

# NUCLEAR PHYSICS

→ Very few: we can change things  
by agreement (wechat group  
/ private message / here)

→ Online classes suboptimal

→ Recommendations :

1) Take notes ( imagine that you were on the classroom & I was writing on the blackboard)

2) Try to make questions

[ 3) Try to check the lessons in advance ]

→ not possible to make exam

→ exercise (

outline

→ easy

- 1) INTRODUCTION (nuclear forces  
→ what are their  
general features)
- 2) TWO-NUCLEON SYSTEM (scattering  
theory)  
↳ dry → interesting, but difficult
- 3) THREE-BODY SYSTEMS (not always)

## 4) NUCLEAR MODELS

↳ easy



- class
- some-time for the lessons
  - some short breaks
  - some time for questions

NUCLEAR PHYSICS (1)



→ [SCALES IN PHYSICS]

尺度

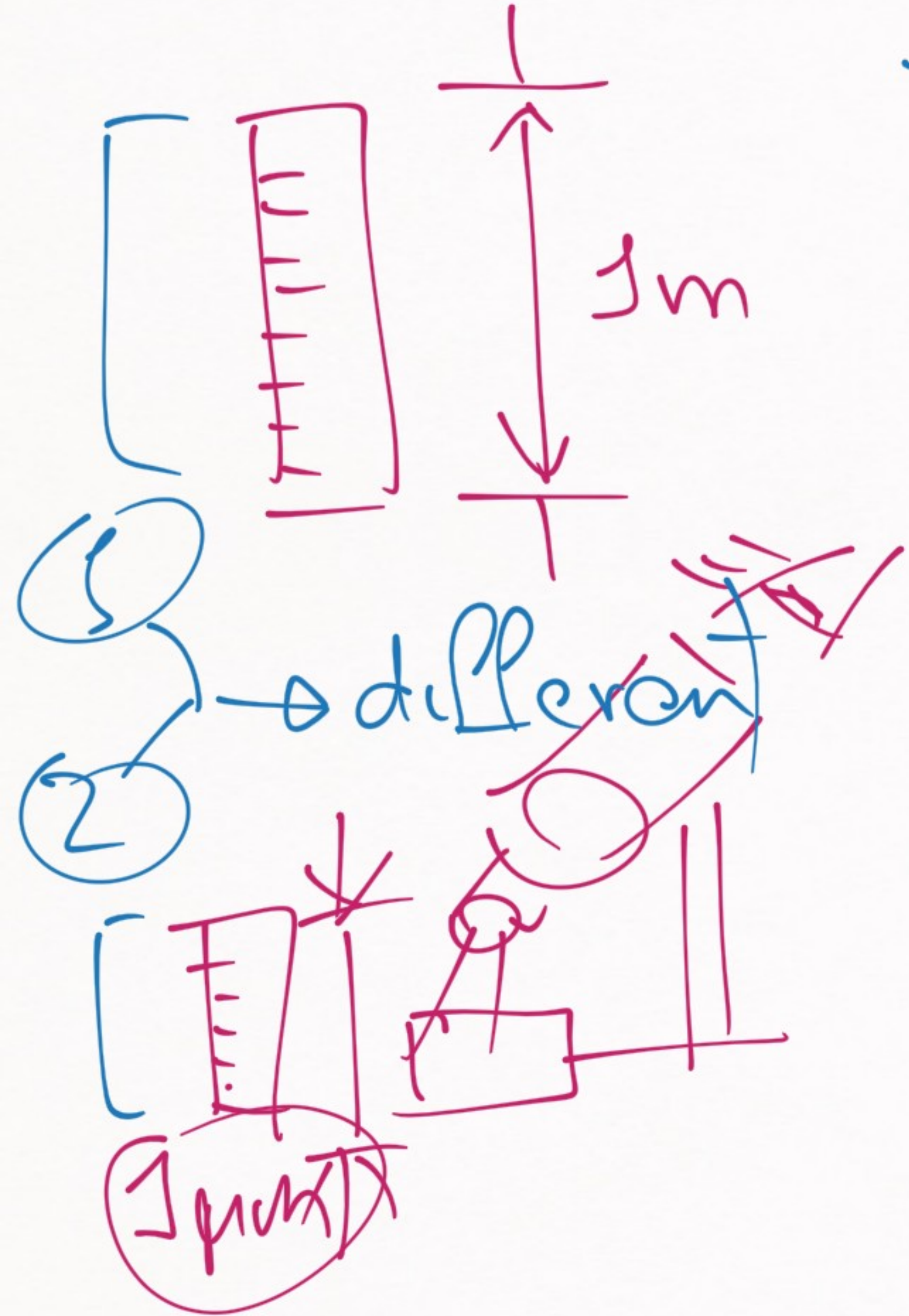
→ [Qualitative discussion]

