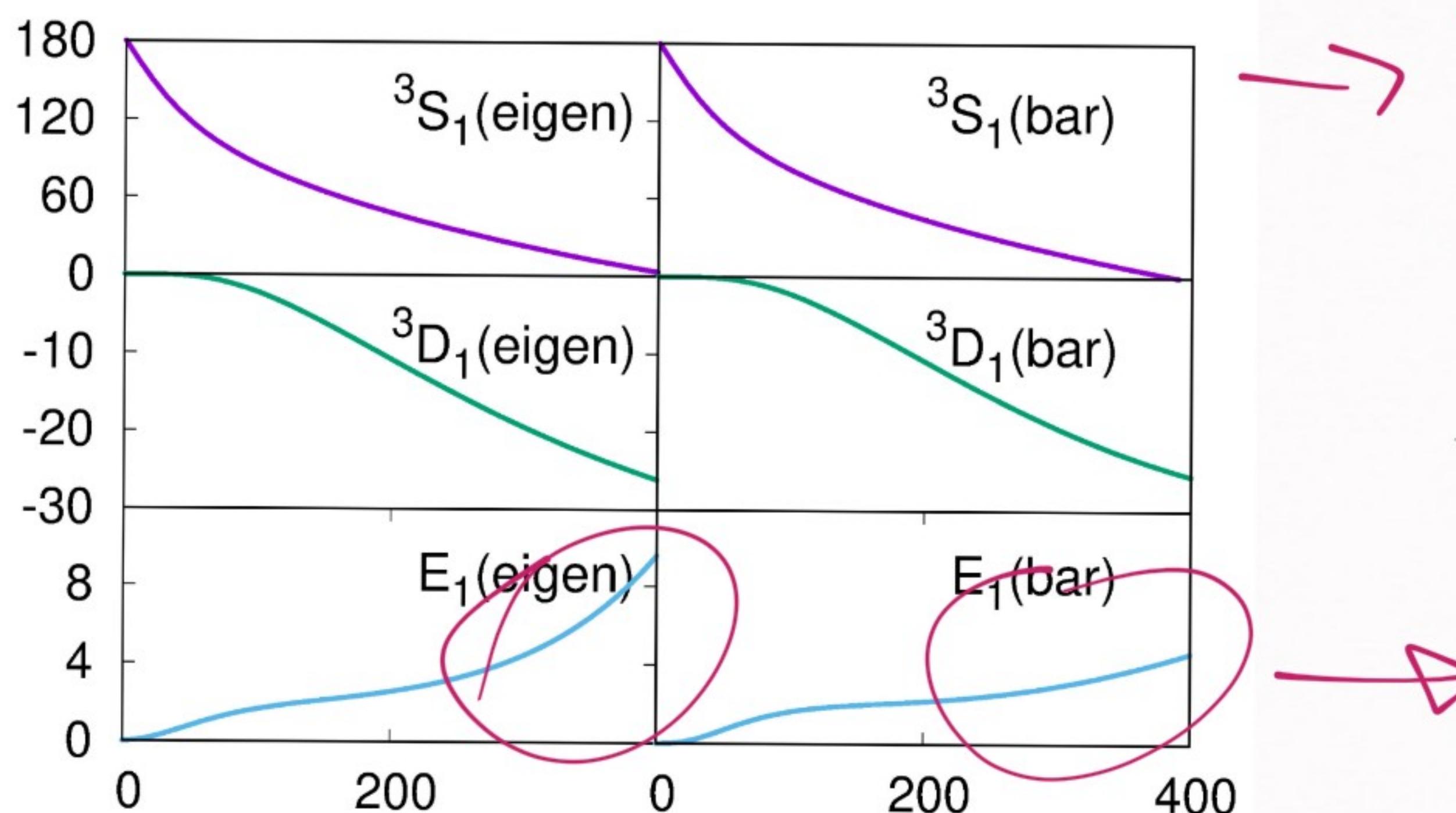


↑ the same,

$$\begin{pmatrix} e^{i\bar{\delta}_1} & 0 \\ 0 & e^{i\bar{\delta}_2} \end{pmatrix}$$

bit can be written in many forms

→ These are the most used parametrization
(as many as you want)



→ Deuteron channel
 $^3S_1 - ^3D_1$
only important difference

TO SUMMARIZE

→ tensor force creates a few
technical complications
for the phase shifts

(important if you study NN in detail)

— G —

———
[OBE MODEL]
———

We have seen a lot about the generic
description of the two nuclear systems

[One boson exchange (OBE) model]

First qualitatively successful description
of the nuclear force



→ very intuitive (direct extension of Yukawa's idea of meson exchange)

Historical notes

- 1) Before QCD
 - 1.a) Plan theories of the SD's → failed
 - 1.b) OBE model → works + easy
- 2) After QCD
 - 2.a) QCD-inspired models
 - 2.b) EFT methods (ongoing,
Scenes-to-work)

OBE model → easy-to-understand description

→ Why was it proposed?

