

NUCLEAR PHYSICS | ②

Why nuclear physics is difficult?

RECAP

→ A general principle:

[Scales are important to
understand physics]

✓ check → typical size of a system
→ typical momentum
of a system

→ Natural problems

→ natural scale is easy to find

→ easy to estimate the size

of an observable on

purely dimensional ground

(numerical factors $\sim \mathcal{O}(1)$)

Example \rightarrow [HYDROGEN ATOM]

$$\left[-\nabla^2 - \frac{2}{a_B r} \right] \Psi(\vec{r}) = -\gamma^2 \Psi(\vec{r})$$

$$\rightarrow \Psi(\vec{r}) = \frac{1}{\sqrt{4\pi}} \frac{2}{a_B^{3/2}} e^{-r/a_B} \quad (\text{solution})$$

SCALES



a_B (length scale)

$\alpha_B = \frac{1}{a_B}$ (momentum scale)

$$\alpha_B = \frac{h c}{m_e \alpha} \approx 5.3 \cdot 10^{-4} \text{ m} = 0.53 \text{ \AA}$$

$$\Phi_B = m_e \alpha \approx \frac{0.511 \text{ MeV}}{137} \approx 3.7 \text{ KeV}$$



We can define natural scales

$$\Phi_B = \frac{1}{\alpha_B}$$

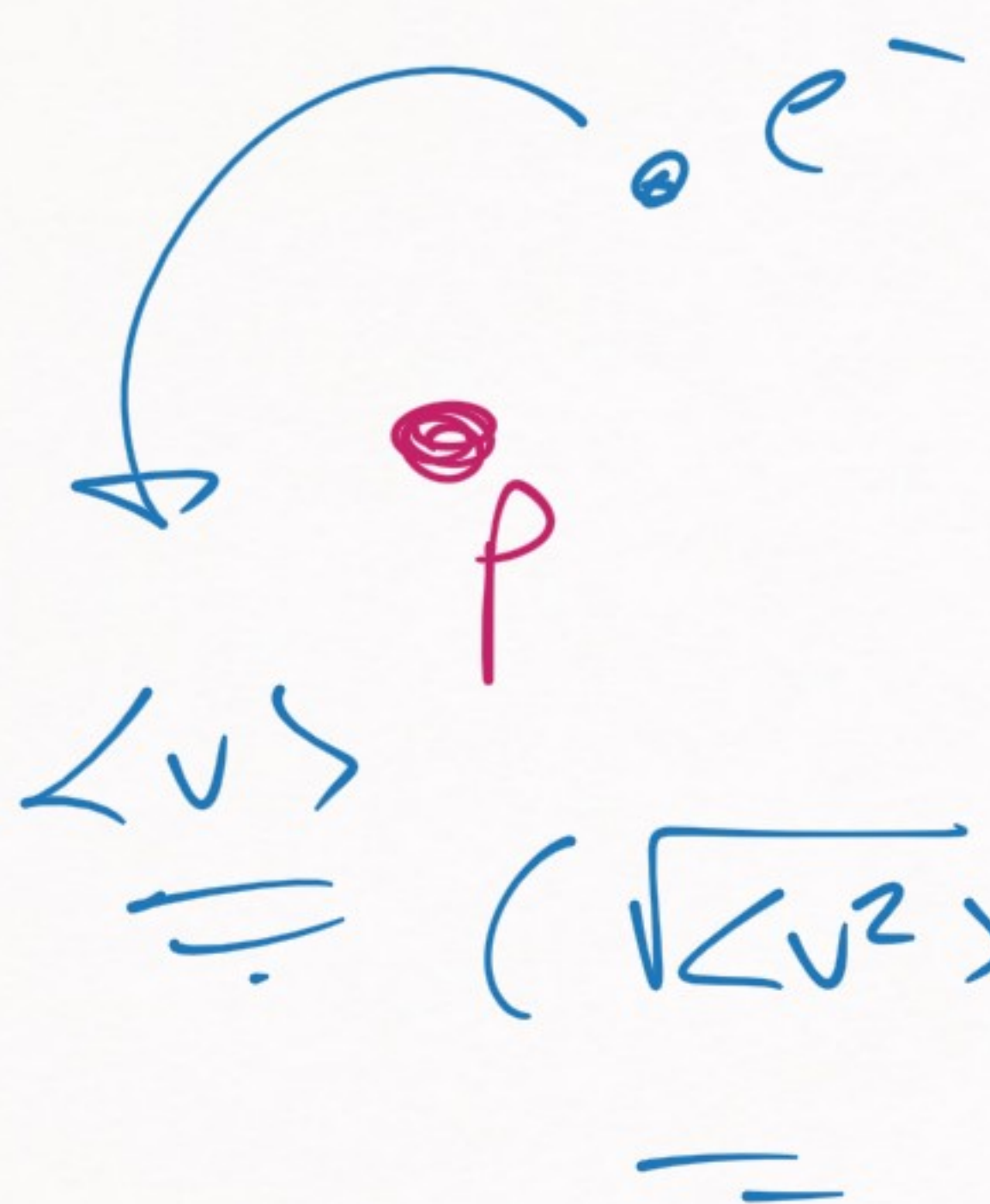
How do we use this knowledge
about scales?

$$E_B = -\frac{\gamma^2}{2\mu} \leadsto \sqrt{\gamma} \sim \left[\frac{1}{a_B} \right] \times O(1)$$

$$\sqrt{\langle r^2 \rangle} \sim O(1) \times a_B$$

$$\frac{\sqrt{\langle p^2 \rangle}}{m_e} \sim \langle v \rangle \sim O(1) \times \frac{Q_B}{m_e} \rightarrow \text{Let's see this one}$$

→ Let's see this example of the speed of an electron within the H atom



$$\sqrt{\frac{\langle p^2 \rangle}{m_e^2}} = \sqrt{\langle v^2 \rangle} = \mathcal{O}(1) \times \frac{\hbar}{m_e} = \mathcal{O}(\hbar)$$

$$p = mv \quad \mathcal{O}(\hbar) = \mathcal{O}(1) \times \frac{m_e \alpha}{m_e} = \mathcal{O}(\alpha)$$

$$\mathcal{O}(\hbar) = \mathcal{O}(1) \times \alpha, \quad \alpha \approx \frac{1}{137}$$

$$\sqrt{\langle v^2 \rangle} \sim \frac{1}{137} c \rightarrow \text{what does this mean?}$$

=

$$\rightarrow \text{natural units: } \left(\frac{v}{c} \right) \quad (c=1)$$

$$\sqrt{\langle v^2 \rangle} \sim \frac{c}{137} \sim \frac{3 \cdot 10^5 \text{ km/s}}{137} \approx \underline{2200 \text{ km/s}}$$

Expectation (from dimensional argument
/ from naturalness): $\sqrt{2VZ} \sim \alpha$

Explicit calculation: $\hat{p}^2 = -\vec{\nabla}^2$
 $\psi(\vec{r}) = \frac{1}{\sqrt{4\pi}} \frac{2}{a_B^{3/2}} e^{-r/a_B}$

$$\sqrt{\frac{\langle 4 | \hat{p}^2 | 4 \rangle}{m_e^2}} = \alpha \rightarrow \boxed{\text{spot on}}$$

[Take-home message]

dimensional
estimate

$$\gamma \sim 1/a_B$$

$$\sqrt{\langle r^2 \rangle} \sim a_B$$

$$\sqrt{\langle v^2 \rangle} \sim \frac{1}{\alpha}$$

actual
calculations

$$\gamma = 1/a_B$$

$$\sqrt{\langle r^2 \rangle} = \sqrt{3} a_B$$

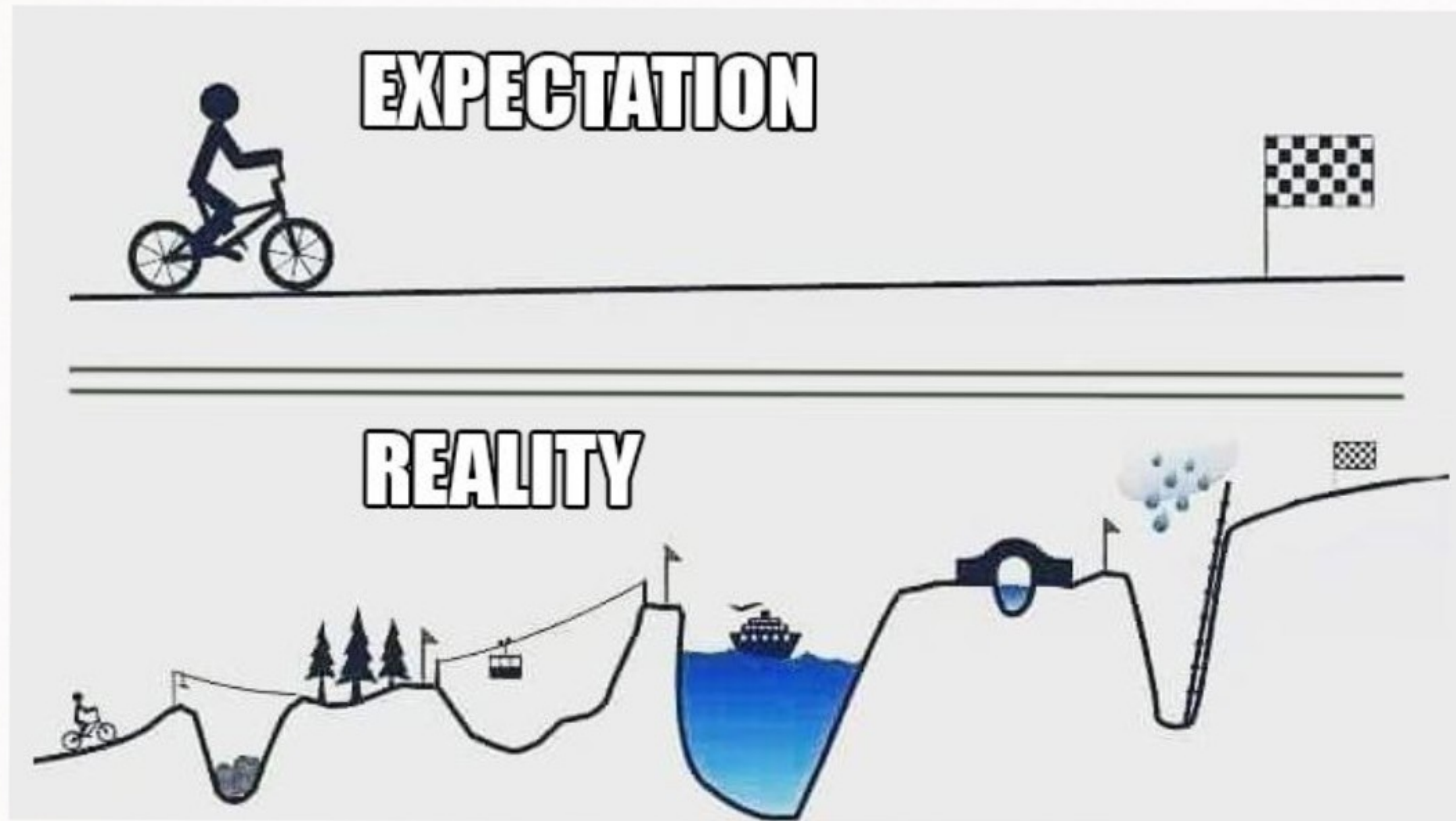
$$\sqrt{\langle v^2 \rangle} = 1/\alpha$$

→ this type of arguments & claims
"naturalness" do work in many cases

(not always the case)

(but very often it is the case)

But life is in general not always so easy



→ naturalness assumption

→ real world

[WHEN WILL NATURALNESS WORK?]