Doing physics → NUCLEAR PHYSICS  
PARTICLE PHYSICS  
COSMOLOGY  
GENERAL RELATIVITY

[ MANY TYPES  
OF PHYSICS ]

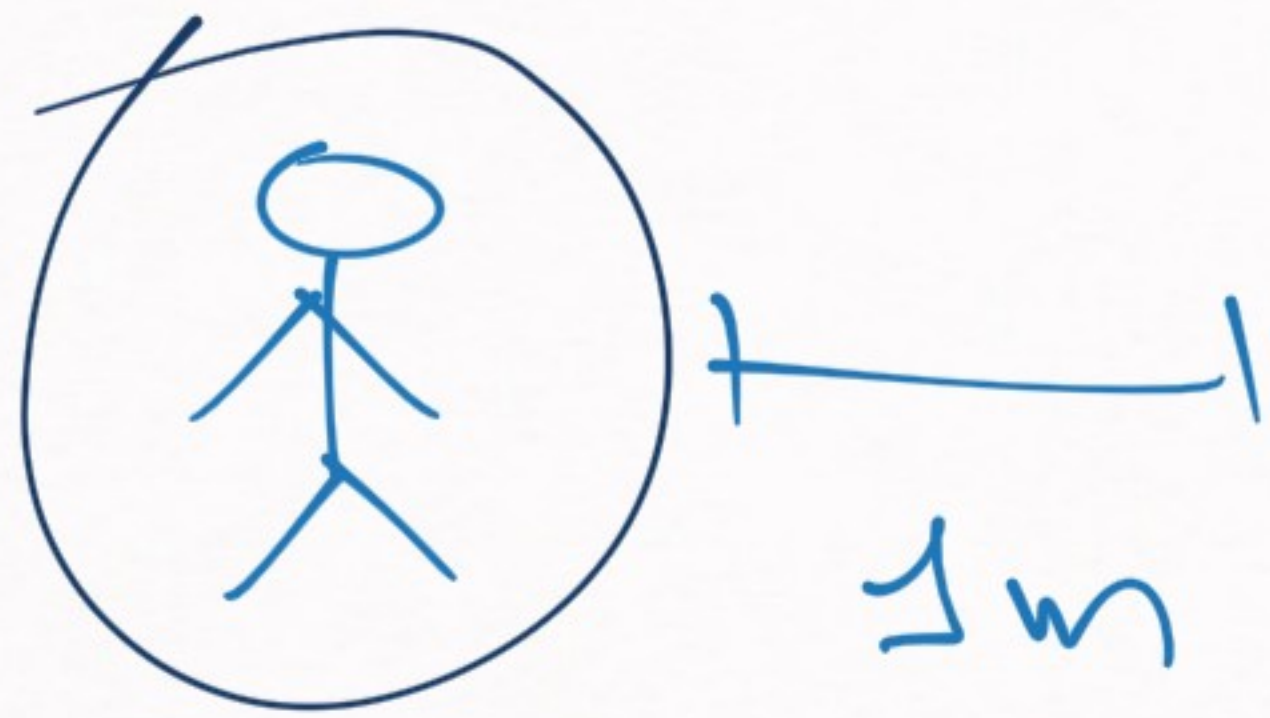
WAY?

