

# NUCLEAR PHYSICS ⑨

## RENORMALIZATION



**RECAP**

① Renormalization means:

"Physics at long-distances does not depend on short-distance details"

**Example**

$$\left[ V(\vec{q}), m \right] \xrightarrow{|\vec{q}|^2 \ll m^2} V_{\text{eff}}(\vec{q}) = C_0 + C_2 \vec{q}^2 + C_4 q^4$$

$$\left[ \dots \right] \xrightarrow{|\vec{q}|^2 \ll m^2} -\frac{g_4^2}{m^2} + \text{corrections} + \dots$$

(Non-relativistic two-body problem)

$$\boxed{V_{\text{true}}(\vec{r})} \rightarrow \left[ V_{\text{eff}}(\vec{r}) = \alpha + \langle \vec{r} \rangle^2 + \langle \vec{r} \rangle^4 + \dots \right]$$

Characteristic  
scale  $m$

(e.g. Yukawa)

→ In most cases, I always  
obtain the same  
effective potential at  
low energies ~

→ Insensitivity to short-range details



$$V_{\text{eff}}(\vec{r}) = V_0 + C_2 r^{-2} + \dots$$

↓

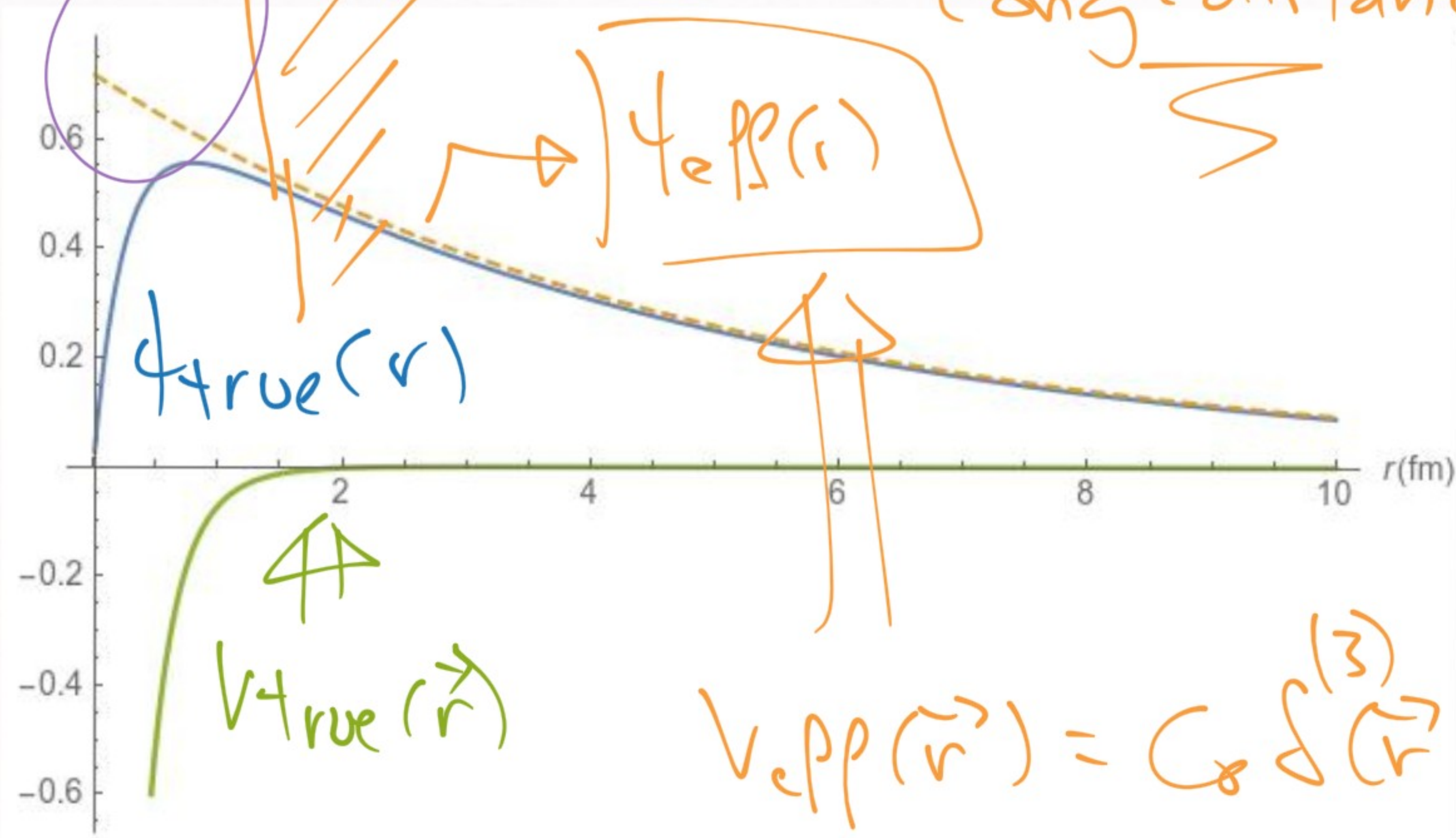
$$\Psi_{\text{true}}(\vec{r}) \xrightarrow{mr \gg 1} \Psi_{\text{eff}}(\vec{r}) = \frac{A_S}{\sqrt{4\pi r}} e^{-\gamma r}$$

For most situations, at low energies  
I can approximate the two-body system  
with

$$V_{\text{eff}}(\vec{s}) \quad \& \quad T_{\text{eff}}(\vec{s})$$

Short-distances

Long-distances



$$\psi_{app} \approx \psi_{true}$$

This contains the basic ideas

[Renormalization] → easy to understand at the conceptual level

the difficult part is implementation

(e.g. check previous calculation)

Why is difficult to implement renormalization at the mathematical level?

$$V_{\text{eff}}(\vec{q}) = C_0 + C_2 \vec{q}^2 + C_4 \vec{q}^4 + \dots$$

$$V_{\text{eff}}(\vec{r}) = C_0 \delta^{(3)}(\vec{r}) + C_2 \nabla^2 \delta^{(3)}(\vec{r}) + C_4 \nabla^4 \delta^{(4)}(\vec{r}) + \dots$$

How to use them?