NATURAL PROBLEM

(i) CHARACTERISTIC SCALE (SIZE/MOMENTUM)

H atom \rightarrow (electron mass)

\times (strength of Coulomb)

$\alpha_B = m_e \alpha \ll 3.7 \text{ keV}$ $\left(\frac{\Delta}{\alpha_B} = a_B \ll 0.53 \text{ \AA} \right)$

→ It is easy to find a natural scale

(ii) EVERY OBSERVABLE IS $O(1)$
IN TERMS OF THE CHARACTERISTIC
SCALE

$$[\langle \hat{O} \rangle] = [L]^n \rightsquigarrow \langle \tilde{O} \rangle \sim \underline{\underline{O(1) a_B^n}}$$

NATURAL PROBLEM

→ characteristic scale

→ \forall natural ($\omega(1)$) in this scale

MESSAGE →

NATURAL PROBLEMS
ARE EASY TO
DEAL WITH!

Another example of a natural problem:

BOILING POINT OF CESIUM



liquid/gas
transition

→ Use scale argument

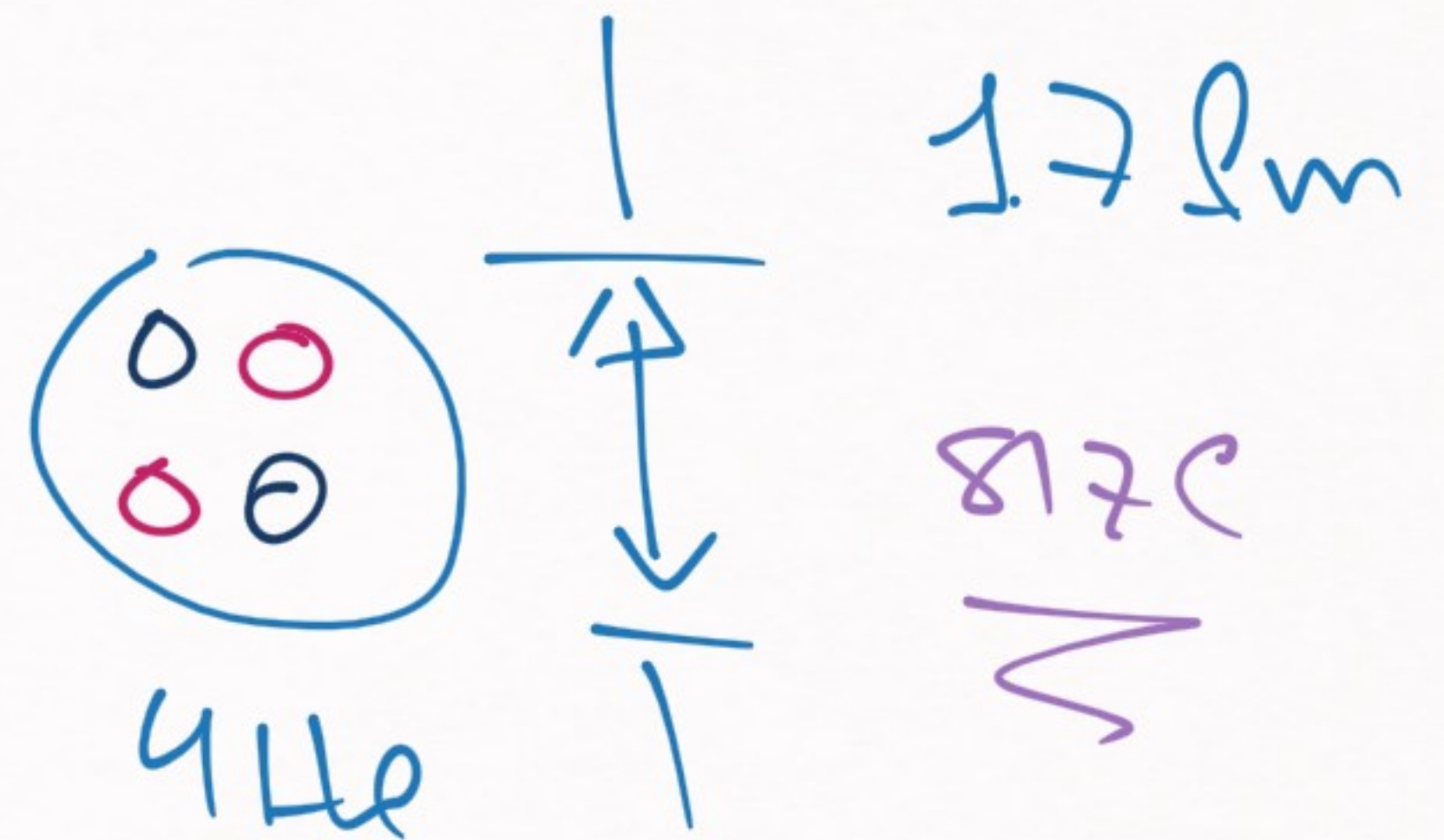
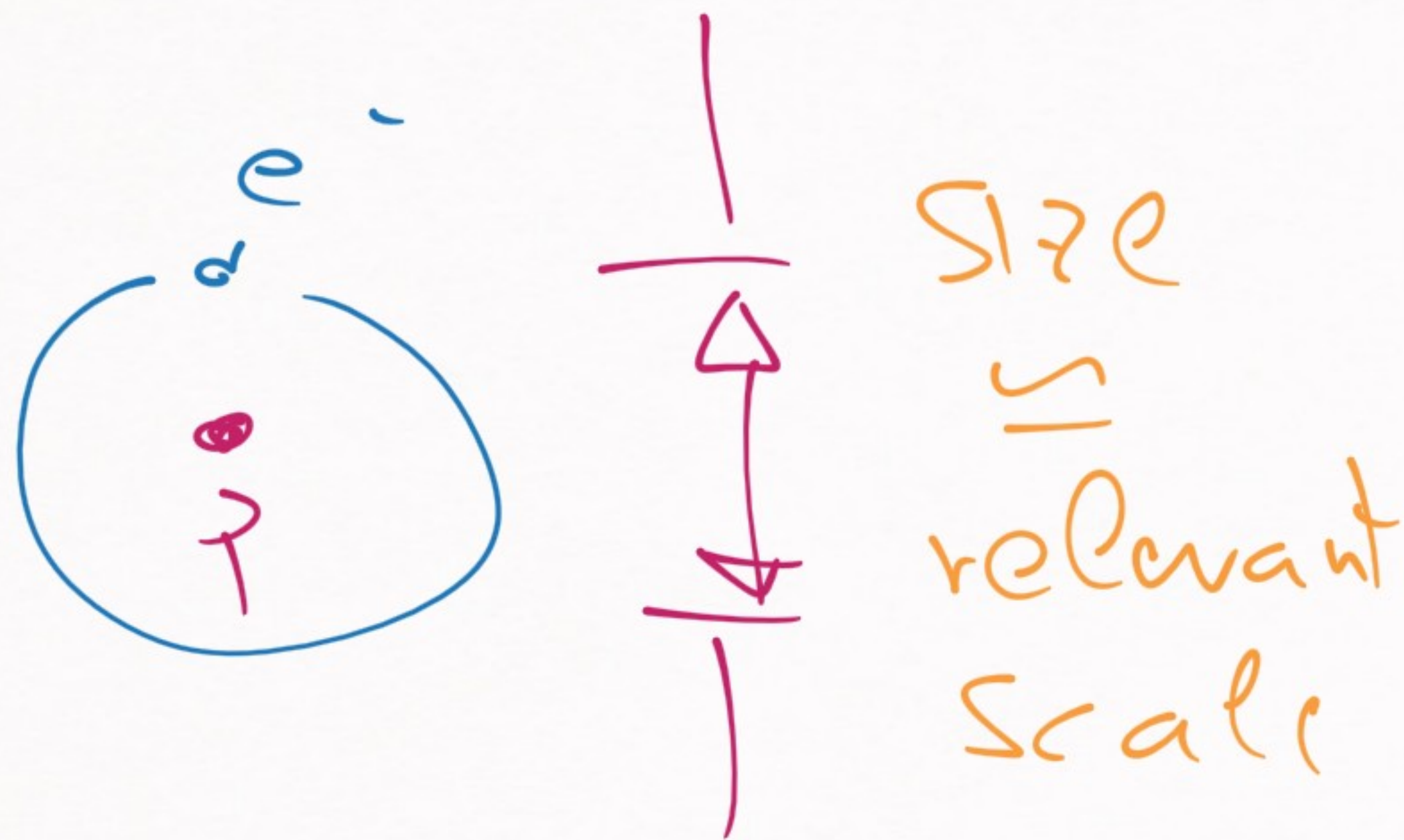