OTHER  
SCIENCES



HOW WE SEE THE WORLD  
 DEPENDS ON  
 THE DISTANCE SCALE  
 WE ARE USING  
 TO SEE IT





[ PRETTY OBVIOUS ]  
OBSERVATION

humans → [ natural scale of the order  
of meters ]



(1-2)km small



OBVIOUS OBSERVATION



( $\theta$  depends on  
the distance scale  
we are using  
to see the world)

can be used  
for improving  
our understanding  
of nature



LET'S SEE HOW THIS TRANSLATES  
IN THE APPEARANCE OF  
DIFFERENT TYPES OF  
KNOWLEDGE

CONSIDER  $\Delta$  LENGTH SCALE

→ FIELD OF KNOWLEDGE

FEW  
EXAM-  
PLS

DISTANCE SCALE

$\infty$

$\leq 0$

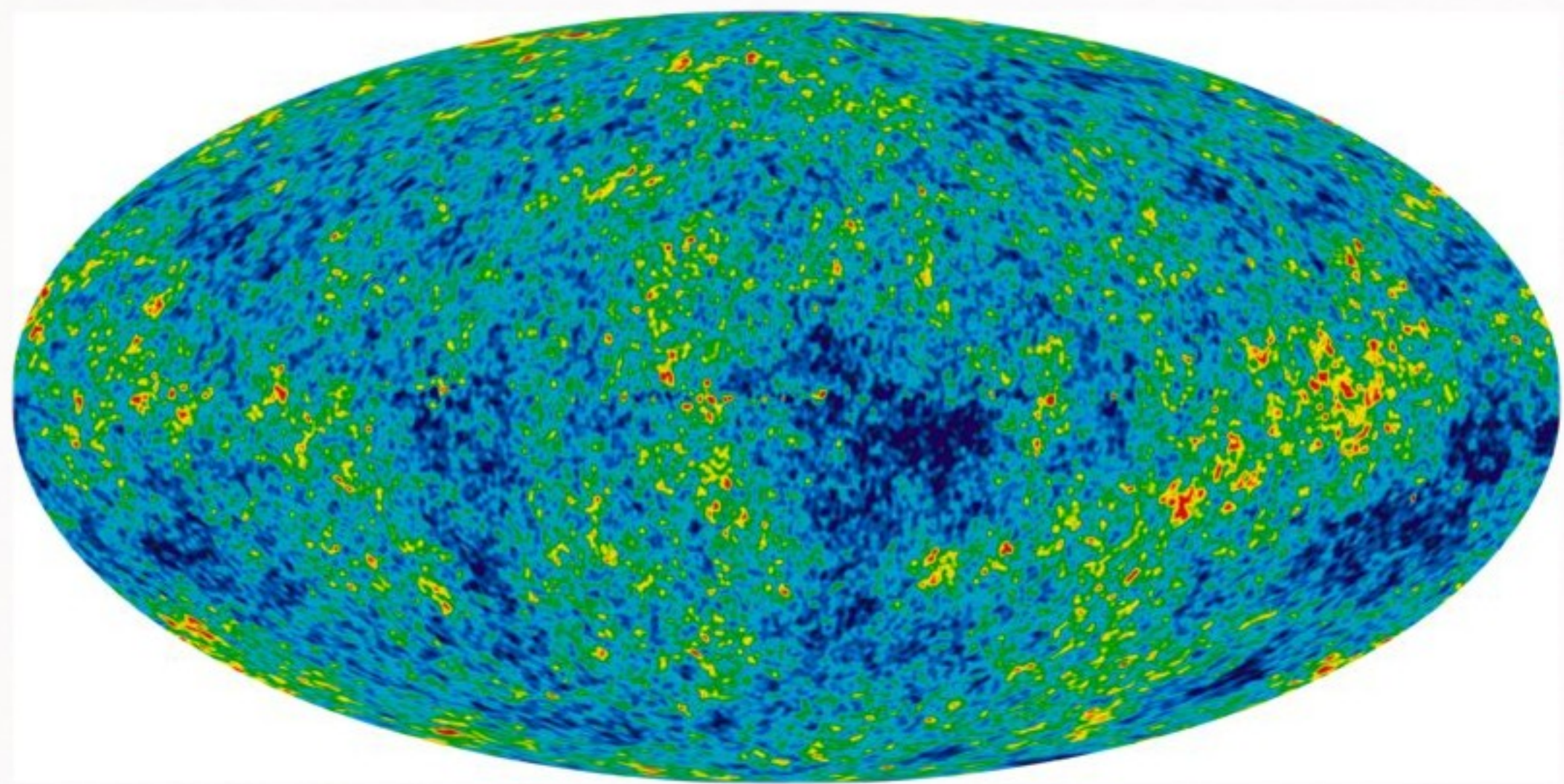


$\infty$  to  $\pi\alpha\upsilon$   
( $\epsilon\upsilon$  to  $\rho\alpha\mu$ ) } all is one