~



$$C_0 \left[ \rho^{(3)}(\vec{r}) \right]$$

→ PROBLEM

VARIATIONAL PRINCIPLE

$H \rightarrow E_{\text{solution}}$  or  $E_{\text{sol}}$

$$\langle \psi_{\text{sol}} | H | \psi_{\text{sol}} \rangle \leq \langle \psi_{\text{t}} | H | \psi_{\text{t}} \rangle$$

↓  
true wf of a system minimizes the energy

$\psi_{\text{t}} \rightarrow$  trial wave function  
(anything we want)

$$V_{eff} = C_0 \delta^{(3)}(\vec{r})$$

$$\psi_{trial} = \frac{\Delta_5}{\sqrt{4\pi}} \frac{e^{-\gamma r}}{r}$$

$$\langle \psi_{trial} | V_{eff} | \psi_{trial} \rangle$$

$$\propto C_0 \times (\infty)$$

$$\text{if } C_0 < 0 \Rightarrow$$

$$\rightarrow -\infty$$

$$\langle \psi_{trial} | H | \psi_{trial} \rangle \rightarrow -\infty$$

$$\langle \psi_{sol} | H | \psi_{sol} \rangle$$

$$\leq \langle \psi_{trial} | H | \psi_{trial} \rangle$$

## PROBLEM

$$V_{eff} = C_0 \delta^{(3)}(\vec{r}) \text{ w/ } C_0 < 0$$

↪ no ground state

→ describes a two-body system  
that collapses

→ not what we want

SOLUTION



Two steps  
↳

1) REGULARIZATION

2) RENORMALIZATION



1) REGULARIZATION  $\rightarrow$  "make things regular"

$$V_{\text{eff}}(\vec{r}) = C_0 \left[ \delta^{(3)}(\vec{r}) \right]$$

REGULARIZE

singular,  $|\delta^{(3)}(\vec{0})| \rightarrow \infty$

$$\approx C_0 \left[ \delta^{(3)}(\vec{r}; R_c) \right]$$

$$|\delta^{(3)}(\vec{0}; R_c)| < \infty$$

→ Regularization can be done in many ways

$\delta^{(3)}(\vec{r}; R_c)$  1.a) finite at  $|\vec{r}| = 0$

1.b)  $R_c \rightarrow 0$ ,  $\delta^{(3)}(\vec{r}; R_c)$

$$\int d^3\vec{r} \delta^{(3)}(\vec{r}; R_c) = 1$$



→  $\delta^{(3)}(\vec{r})$



Gaussian regulator  $\rightarrow \delta^{(3)}(\vec{r}; R_c) = \frac{e^{-(r/R_c)^2}}{\pi^{3/2} R_c^3}$

Delta-shell regulator  $\rightarrow \delta^{(3)}(\vec{r}; R_c) = \frac{\delta(r - R_c)}{4\pi R_c^2}$

(Infinite possibilities)  
( $\delta$  is not singular in one dimension)

You can also do it in p-space

$$\begin{array}{ccc} \downarrow & \xrightarrow{\Lambda \text{ cutoff}} & \frac{\Lambda^2}{\Lambda^2 + g^2} \\ (\delta^{(2)} \text{ in } p\text{-space}) & & \left\{ \begin{array}{l} \Lambda \rightarrow \infty \\ \rightarrow 1 \quad \checkmark \\ g \rightarrow \infty \\ \rightarrow 0 \quad \checkmark \\ \text{(regular)} \end{array} \right. \end{array}$$
$$1 \xrightarrow{\quad} e^{-\left(\frac{g^2}{\Lambda^2}\right)}$$

(same features)





REGULARIZATION

→ Infinite possibilities



PROBLEM



The cutoff is not a physical parameter

$\Rightarrow$  ~~⊗~~

⊗ ⇒

The cutoff is a parameter  
of our theory  
(not of nature)

⇒

Results should not depend  
on the cutoff

⇒ ⊗

~~⊗~~ ⇒  $\mathbb{D}$  We need to achieve cutoff independence

## 2) RENORMALIZATION

$$V_{eff}(\vec{r}) = C_0 \delta^{(3)}(\vec{r}) \xrightarrow{\text{①}} C_0 \delta^{(3)}(\vec{r}, R_c)$$

$$\xrightarrow{\text{②}} \boxed{C_0(R_c) \delta^{(3)}(\vec{r}, R_c) + \text{something}} \Rightarrow \text{else} \text{⊗}$$