the element

铯

→ a part of this calculation is not trivial

↳ identification of the relevant scales

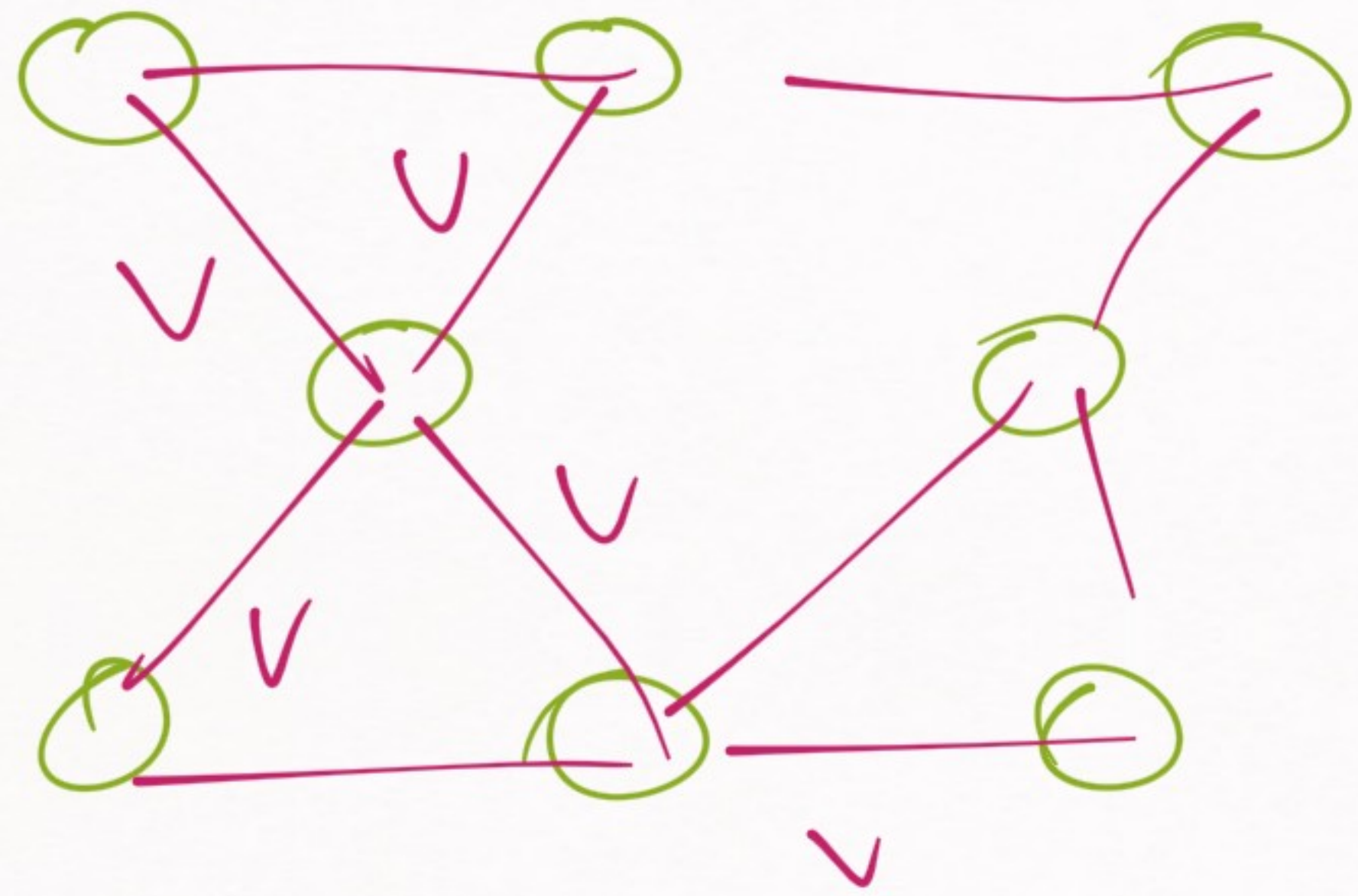


SCHEMATICS OF THE ARGUMENT I WILL USE

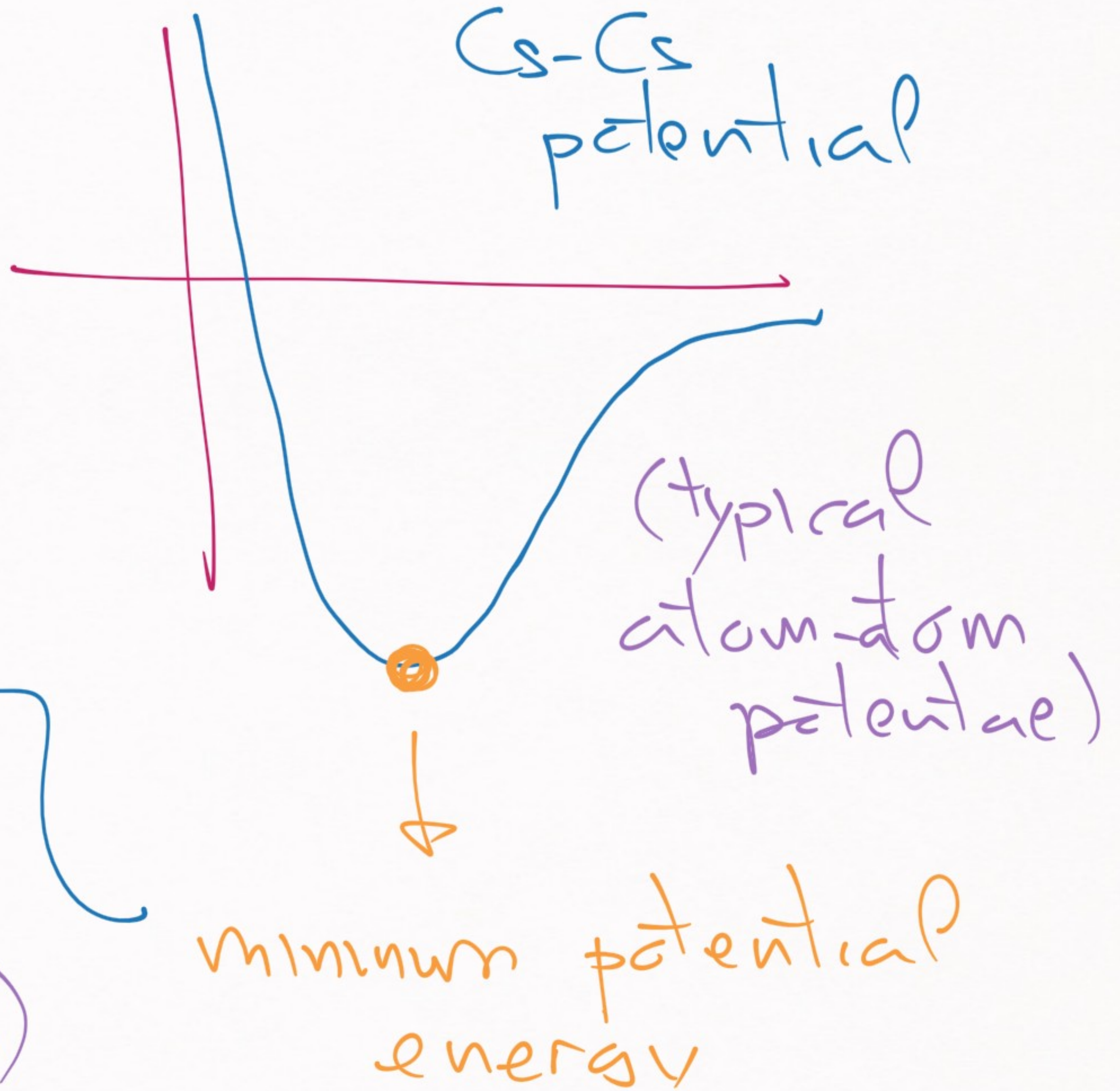
→ A group of Cs atoms, they will have a natural average speed $\sqrt{\langle v^2 \rangle}$

→ If the thermal speed of atoms is similar to this natural

speed \Rightarrow Cs needs to be a gas



average separations
 &
 average speed
 (momentum)



→ Estimate the expected momentum of Cs atoms (when ignoring external factor such as temperature)

↙
Objective #1

→

We need info
about Cs

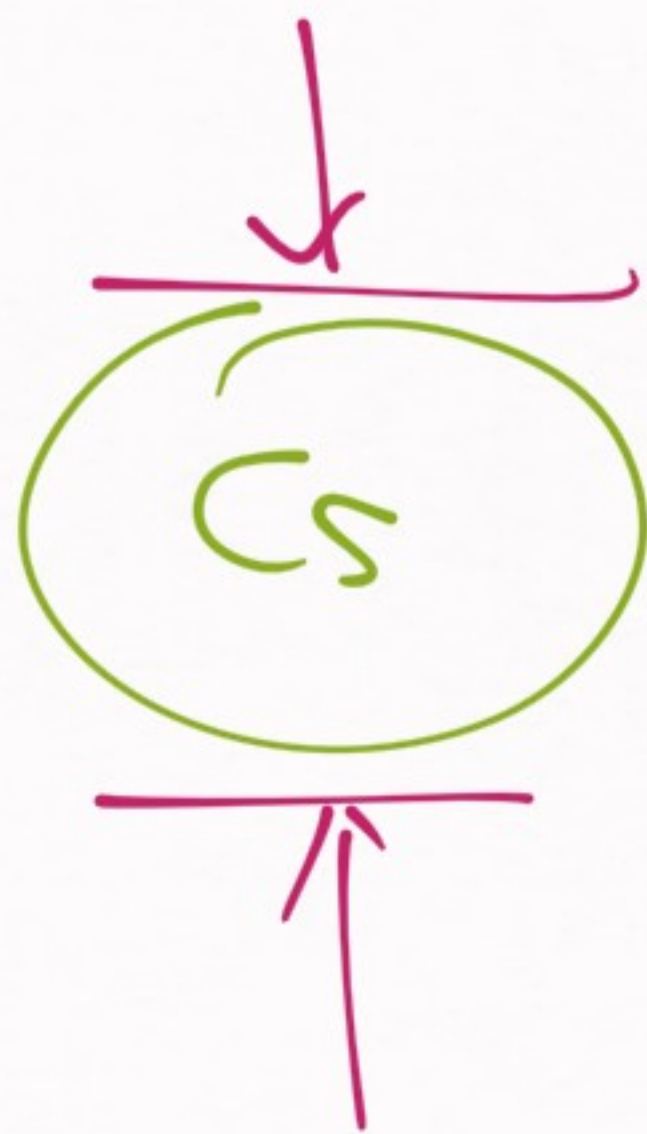
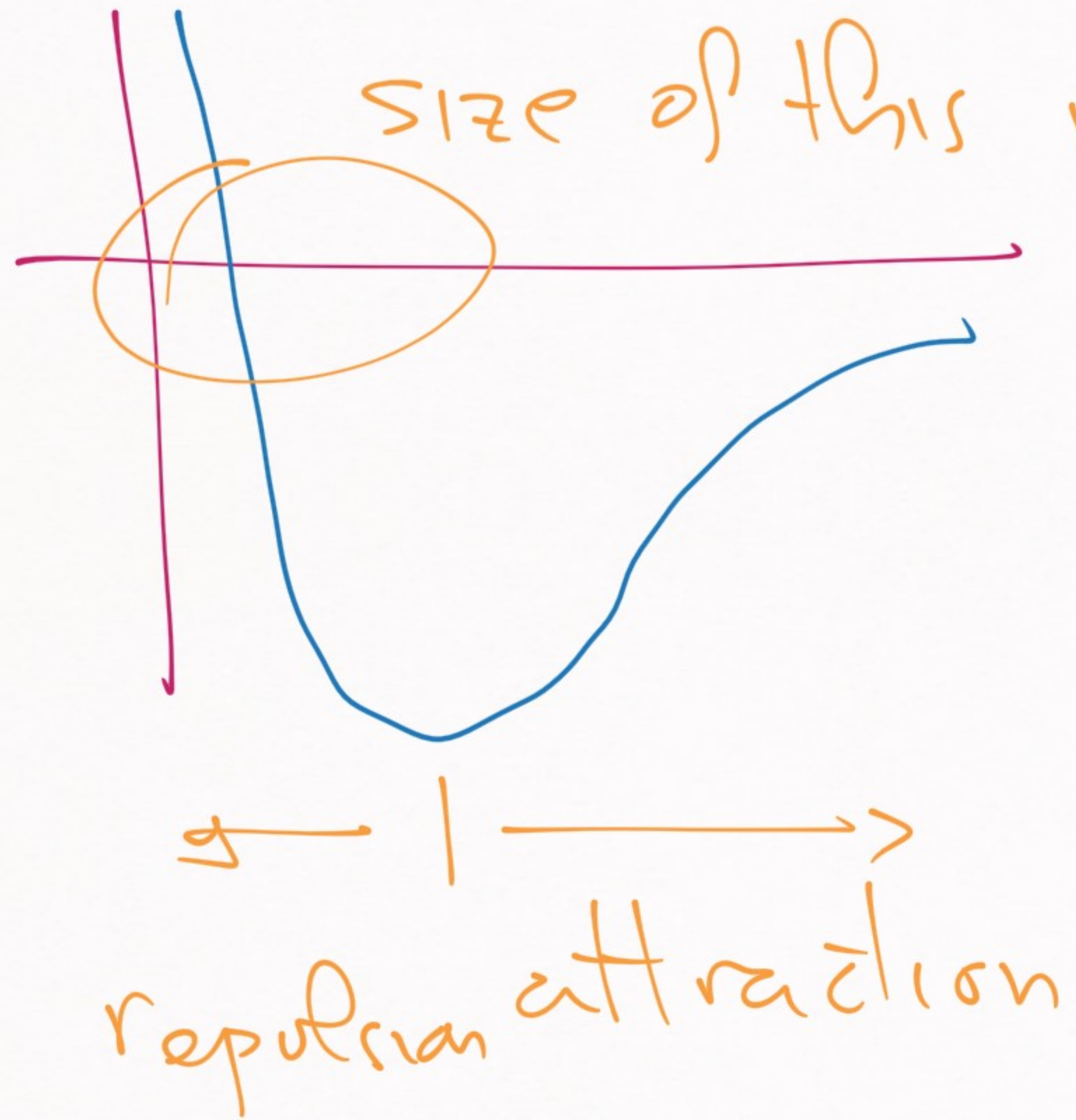
C_s