1) Yukawa's idea: $\boxed{\dots \pi \dots}$ → First step in NN description

2) How to go beyond this?

2.a) Multi-pion theory

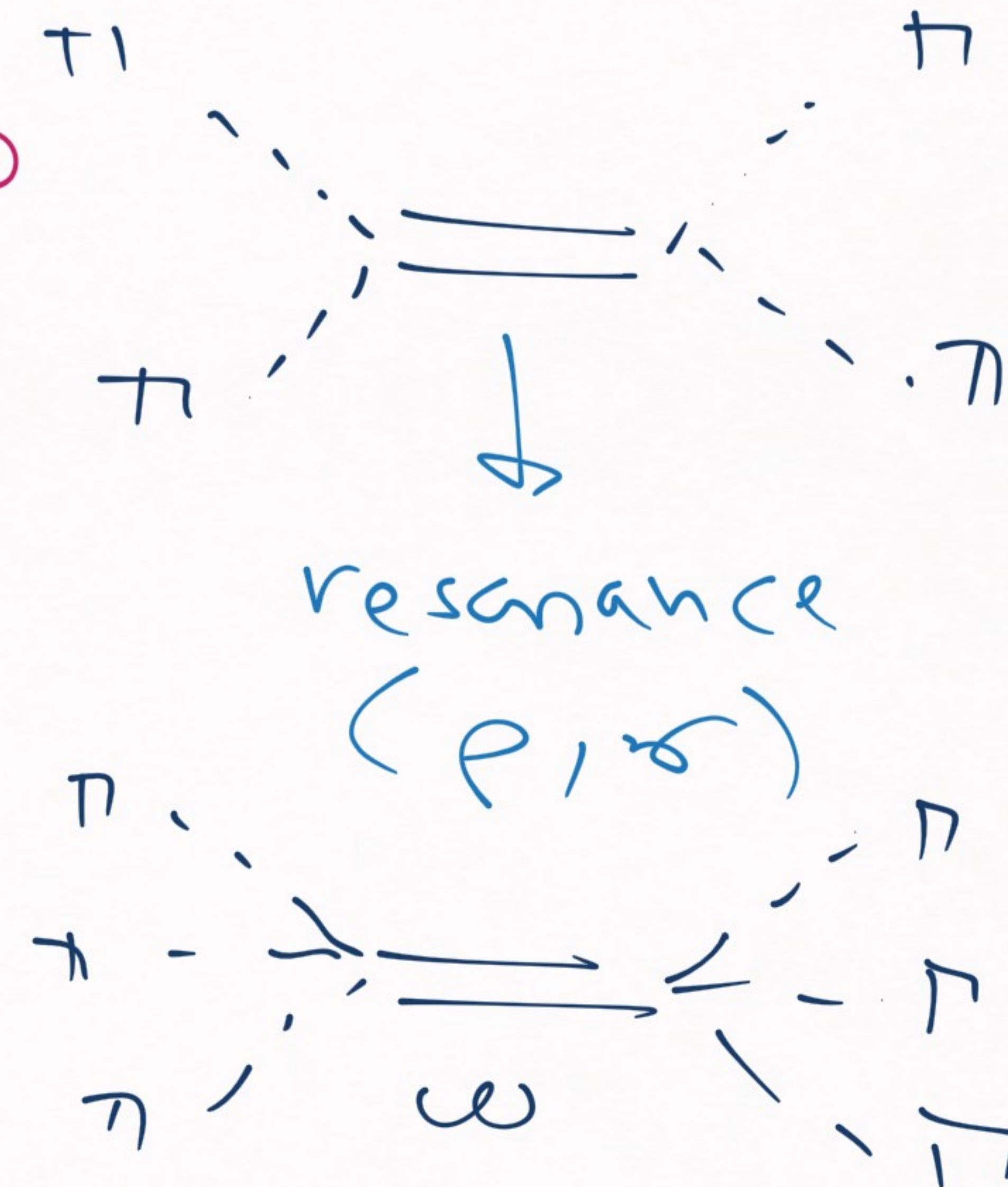
$$| \dots | \Rightarrow | \gamma_i |, k'_- , | \gamma_- |$$