@ = D this something else is a condition

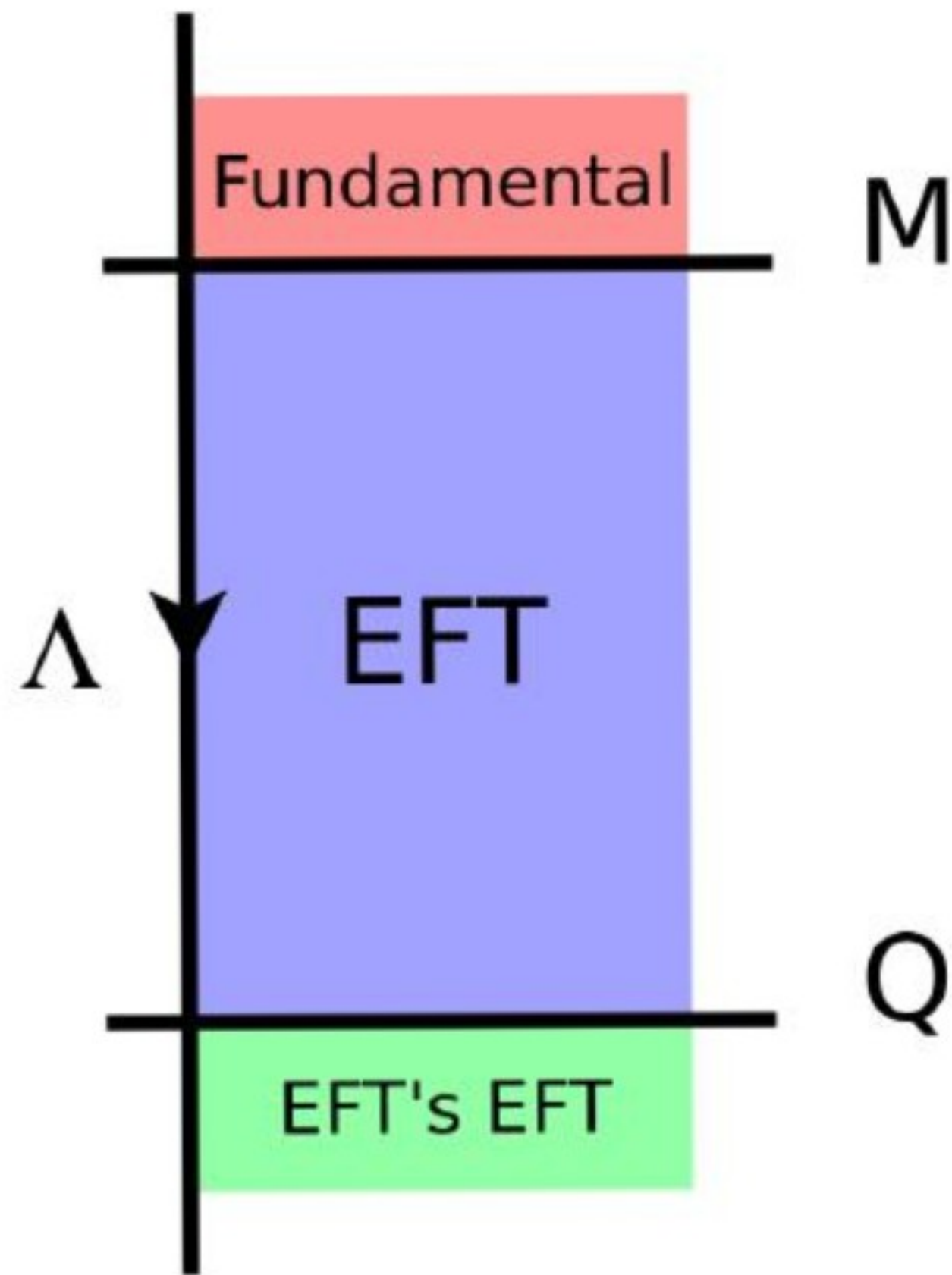
→ Results should not depend on  
the cutoff  $R_c / \Lambda$

↙  $\langle \psi | \hat{O} | \psi \rangle = \langle \hat{O}(R_c) \rangle$ ,  $\left[ \frac{d}{dR_c} \langle \hat{O}(R_c) \rangle = 0 \right]$

$$\Rightarrow \left[ \frac{d}{dR_c} \langle \hat{O}(R_c) \rangle = 0 \right]$$

It's abstract,  
but simple

$$\Rightarrow \left[ \begin{array}{l} C_0(R_c) \text{ must} \\ \text{have a specific} \\ \text{dependence} \\ \text{on } R_c \end{array} \right]$$



Physics is unique, but choice of theory depends on resolution  $\Lambda$ :

- ▶  $\Lambda \geq M$ : Fundamental
- ▶  $M \geq \Lambda \geq Q$ : EFT

For equivalent descriptions:

$$\frac{d}{d\Lambda} \langle \Psi | \mathcal{O} | \Psi \rangle = 0$$

Renormalization group invariance



$$\frac{d}{dR_c} \langle \hat{\mathcal{O}}(R_c) \rangle = 0$$

$$\frac{d}{d\Lambda} \langle \hat{\mathcal{O}}(\Lambda) \rangle = 0$$



How to do this?

→ Solve  $V_{\text{eff}}(\vec{r}, R_c) = C_0(R_c) f^{(3)}(\vec{r}; R_c)$

→ Obtain some observable  
as a function of  $C_0(R_c)$

→ We force  $\frac{d}{dR_c}(\ ) = 0$

# SIMPLE EXAMPLE 1

$$V_{\text{eff}}(r; R_c) = \frac{C_0(R_c)}{4\pi R_c^2} \delta(r - R_c) \rightarrow \text{bound state}$$

$$-\gamma = -\frac{1}{R_c} + \cancel{O(\delta R_c)} = Z_M \frac{C_0(R_c)}{4\pi R_c^2} \Rightarrow$$



$$\Rightarrow \gamma = -\frac{\lambda}{R_c} - \frac{2\mu}{4\pi R_c^2} C_0(R_c) \quad \left( \frac{\gamma'}{2\mu} = \beta \right)$$

$$\boxed{\frac{d\gamma}{dR_c} = 0}$$

$$\Rightarrow \frac{d}{dR_c} \left( \frac{\lambda}{R_c} + \frac{2\mu}{4\pi R_c^2} C_0(R_c) \right) = 0$$

$$\Rightarrow \boxed{C_0(R_c) \sim R_c}$$

# EXAMPLE

1)  $V_{eff}(\vec{r}; R_c) = C_0(R_c) \frac{|r - R_c|}{4\pi R_c^2}$

2) Binding energy/wave number  
is known ( $B$  or  $\gamma$ )