$^{133}C_s \rightarrow$ atomic number: 133

$m(C_s) \approx 130 \text{ GeV}$ (approximately)

why? \rightarrow 133 nucleons

$m_p \approx m_n \approx 1 \text{ GeV}$
(0.94 GeV)

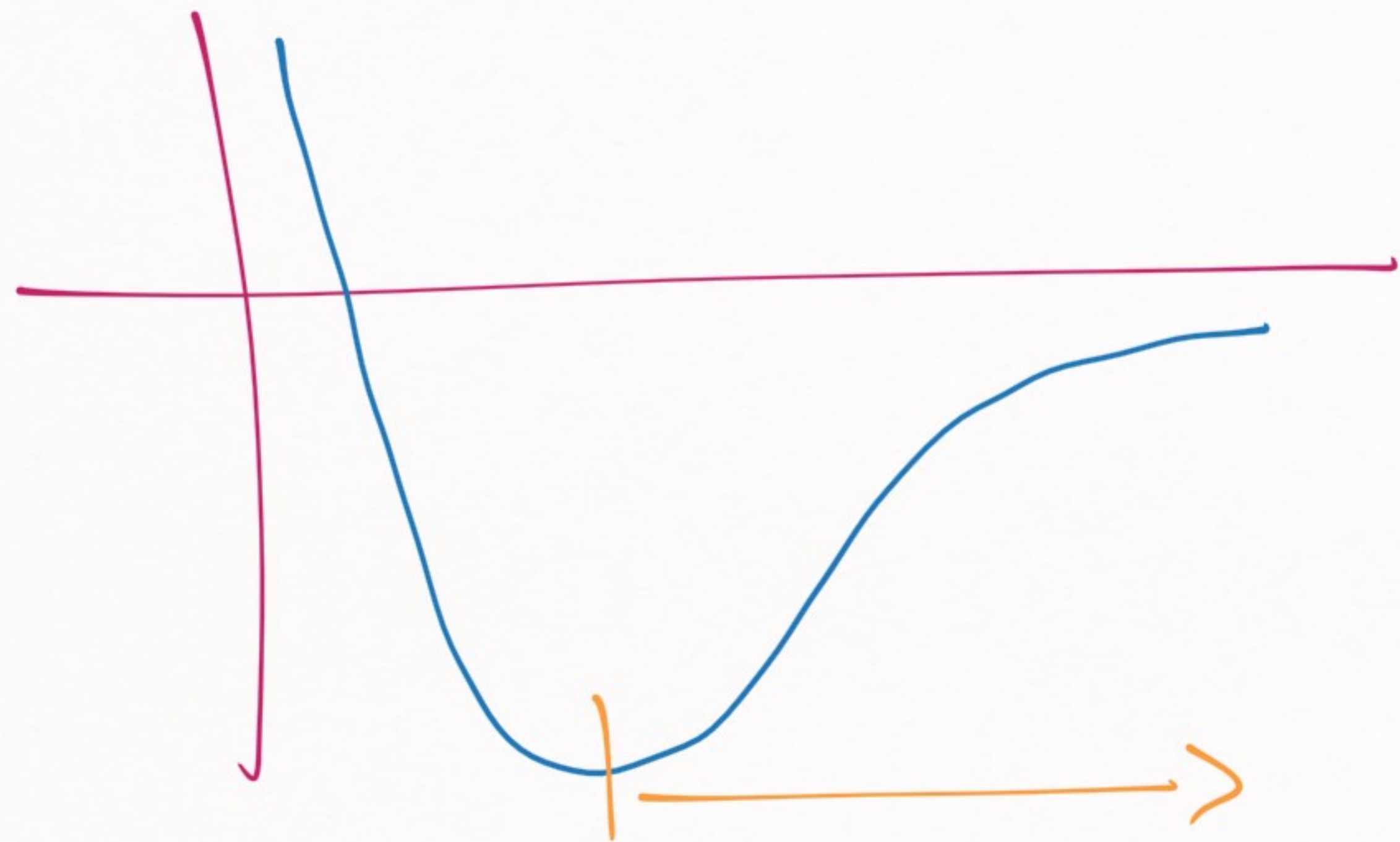


size of Cs

≈ 7 a.u.

≈ 7 a.u. = ≈ 7 a_B

≈ 7 a_B



attraction

typical atomic potential

$$V_{AA}(r) = \frac{C_{12}}{r^{12}} - \frac{C_6}{r^6}$$

(Lennard-Jones potential)

short-range repulsion

long-range attraction

Long-range attraction \rightarrow well known

\rightarrow given by the van der Waals potential

$$V_{\text{vdW}}(r) = -\frac{C_6}{r^6} = -\frac{1}{2\mu} \frac{\overset{4}{R_6}}{r^6} \rightarrow \text{length scale}$$

Why this step? \rightarrow I want a scale

\downarrow

$$\frac{C_G}{r_G} = -\frac{1}{2\mu} \frac{R_G^4}{r_G}$$

$$\rightarrow R_G = \left(2\mu (C_G)^{1/4} \right)$$

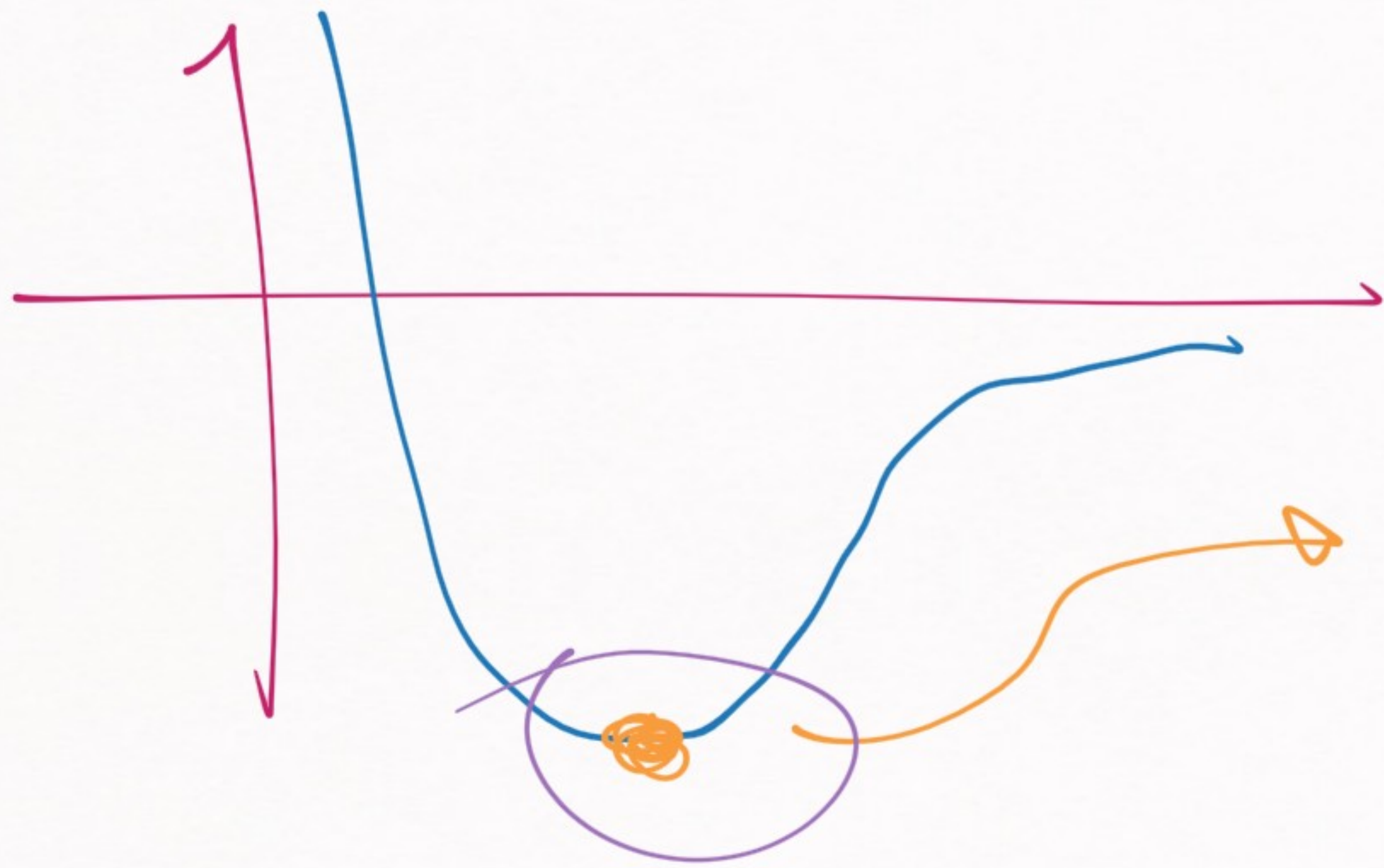
C_G, μ are well-known \rightarrow $R_G \approx 200 \text{ a.u.}$