more pions \rightarrow conceptually straightforward

In the 50's people didn't understand
pion dynamics / renormalization

→ Paired

2.b) ORC mode^P



$$|-\gamma^5 \gamma^7 \gamma^8 \gamma^9 \gamma^{10}| \rightarrow |\gamma^5 \gamma^9| \rightarrow ||$$

IDEA

reduce multi-pion exchange

→ exchange of a
single meson



RESULT OF THIS IDEA IS REALLY EASY:

$$\left[\begin{array}{c} \vdots \\ \vdots \\ \vdots \end{array} \right] \Rightarrow \left[\begin{array}{c} \cdots \\ \pi, \delta, \\ e, \omega \end{array} \right]$$

(+ sometimes a few
more)

Who are the important mesons?

1) The pion(π) : $J^P = 0^-$, $J=1$, $m_\pi = 140 \text{ MeV}$

(Job \rightarrow explain Q_d in the deuteron)

2) The sigma(σ) : $J^P = 0^+$, $J=0$, m_σ

$\approx 800 \text{ MeV}$

(Job \rightarrow strong mid-range attraction)

3) The rho(ρ): $J^P = 1^-$, $J = 1$, $m_\rho \Rightarrow 700\text{ MeV}$

(Jcb \rightarrow the tensor force of the pion is
too strong & the rho counter (π)

4) The omega(ω): $J^P = 1^-$, $J = 0$,
 $m_\omega = 780\text{ MeV}$

(Jcb \rightarrow provide short-range
repulsion)

GBE Potential:

$$V_{\text{GBE}} = V_T + V_C + V_W + V_\sigma + \dots$$

→ sum of individual contributions

→ Very simple description
of NN forces

$V_A(\vec{g}) \rightarrow$ (we have shown it before)

$$V_R(\vec{g}) = -\frac{g_0^2}{\vec{g}^2 + m_0^2}$$

$$V_W(\vec{g}) = \frac{g_W}{\vec{g}^2 + m_W^2} + \frac{(l_W + g_W)}{4M_N} \frac{(\vec{\gamma}_1 \wedge \vec{g}) \cdot (\vec{\gamma}_2 \wedge \vec{g})}{\vec{g}^2 + m_W^2}$$

$$V_F(\vec{g}) = \vec{\gamma}_1 \vec{\gamma}_2 \cdot [V_W / g_W \rightarrow g_F, l_W \rightarrow l_F / m_W \rightarrow m_F]$$

Some simplifications

$$\hookrightarrow g_e \gg g_r \Rightarrow g_r \approx 0$$

$$f_w \ll g_w \Rightarrow g_w \approx 0$$

\rightarrow Very common simplifications

[CORF potential] \rightarrow some contributions
to the potential
are singular

V_H, V_P, V_W

(Tensor forces)