3) This forces  $C_0(R_c)$   
to depend on  $R_c$

We can show more examples

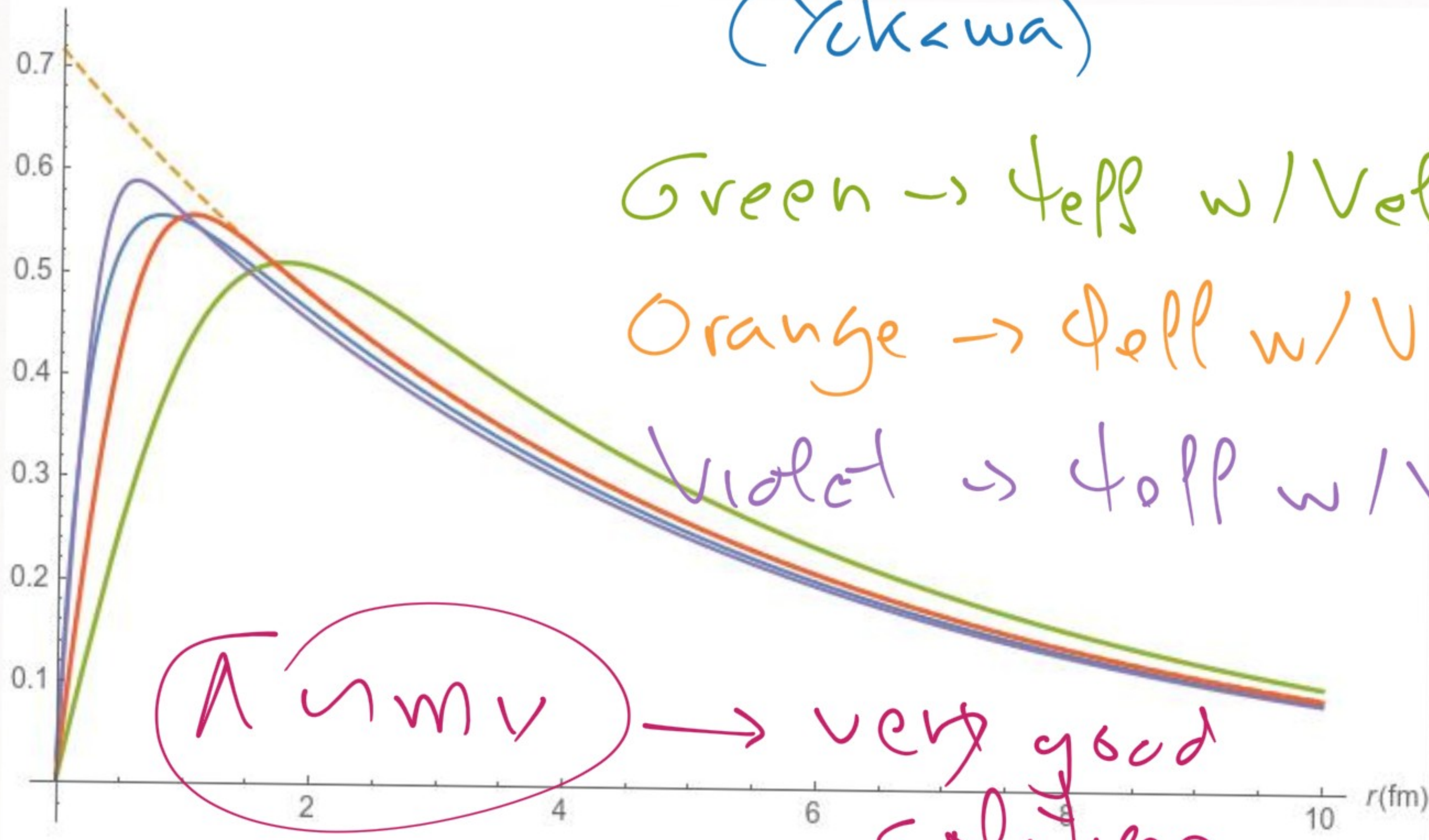


$$V_{\text{eff}}(\vec{q}; \Lambda) = C_0(\Lambda) e^{-\left(\frac{\vec{q}}{\Lambda}\right)^2}$$

w/  $C_0(\Lambda)$  such that it reproduces  
a given bound state

(Example  $\rightarrow$  Yukawa tax model from  
part (a))

Blue  $\rightarrow \phi_{true} \text{ w/ } V_{true}, \gamma \approx 0.075 \text{ m},$   
 (Yukawa)



Green  $\rightarrow \phi_{eff} \text{ w/ } V_{eff} \text{ and } \Lambda = \frac{m\gamma}{2}$

Orange  $\rightarrow \phi_{eff} \text{ w/ } V_{eff} \text{ and } \Lambda = m\gamma$

Violet  $\rightarrow \phi_{eff} \text{ w/ } V_{eff} \text{ and } \Lambda = 2m\gamma$

$\Lambda \approx m\gamma \rightarrow$  very good solution

TAKE-HOME MESSAGE

$$\omega(\Lambda) e^{-(g^2 \Lambda^2)} \longleftrightarrow \frac{g^2}{4\pi} \frac{e^{-m\gamma r}}{r}$$

SAME RESULTS IF  $\gamma \ll m$

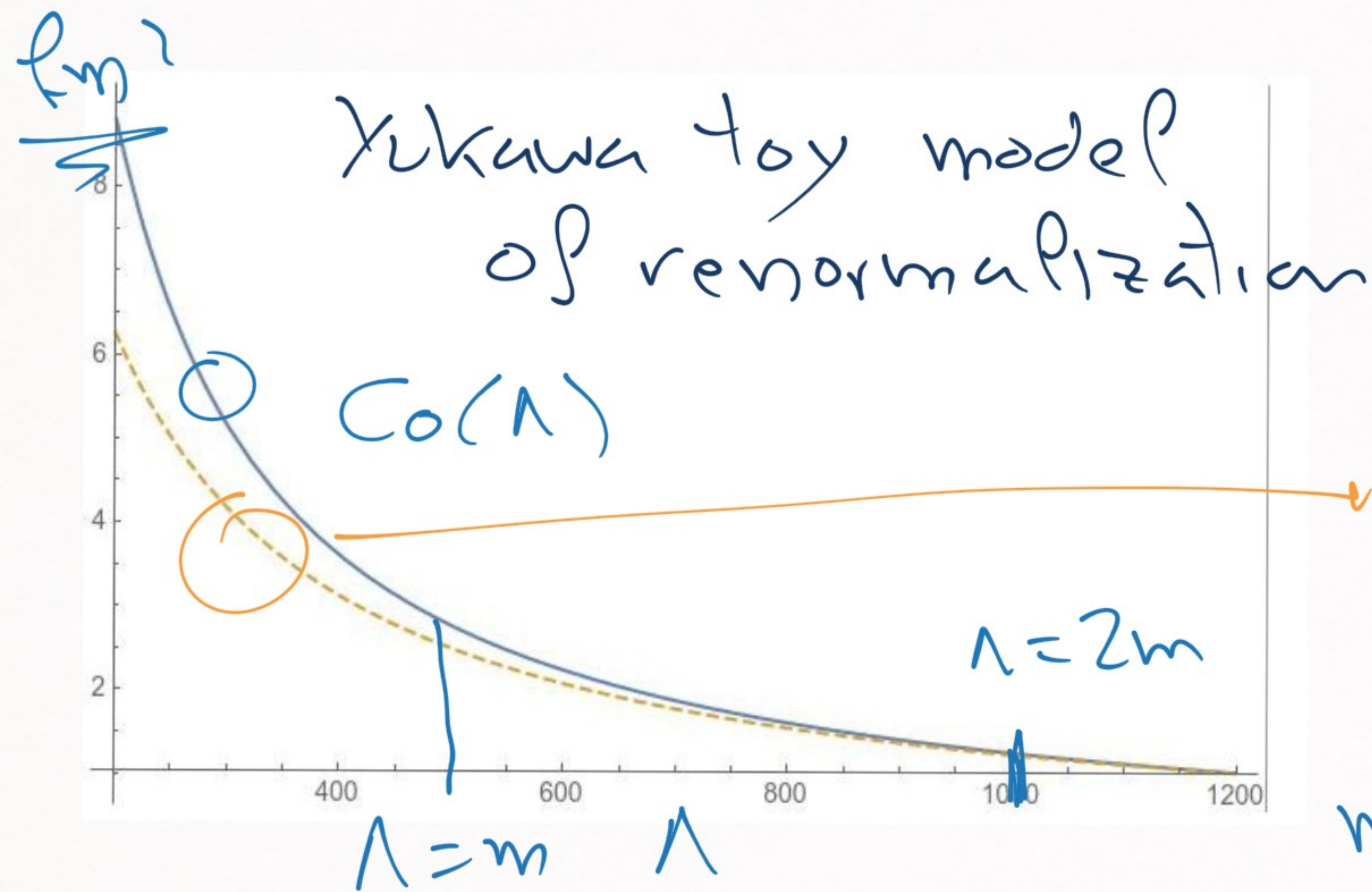
→ Insensitivity to short distance ( $m$ )