Two possible relevant scales:

$$R(CS) \equiv R_5 = 7 \text{ aB}$$

$$R(vdW) \equiv R_6 = 203 \text{ aB}$$

→ they are still not the scale
we are looking for



Cs atoms will try to get into minimum energy configurations

$E_{kinetic} \approx E_{potential}$

$$\rightarrow E_{kin} = \frac{1}{2} m \langle v^2 \rangle$$

$\langle v^2 \rangle$

\downarrow

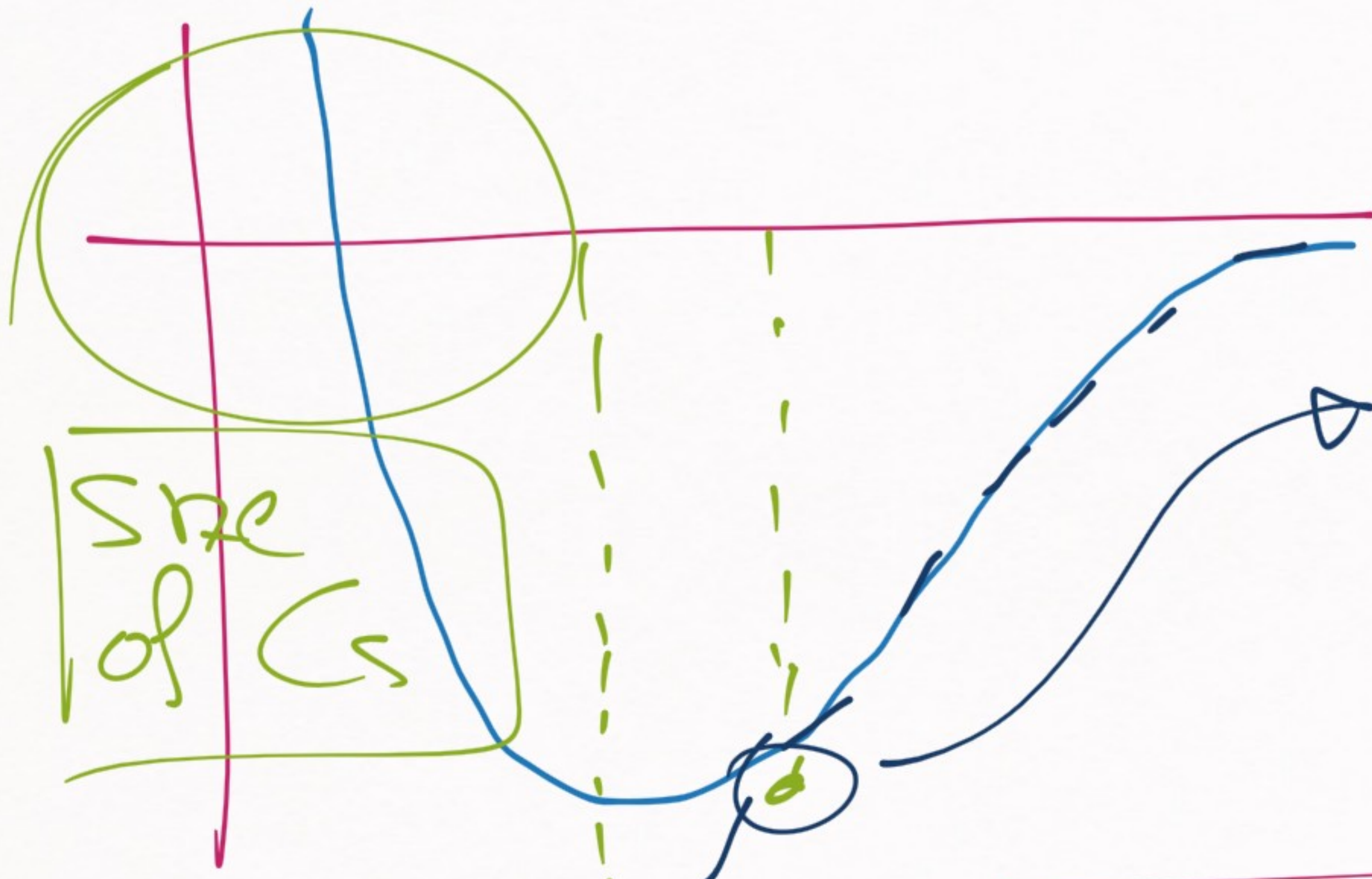
AIM \rightarrow get $\langle v^2 \rangle$ for a group of Cs atoms (or $\langle E_{kin} \rangle$)

(Limitation \rightarrow only dimensional arguments)

\leftarrow Looking at previous page

$$\langle E_{kin} \rangle = \langle v_{min}^2 \rangle \rightarrow$$

we want to estimate this



$$V_{min} \approx V_{vdW}(P \approx P_S)$$

$$V_{vdW}(P_{CS}) < V_{min}$$

van der Waals \rightarrow we know

$R_S \lesssim 7a_B$ $\Rightarrow D$ (50% more maybe)
(approximations)

$\Delta =$
 $10a_B$

$$V_{\min} \lesssim V_{vdW} (R = 10a_B)$$

$$V_{\min} = -\frac{1}{2\mu} \frac{R_0^4}{R_{\min}^6}$$

$$R_{\min} \leq 10 \text{ a.u.}$$



↳ From here I will get the expected size of the speed/energy of Cs atom

$$|\langle E_{\min} \rangle| \approx |V_{\min}| \Rightarrow \underbrace{\frac{1}{2\mu} Q_{CS}^2}_{\text{kinetic energy}} = \frac{1}{2\mu} \frac{R_6^4}{R_{\min}^6}$$

→ doing the calculation

$$Q_{CS} \approx 20 Q_B \approx \frac{20}{a_B}$$

↓
3.7 keV



$$\langle R \rangle \sim R_{min}$$

$$\langle E_{kin} \rangle \sim \frac{1}{2\mu} Q_{cs}^2$$

$$\langle v \rangle \sim \frac{Q_{cs}}{m_{cs}}$$

$$Q_{cs} \sim 70 \text{ keV}$$



$$\langle v \rangle \sim 350 \text{ m/s}$$

Compare this $\langle v \rangle_{cs}$ w/ $\langle v \rangle_{\text{thermal}}$

\Rightarrow when $\langle v \rangle_{\text{thermal}} \geq \langle v \rangle_{cs}$

\Rightarrow only gas state is possible

\Rightarrow Boiling point

$$E_T = k_B T$$

$k_B \rightarrow$ Boltzmann constant

$$k_B = \frac{1 \text{ eV}}{11000 \text{ K}}$$

Room temperature
 $T \sim 300 \text{ K}$

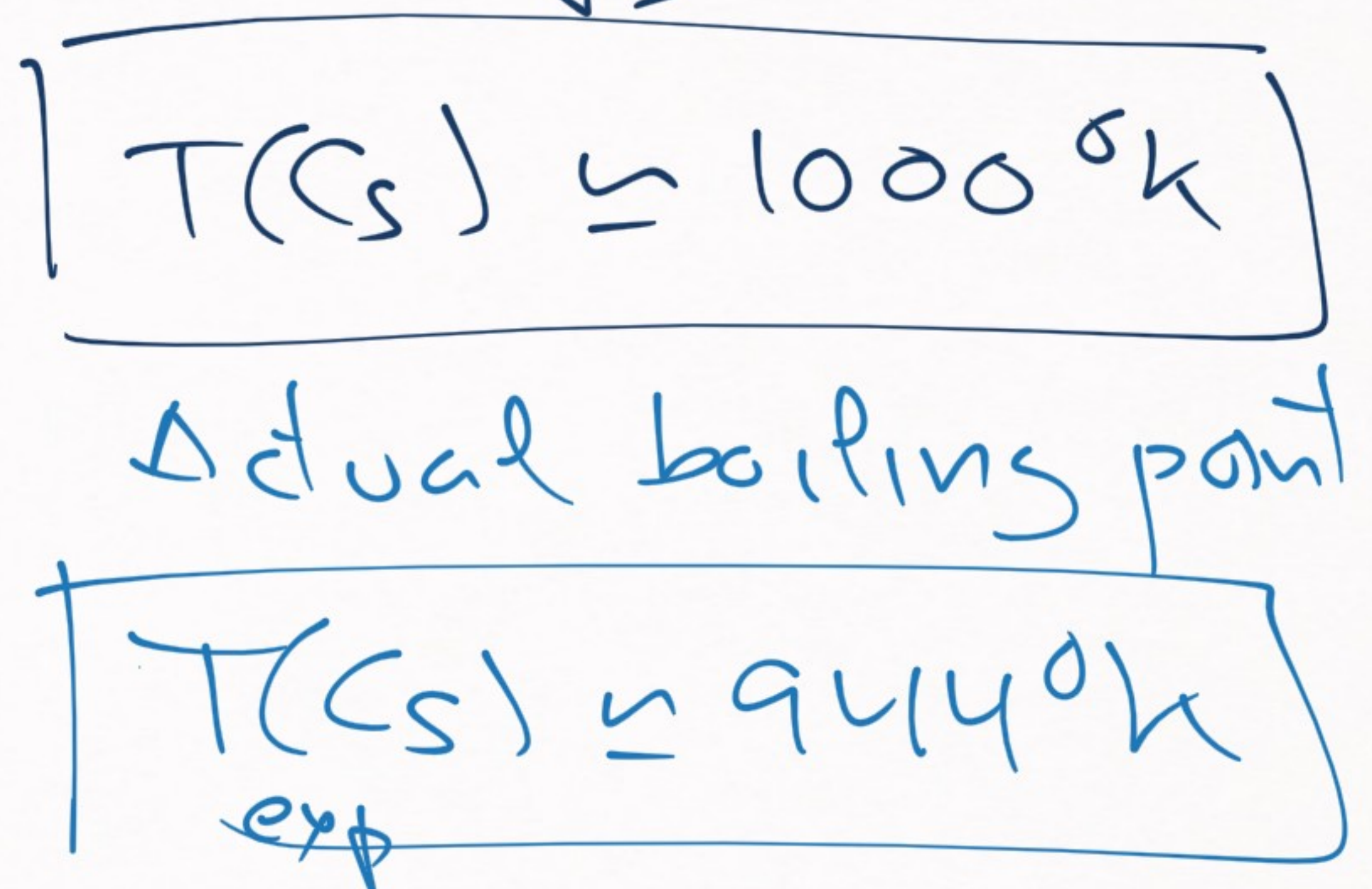
$$E_T \sim 50 \text{ meV}$$

$$\frac{1}{2} m \langle v \rangle^2 = k_B T \quad \rightarrow \quad T = \frac{\frac{1}{2} m \langle v \rangle^2}{k_B}$$

(kinetic = thermal)

$$\langle v \rangle \approx \frac{Q}{mcs} \approx 350 \text{ m/s}$$

$$\langle v_T \rangle \approx 350 \text{ m/s}$$



→ Cs boiling point (natural problem)

Dimensional argument → $T(\text{Cs}) \approx 10000^\circ\text{K}$

Experiment → $T(\text{Cs}) \approx 944^\circ\text{K}$

→ WORKS REALLY WELL

→ FOR NATURAL PROBLEMS YOU CAN
USE REALLY DUMB APPROX.