→ METAPHYSICS

(mostly useless speculation)

$10^9$  light years



COSMOLOGY

General structure  
of universe

(GENERAL RELATIVITY,  
EXPANSION OF UNIVERSE)

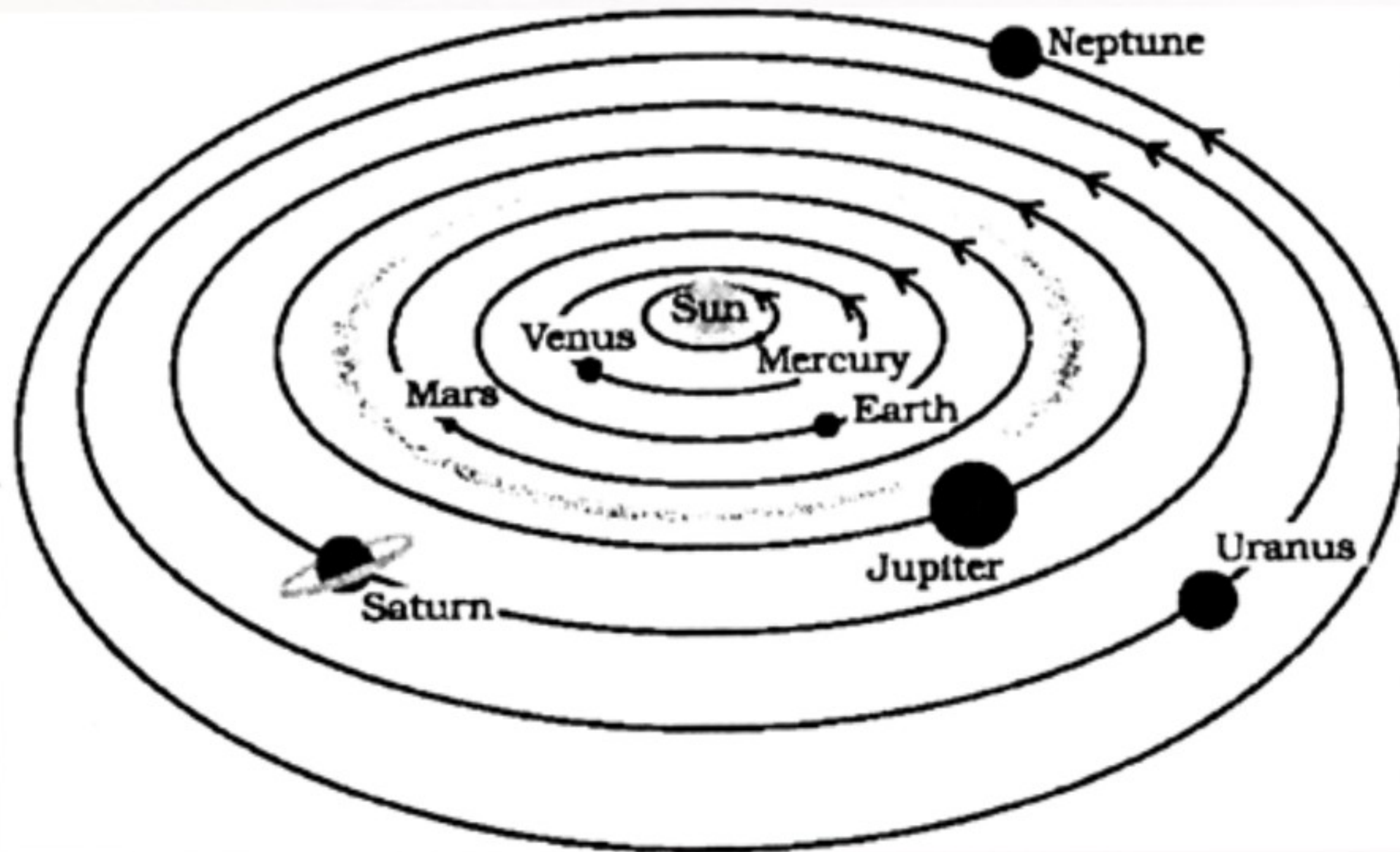
$10^3 - 10^5$  light years  $\rightarrow$  ASTROPHYSICS