→ For most problems, it does not matter  
if we know the "true potential"

REASON → RENORMALIZATION



Renormalization  $\Rightarrow C_0(K_c)$  depends on the cutoff  $\frac{\Lambda}{m}$



Approx  $C_0(\Lambda) \sim \frac{1250}{\Lambda}$

$m = 500 \text{ MeV}$

→

$$\left[ \begin{array}{l} C_0(R_c) \sim R_c \\ C_0(n) \sim \frac{1}{n} \end{array} \right]$$

Fourier-transforms  
 $\rightleftharpoons$

(inverse of each other)

↓

$$\boxed{\text{RGE}}$$

(Renormalization group evolution) of the effective description  
 $\rightleftharpoons$   
 →



$$\text{RGE} \rightarrow C_0(R_c) \sim R_c \quad / \quad C_0(\Lambda) \sim \frac{1}{\Lambda}$$

RGE  
(Renormalization  
group  
equation)

$$\left. \right\} \rightarrow \frac{d}{dR_c} \left[ \frac{C_0(R_c)}{R_c} \right] \sim 0$$
$$\frac{d}{d\Lambda} [ \Lambda C_0(\Lambda) ] \sim 0$$

(so you know the language  
people use)

RECAP → A brief review of renormalization

1) Consider some physical system

1.a) "true potential" → not always necessary

1.b) Check the scales, see if an effective description is possible → if ok

2) Consider an effective interaction

$$V_c(\vec{r}) = c_0 \delta^{(3)}(\vec{r}) \quad , \quad V_c(\vec{q}) = c_0$$

### 3) Regularize & Renormalize

3.a) Regularize (include a cutoff)

$$V = C_0 \delta^{(3)}(\vec{r}) \rightarrow V = C_0(R_c) \frac{\delta(r - R_c)}{4\pi R_c^2}$$

3.b) Renormalize (determine  $C_0(R_c)$ )

from some observable quantity,

e.g. the binding energy)

4) Check if the description works

(if sb is ok, if  $\exists$  a good  
safe separation  $\Rightarrow$  it will work)

→ STEPS WE HAVE TO FOLLOW

→  $\left[ \exists \text{ several ways to formulate} \right]$   
renormalization  $\rightarrow \frac{G_0(R_c)}{N}$

a)  $G_0(R_c)$  from  $R_c$  of some observable  
(like binding energy)

b)  $G_0(R_c)$  from the RGE

$$\left( \frac{d}{dR_c} \langle \psi | \hat{O} | \psi \rangle = 0 \right)$$

Let's  
see  
this  
one

b) → more abstract, but can be useful  
in many situations



$$C_0(R_c) \frac{\delta(r-R_c)}{4\pi R_c^2} \left\{ \begin{array}{l} \text{a) } 2\mu \frac{C_0(R_c)}{4\pi R_c^2} = -\gamma - \frac{1}{R_c} \\ \text{b) } \frac{d}{dR_c} \langle 4 | H | 4 \rangle = 0 \quad \leftarrow \text{second} \\ \text{new} \\ \downarrow \\ \frac{d}{dR_c} [R_c^9 C_0(R_c)] = 0 \end{array} \right.$$

RGE  $\rightarrow$  Problem is usually to transform

$$\frac{d}{dR_c} \langle \psi | \hat{O} | \psi \rangle = 0 \quad \text{into} \quad \frac{d}{dR_c} [R_c^a C(R_c)] = 0$$

(a is some power)  
M

(Simplification)  $\rightarrow$

$$\left[ \frac{d}{dR_c} \langle \psi | H | \psi \rangle = 0 \right] \rightarrow \frac{d}{dR_c} \langle \psi | V | \psi \rangle = 0$$

$$E = \langle \psi | H | \psi \rangle$$

=

- B



$$V_c(\vec{r}) = G_0(R_c) \frac{\delta(r-R_c)}{4\pi R_c^2}$$

$$\frac{d}{dR_c} \langle \psi | V | \psi \rangle = 0$$

$$\langle \vec{r} | \psi \rangle = \frac{\Delta_s}{\sqrt{4\pi}} \frac{e^{-\gamma r}}{r}$$

(For  $r > R_c$ )

$$\langle \psi | V_c | \psi \rangle = G_0(R_c) \int d^3\vec{r} \frac{\delta(r-R_c)}{4\pi R_c^2} |\langle \vec{r} | \psi \rangle|^2$$