(AND THEY WILL WORK)

$$T_{\text{dim}}(Cs) \leq 1000^\circ K \quad T_{\text{real}}(Cs) \leq 944^\circ K$$

NATURAL PROBLEMS \rightarrow EASY

(usually \exists one relevant scale)

\rightarrow most problems in physics involve several scale (multiscale problems)

\rightarrow corrections

[How to DEAL w/ THESE CORRECTIONS]

→ [+1 ATOM]

1) most important scale $a_B = \frac{1}{m_e \alpha} \left(\frac{1}{9k} = Q_E \right)$

2) but \Rightarrow other scales

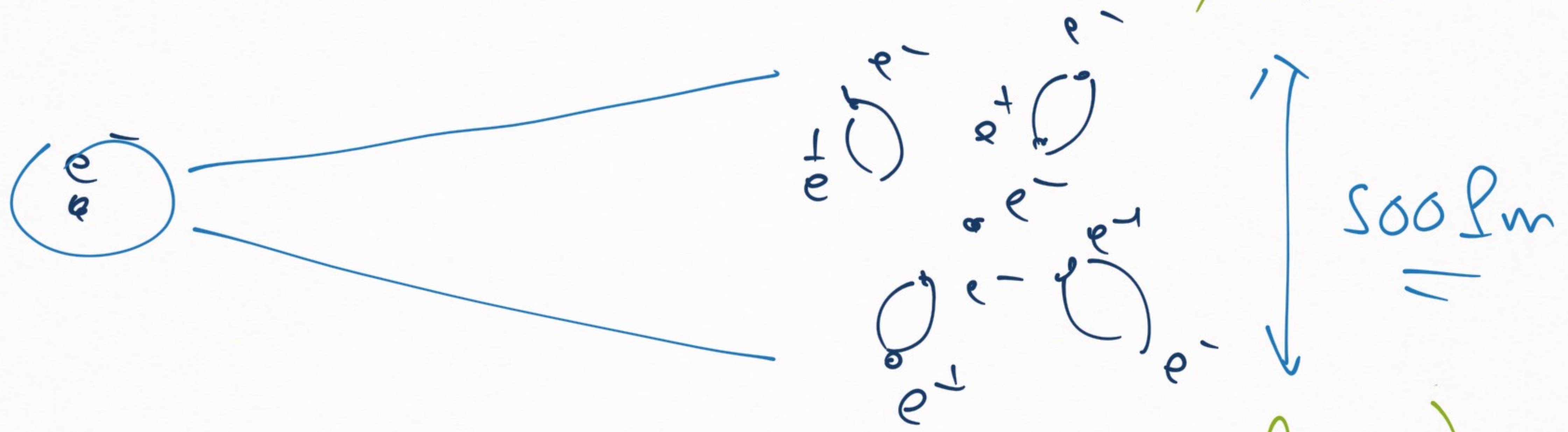
r_e, r_p → electron & proton size
?

Other scales →

ARE THEY BIG OR SMALL?

→ electron size (isn't the electron point-like?)

yes, but...



... vacuum polarization.

→ What is vacuum polarization?

→ creation of electron-positron pairs
that change the electromagnetic
properties of vacuum

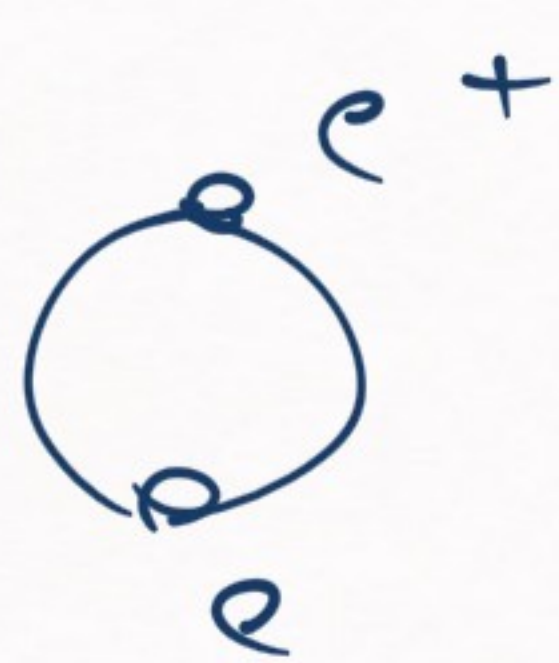
→ What is the size of this effect?



$$\Delta E \Delta t \leq \frac{\hbar}{2}$$

(uncertainty
relations)

$$\Delta x \Delta p \sim \hbar$$



$$\Delta x \sim \frac{\hbar c}{\Delta p}$$

$m_e c$



effect from
AFT

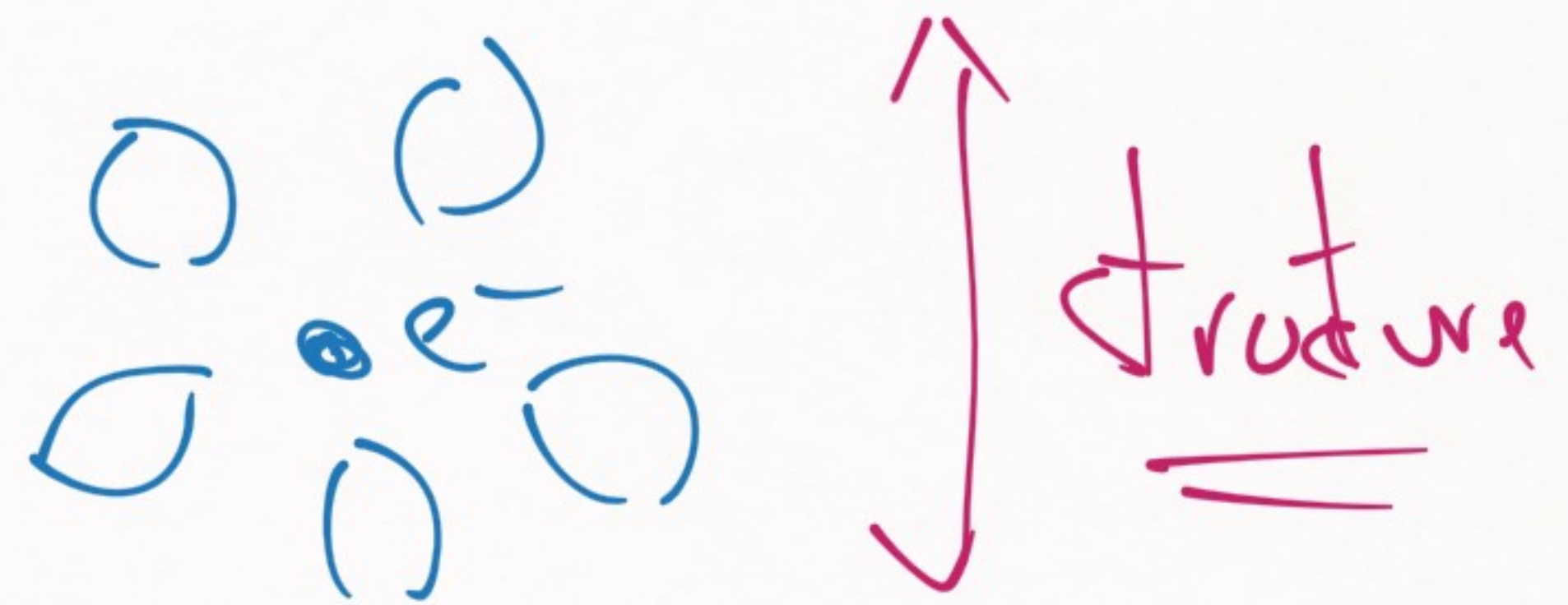
"size" of
electron
~

$$\rightarrow \left[\frac{1}{m_e} \sim \frac{\hbar c}{m_e} \sim 400 \text{ fm} \right]$$