$\rightarrow$  we will worry  
about things  
like these  
W

100 a.u. (astronomical units)

→ [size of the solar system]



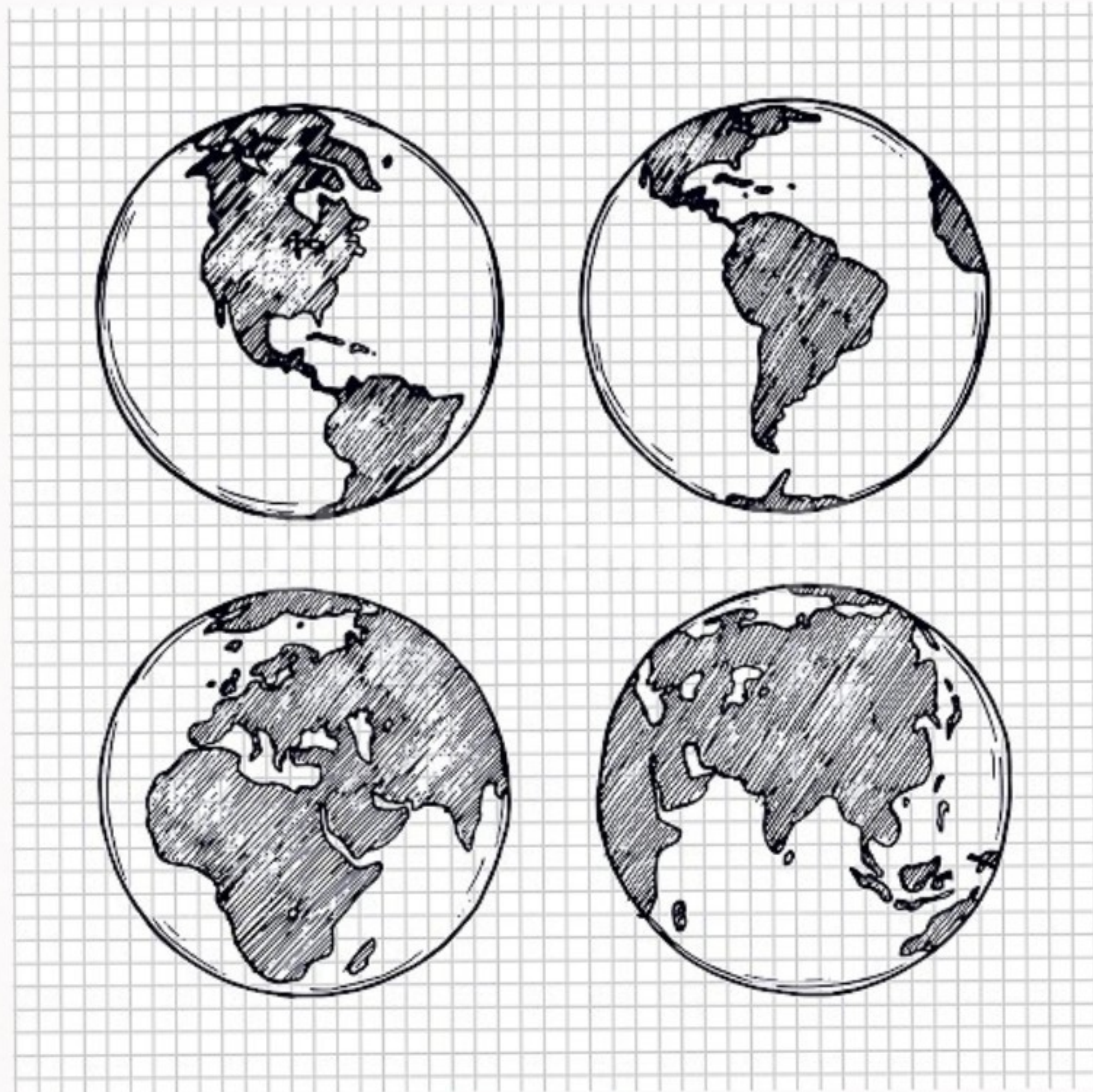
WE WILL WORRY ABOUT

→ CLASSICAL MECHANICS

→ NEWTONIAN GRAVITY

$10^4$  km

→ size of the earth



→ GEOGRAPHY  
→ ECOLOGY  
→ CLIMATE  
SCIENCE