= 0

$$\theta \approx C_0(R_c) | \psi(R_c) |^2$$

$$\langle \psi | V | \psi \rangle \approx C_0(R_c) | \psi(R_c) |^2$$

$$\frac{d}{dR_c} \langle \psi | V | \psi \rangle = \frac{d}{dR_c} [C_0(R_c) | \psi(R_c) |^2] \leq 0$$

$$\psi(R_c) \approx \frac{1}{R_c} \quad (\text{check previous slide})$$

With this simplification:

$$\frac{d}{dR_c} \left[ \frac{C_0(R_c)}{R_c^2} \right] \approx 0 \Rightarrow C_0(R_c) \approx R_c^2$$

Not completely correct (previously we

(we are ignoring  $T$ )

but it's quite good  $\approx$

saw that  
 $C_0(R_c) \approx R_c^2$ )

$\frac{d}{dt} \langle 41014 \rangle \approx 0$  → quick & dirty  
trick  
to derive the RTE  
of  $\langle \rho \rangle$

CAVEAT: IT WILL FAIL

FOR NON-PERTURBATIVE

PROBLEMS (BECAUSE IT IGNORES  $\langle T \rangle$ )

→ We will use this trick to explore  
 the existence of different ~~GFs~~  
 for the same problem



"Cup"  
 power  
 counting  
 $\downarrow$

$$\chi = \sum (6 + r_1 \chi + \dots)$$

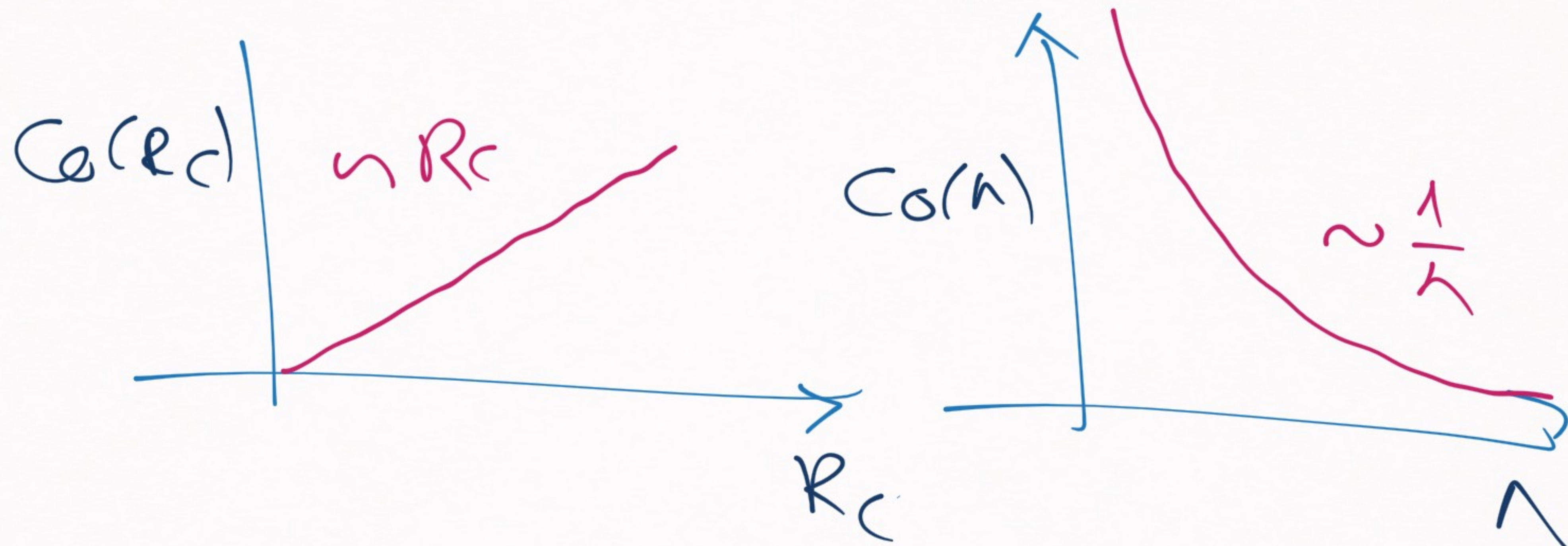


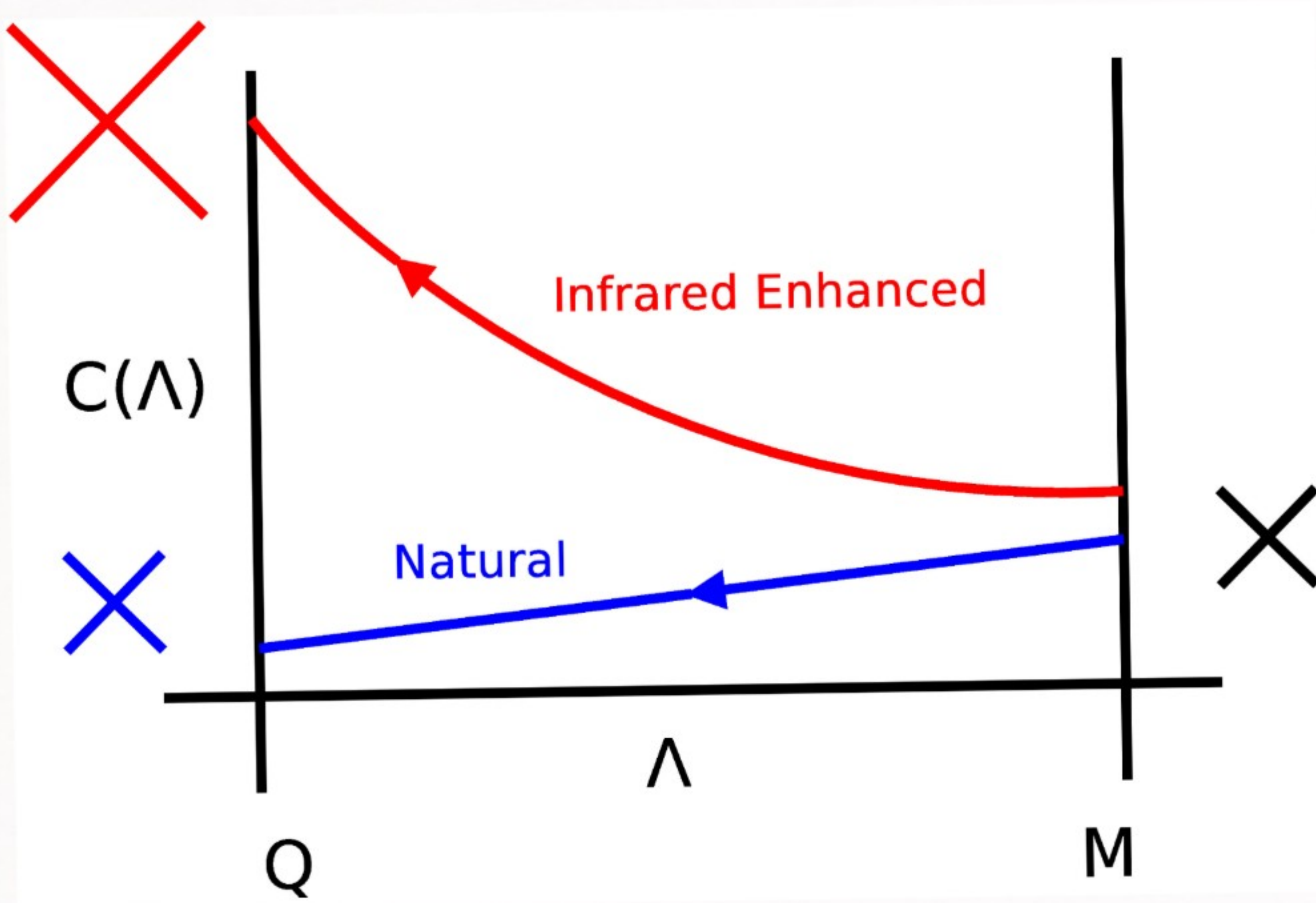
" $\mathbb{P}_0$ "  
 power  
 counting  
 $\downarrow$

$$\chi = \sum (d_0 + \frac{d_1}{X} + \dots)$$

$\Rightarrow$  TWO BODY SYSTEM IS THE SAME

IP  $\Rightarrow$  a bound state  $\Rightarrow$   $C_0(R_c)$





RGEs have several solutions



Depend on our assumptions

# Two-body system

1)  $\Rightarrow$  a shallow bound state  
( $\gamma \ll m$ )