$$a_B \approx 53000 \text{ fm}$$

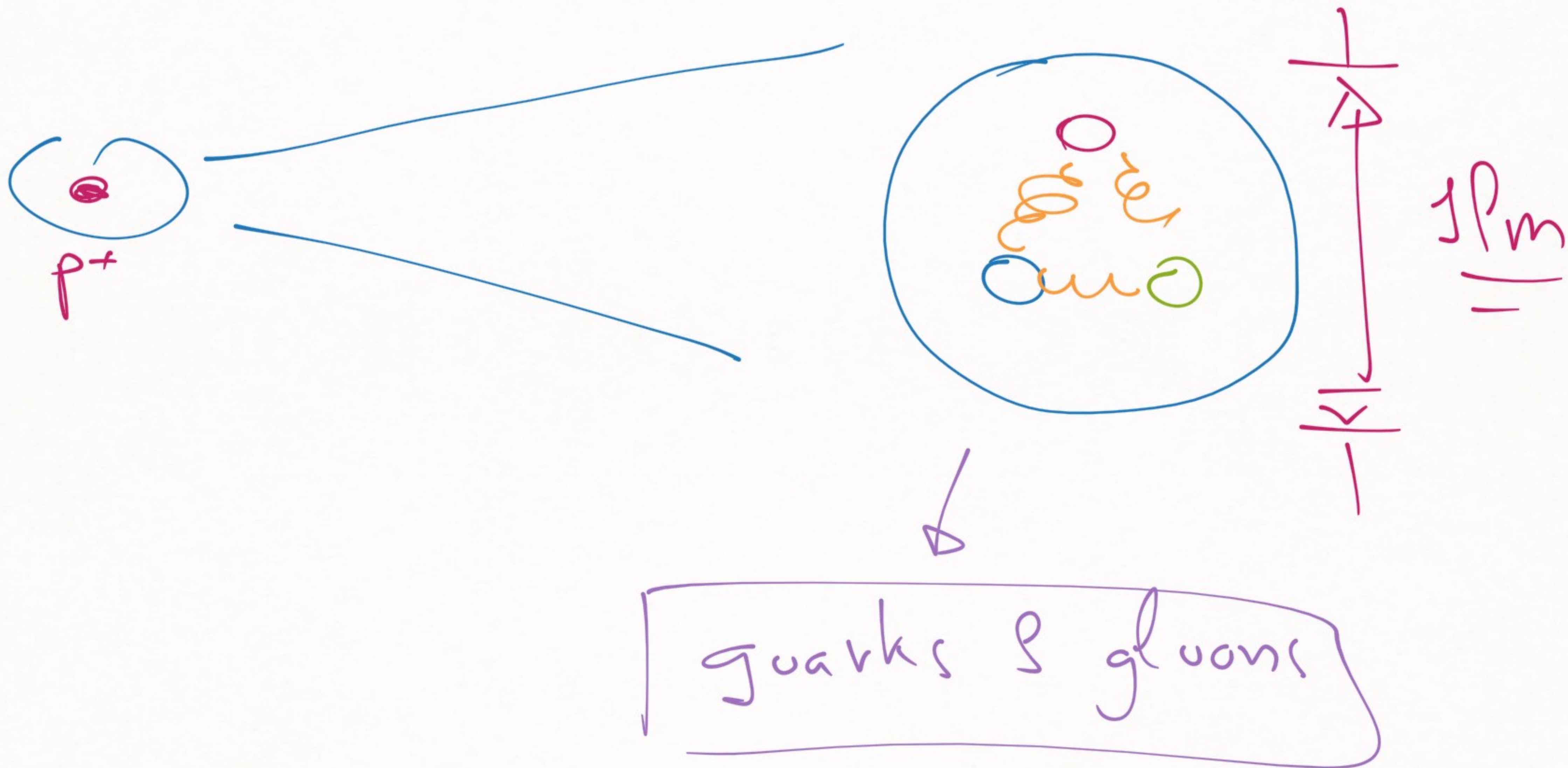
$$\frac{\lambda}{m_e} \approx 400 \text{ fm}$$

e^-
(point-like)

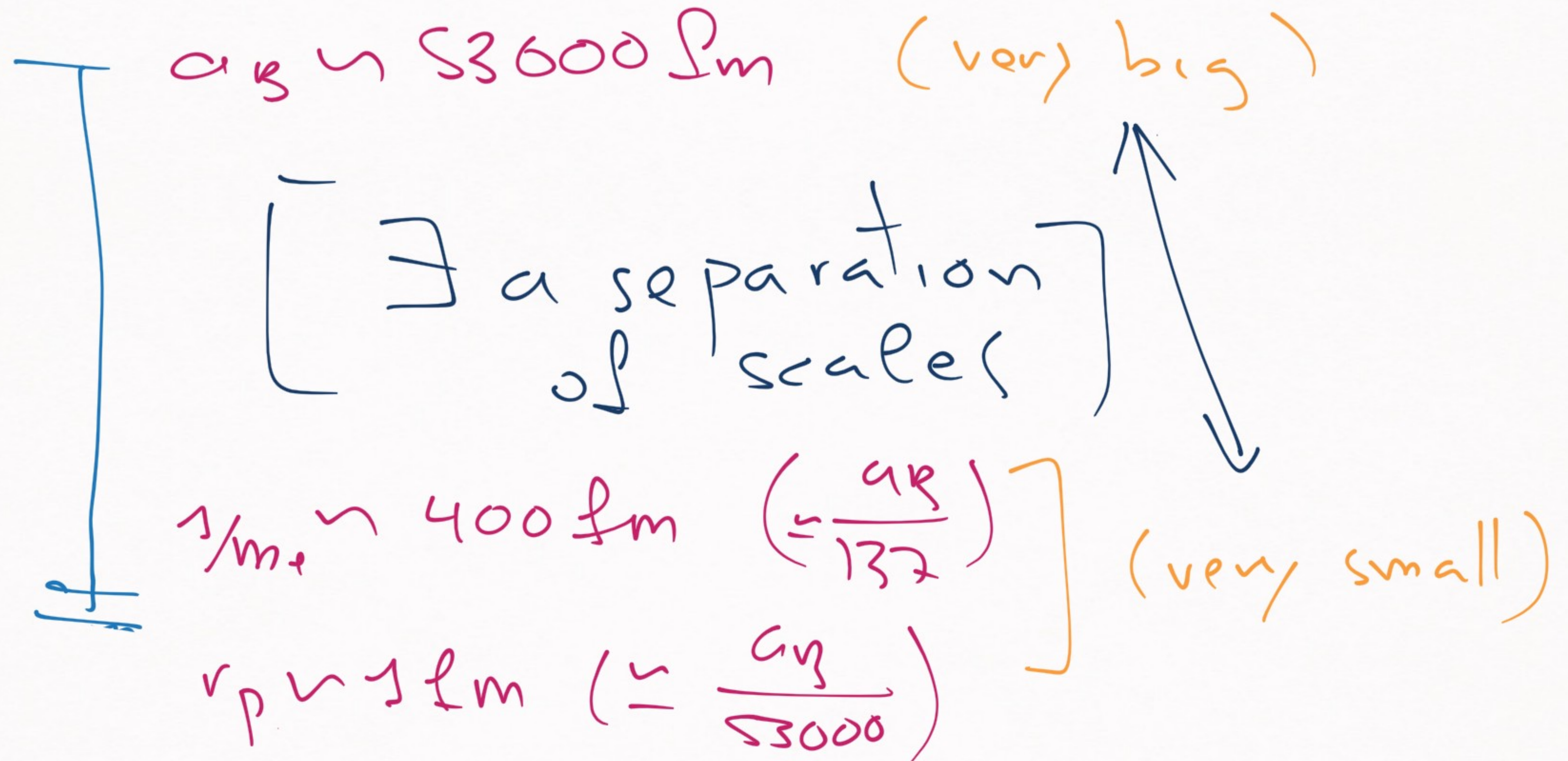


long distances
→ no structure

short distances
→ structure



Hydrogen Atom \rightarrow MULTISCALE PROBLEM



Bigger scales, smaller scales

→ good scale separation

1) Most important properties → bigger scale

2) Smaller scales → corrections

2) Smaller scales:

$$2.a) r_e \sim \frac{1}{m_e} \Rightarrow \frac{r_e}{a_B} \approx \frac{1}{137}$$

(even smaller because dynamics
of orbit r_e/a_B corrections)

$$2.b) \frac{r_p}{a_B} \sim 2 \cdot 10^{-5}$$

EXTRA SCALES \rightarrow \exists CORRECTIONS

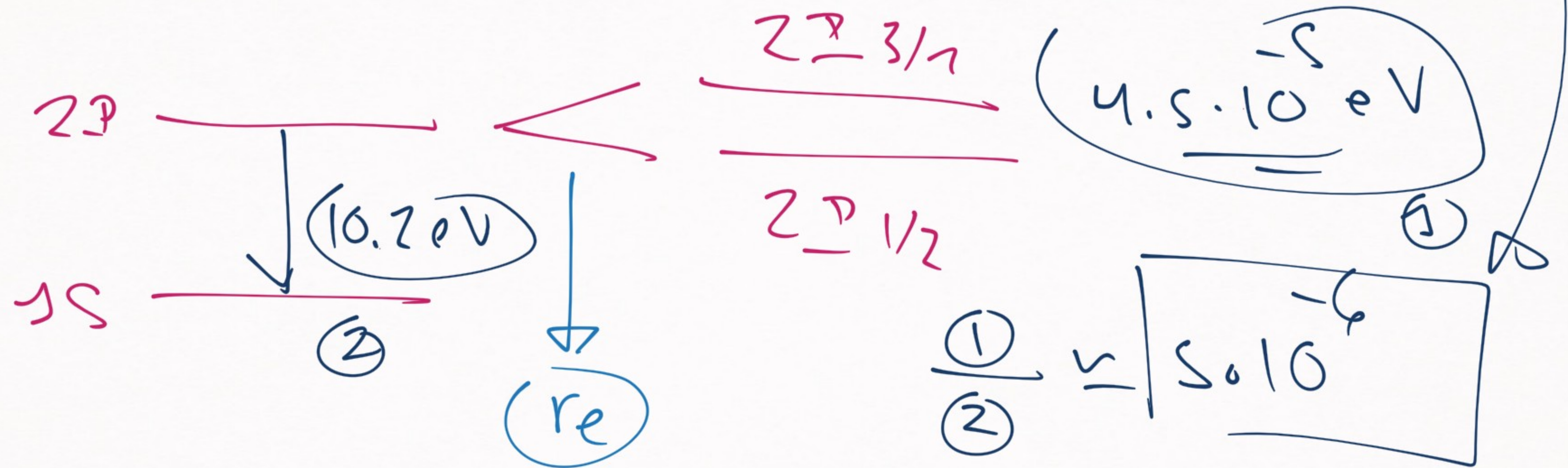
(CORRECTION 1) electron size
relativistic in nature

\swarrow

$$\left(\frac{r_e}{a_B}\right)^2 \sim \left(\frac{1}{137}\right)^2 \sim \underline{\underline{5 \cdot 10^{-5}}}$$

EXPECTATION \rightarrow r_e corrections of β
a relative size of $\sqrt{5 \cdot 10^{-5}}$

REALITY \rightarrow fine-structure corrections



not perfect (factor of 10 difference)

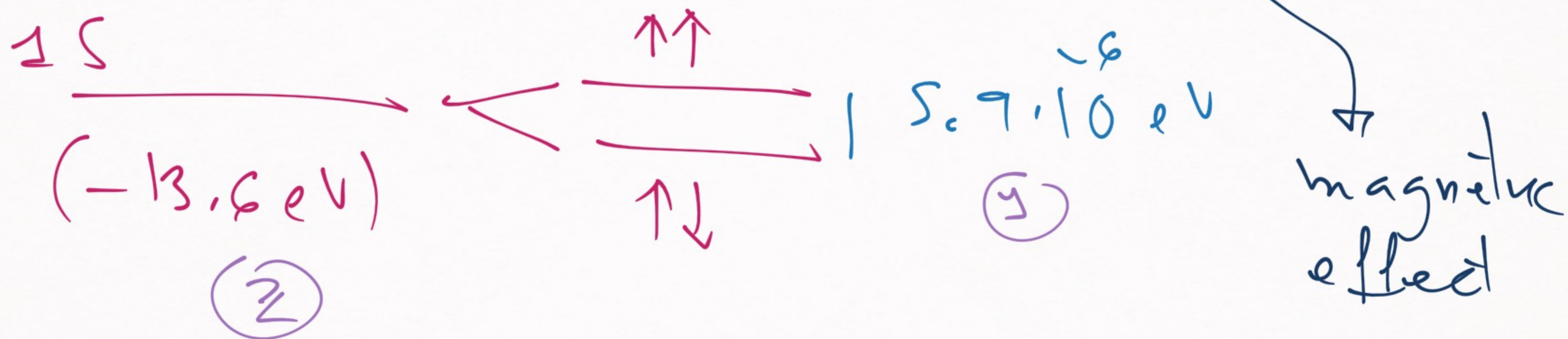
→ but still pretty good

→ \exists -waves, $\left(\frac{1}{2l+1}\right)^2 \sim \left(\frac{1}{3}\right)^2 \sim \frac{1}{9}$