$\sim \frac{1}{r^3}$ at short distances

Regulation \rightarrow form factors

Form factors | avoid $1/r^3$ singular
behavior



Since hadrons have a finite size,

we can include this effect

in the potential

$$V_M(\vec{g}) \rightarrow V_M(\vec{g}) F_M^2(\vec{g})$$

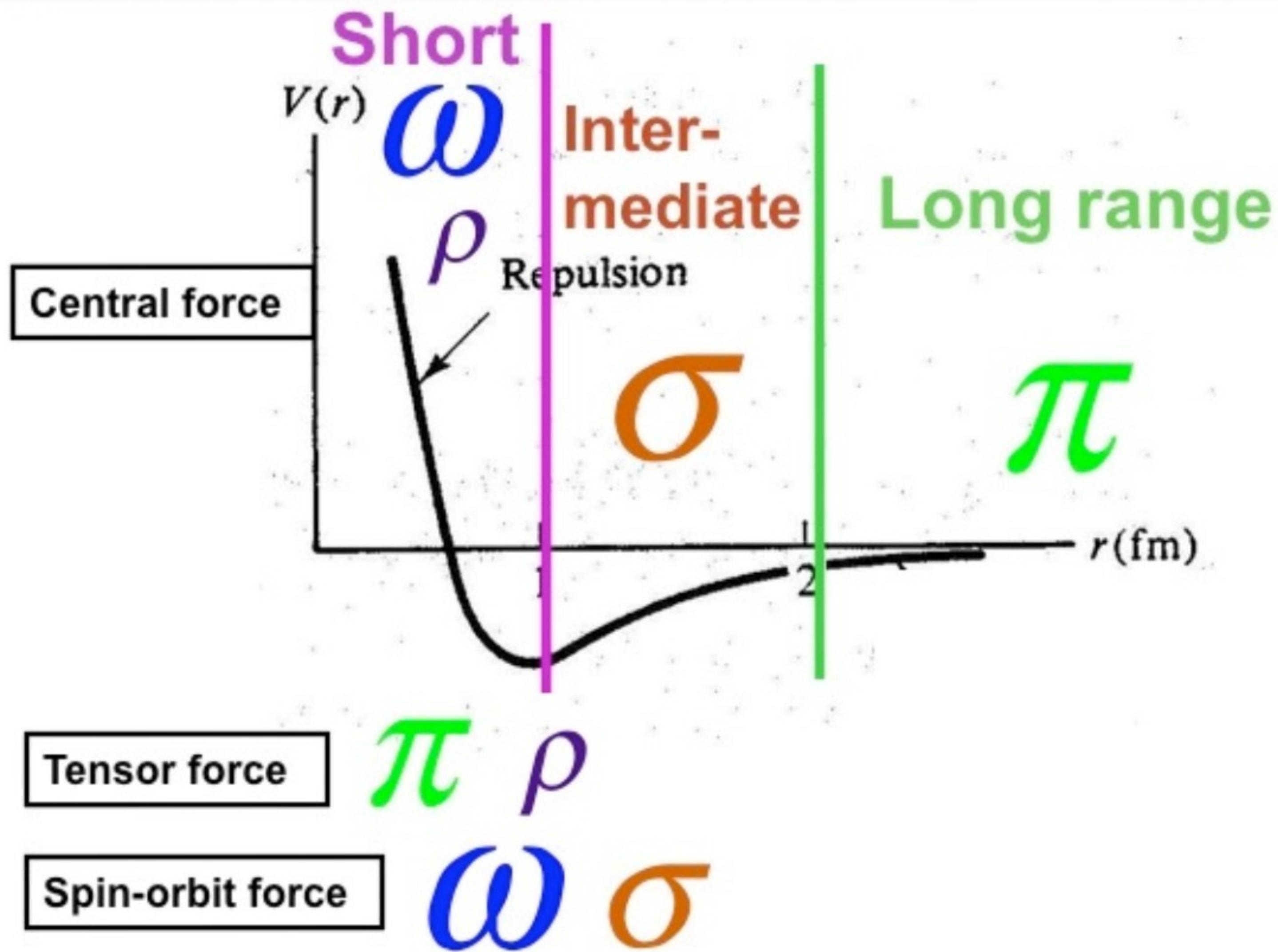
+ To account for
hadron finite-size

$$F_M(\vec{g}) = \left(\frac{R - m^2}{\Lambda^2 + \vec{g}^2} \right) \xrightarrow{\sim} \text{mult. polar form-factor}$$

⇒ does relativistic corrections

$\overbrace{(\delta, \omega)}$ → rel. corrections give rise
to a spin-orbit force

→ important for n-p wave



Summary
 of OBE
 model



Scholarpedia \rightarrow Nuclear Forces by McGehee

Table 1: The four most important mesons and the main characteristics of their contributions to components of the nuclear force.

Meson	Central	Spin-Spin	Tensor	Spin-Orbit
$\pi(138)$	--	weak, long-ranged	strong, long-ranged	--
$\sigma(500)$	strong, attractive, intermediate-ranged	--	--	moderate, intermediate- ranged
$\omega(782)$	strong, repulsive, short- ranged	--	--	strong, short-ranged, coherent with σ
$\rho(770)$	--	weak, short-ranged, coherent with π	moderate, short-ranged, opposite to π	--

\rightarrow Summary of the job done by
every meson //

→ In summary, this is an easy
and intuitive mode!



→ SEE YOU ON TUESDAY