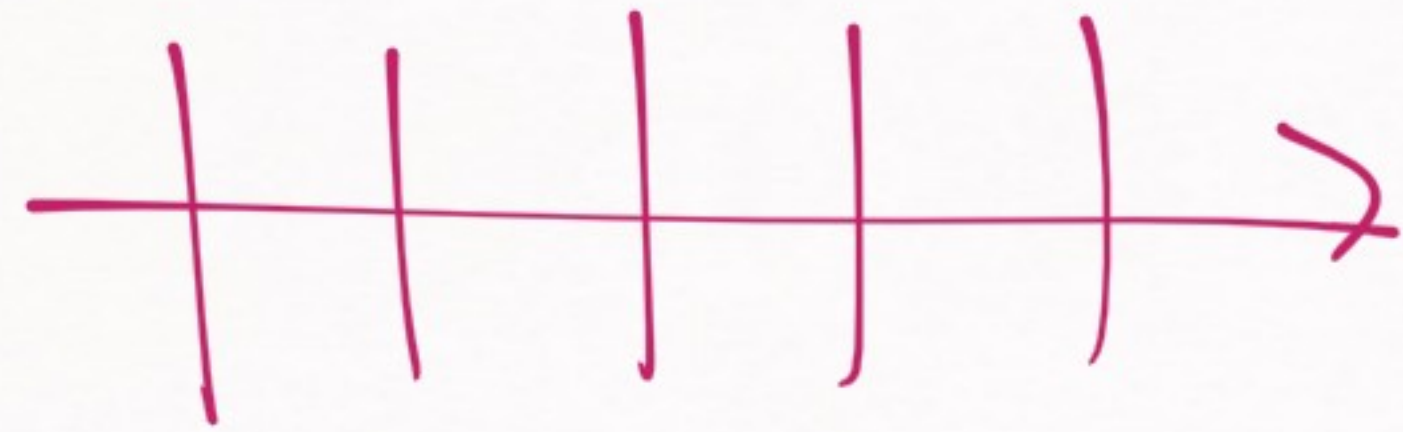
How to do it?  
 $\nearrow$

2)  ~~$\Rightarrow$~~  bound states, but we have  
instead a perturbative system

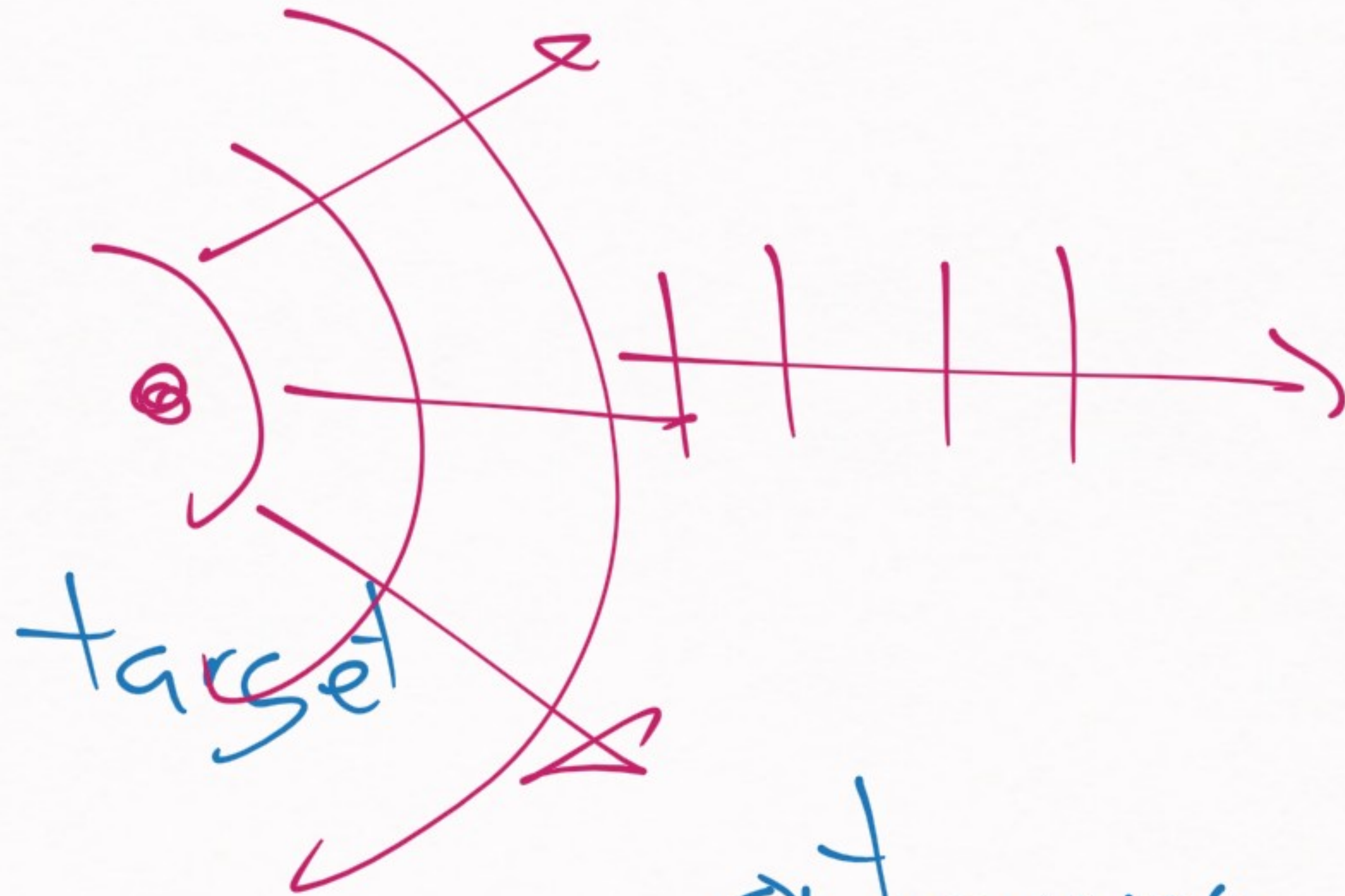


→ I will have to advance scattering theory (if you don't know scattering theory, ignore the following explanations until later)

## 2) SCATTERING BY A WEAK POTENTIAL



incoming particle



outgoing scattered particle

$$\psi(\vec{r}) \rightarrow e^{i\vec{k}\cdot\vec{r}} + f(\vec{k}) \frac{e^{ikr}}{r}$$

incoming wave                      scattered wave

$f(\vec{k}) \rightarrow$  scattering amplitude

$$\frac{d\sigma}{d\Omega} = |f(\vec{k})|^2$$

Scattering by a weak potential:

→ Born approximation

$$f(\hat{k}) \approx -\frac{m}{2\pi} \int d^3\vec{r} V(\vec{r}) e^{i\vec{k}\cdot\vec{r}}$$

(check it in your QM textbook)

→ Born approximation  $\Rightarrow \beta(\vec{k})$  small

$\psi(\vec{r})$  →  $\psi(\vec{r}) \approx e^{i\vec{k} \cdot \vec{r}}$  + small corrections

input for RGE

$$\langle \psi | V_c | \psi \rangle = \int d^3r \int d^3r' G(\vec{r}, \vec{r}') |e^{i\vec{k} \cdot \vec{r}}|^2$$

$$\textcircled{1} \quad \leq C_0(R_c) \quad \Leftrightarrow \quad \langle 4|V_c|4 \rangle \leq C_0(R_c)$$

$$\boxed{RGE} \quad \Leftrightarrow \quad \frac{d}{dR_c} \langle 4|V_c|4 \rangle = 0$$