(remember: \exists numerical factors)

CORRECTION 2) proton size correction
 (expectation $\rightarrow (\frac{r_p}{a_0}) \sim 2 \cdot 10^{-5}$)

Hyperfine corrections



LESSONS

Systems that are

1) natural

2) good separation of scapel

NATURAL
SYSTEMS
—

→ [very easy to deal with]

→ good approx. w/ little effort

COROLLARY Systems that

1) are not natural

2) have a poor scale separation

→ **DIFFICULT**

UNNATURAL
SYSTEMS

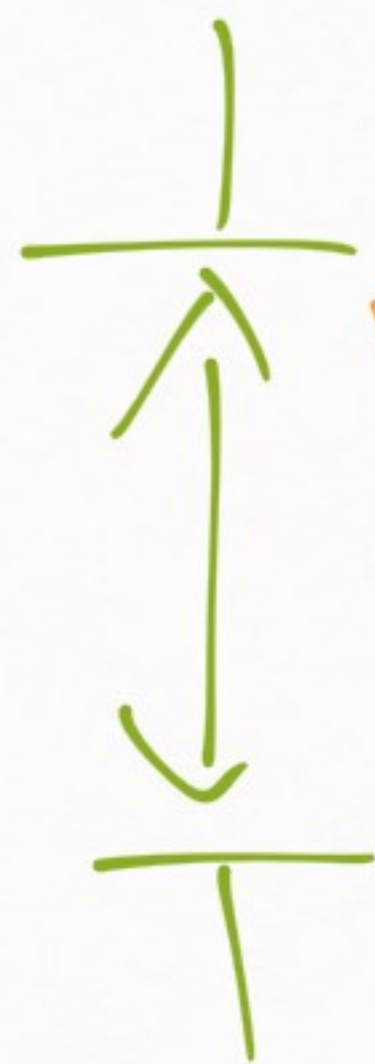
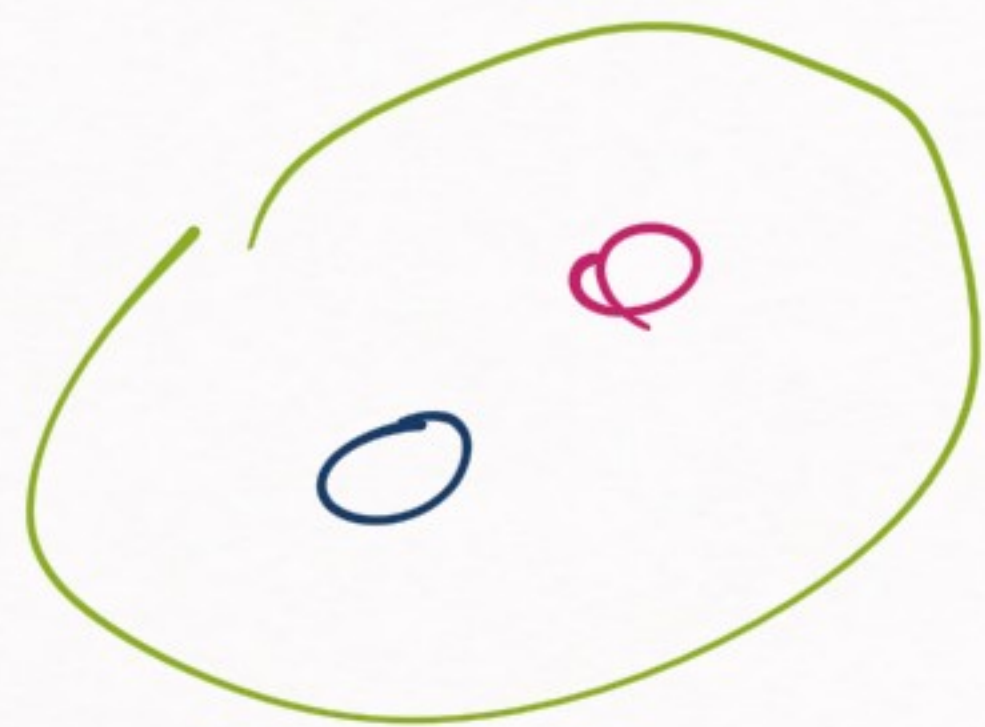
↳ [good for research]

NUCLEAR PHYSICS

NATURAL
or
UNNATURAL

Let's check info
about these nuclear systems:

1) Deuteron \rightarrow np bound state $S=1$
 $I=0$



$r_0 \sim 5 \text{ fm}$

is this big
or small?

2) Range of nuclear forces

$$V_Y(r) = - \frac{g^2}{4\pi} \frac{e^{-mr}}{r}$$

$$V_C(r) = - \frac{e^2}{4\pi} \frac{1}{r}$$

π pion mass

$$m \approx 140 \text{ MeV}$$

$$\frac{1}{m} \approx 1.4 \text{ fm}$$

$$\alpha_Y \approx 15$$

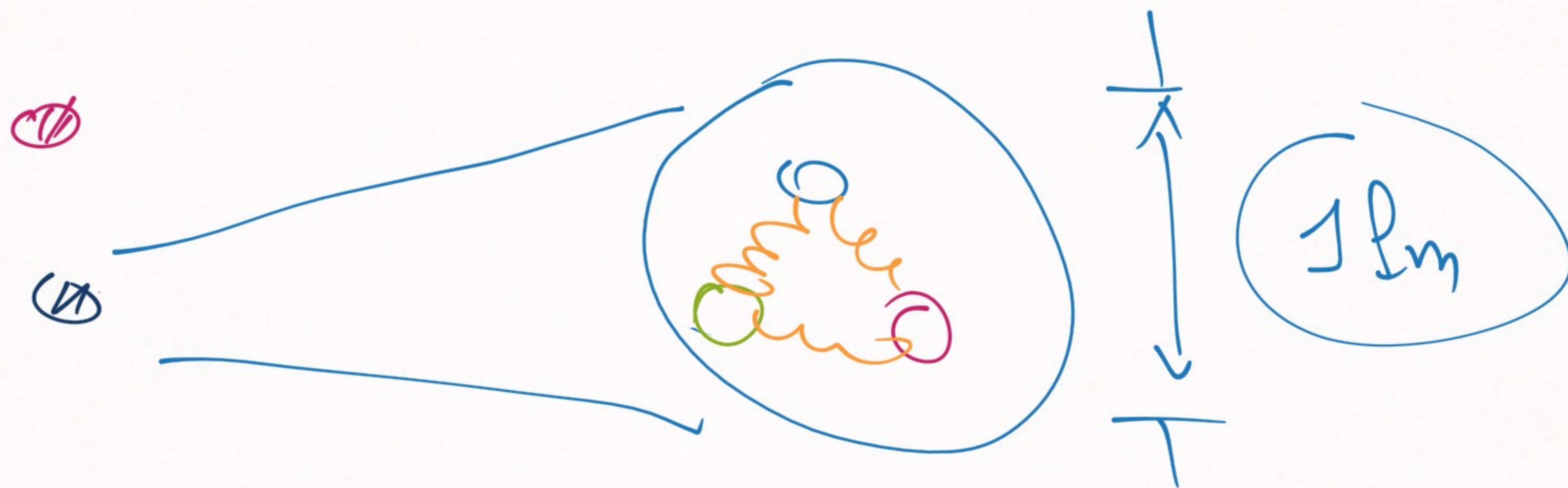
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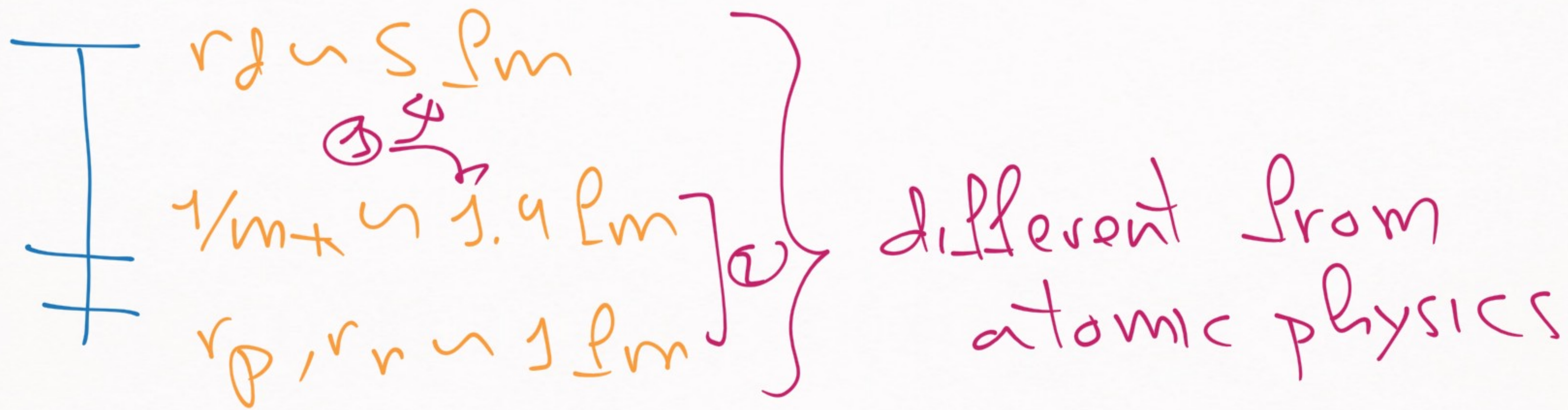
→ wow!!

→ more complex

$1+2) \rightarrow \boxed{r_{d \sim} \sim 3/m\pi} \rightarrow 3 \sim \mathcal{O}(1)$
 "sort of"

3) Other scales \rightarrow size of nucleon r





① not particularly natural

② poor separation of scales

① + ②



NUCLEAR PHYSICS WILL
PROBABLY BE MORE
DIFFICULT THAN
ATOMIC PHYSICS

how much more difficult →

→ we have been studying nuclear forces for 70-80 years

→ yet, we still do not understand nuclear forces well

→ PRETTY DAMN DIFFICULT