$$\underbrace{\hspace{10em}}_{\neq}$$

$$C_0(R_c) \leq \text{constant}$$

$$\boxed{\frac{d}{dR_c} [C_0(R_c)] = 0}$$

Is this true? Can we check it? Yes!

$$\begin{aligned} \rho(\vec{k}) &= -\frac{\mu}{2\pi} \int d^3\vec{r} V(\vec{r}) e^{-i\vec{k}\cdot\vec{r}} \\ &= -\frac{\mu}{2\pi} V(\vec{k}) = -\frac{\mu}{2\pi} [V(\vec{0}) + (\text{expansion in } \vec{k}^2)] \end{aligned}$$

Low energies ( $|\vec{k}|^2 \ll m^2$ )  $\rho(\vec{k}) = -\frac{\mu}{2\pi} V(\vec{0})$

$$V_{\text{eff}}(\vec{q}) = C_0 \rightsquigarrow \rho_{\text{eff}}(\vec{k}) = -\frac{M}{2\pi} C_0$$

$$V_Y(\vec{q}) = -\frac{g_Y^2}{g^2 + m^2} \rightsquigarrow \rho_Y(\vec{k}) = -\frac{M}{2\pi} \left( -\frac{g_Y^2}{m^2} \right) + \text{corrections}$$

$$\rho_Y(\vec{k}) \rightarrow +\frac{M}{2\pi} \left( \frac{g_Y^2}{m^2} \right)$$

$$\rho_C(\vec{k}) \rightarrow -\frac{M}{2\pi} C_0$$

$$C_0 = -\frac{g_Y^2}{m^2}$$



$$\frac{d}{dR_c} [\delta(R_c)] = 0 \rightarrow \text{Weak potential}$$

$$\frac{d}{dR_c} \left[ \frac{\delta(R_c)}{R_c} \right] = 0 \rightarrow \text{Strong potential} \\ \text{w/ shallow} \\ \text{bound state}$$

Message



DEPENDING ON THE SITUATION  
WE WILL HAVE DIFFERENT  
BEHAVIORS FOR Co(Rc)



~ Different RGE's or EAT's



(FINISH FOR  
TODAY)