$10^3$  km

→ countries, ecosystems



→ OUR UNDERSTANDING  
OF THINGS  
CHANGES  
AGAIN



1 km



our normal, ordinary lives

→ SOCIAL INTERACTIONS

URBAN PLANNING

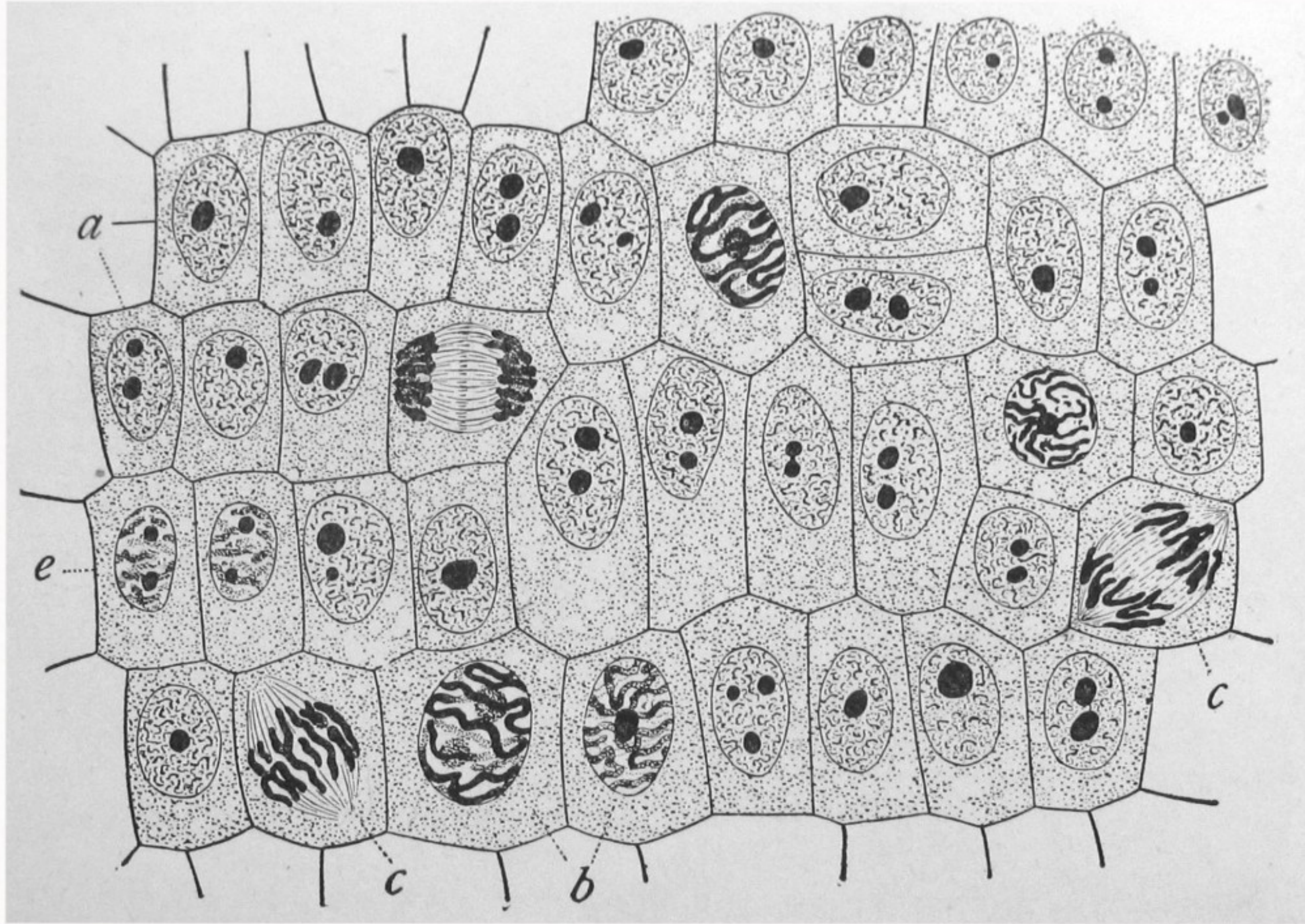
1 m



people, throwing things

(CLASSICAL  
MECHANICS)

$10^{-5} \text{ m}$   $\rightarrow$  cells (BIOLOGY)

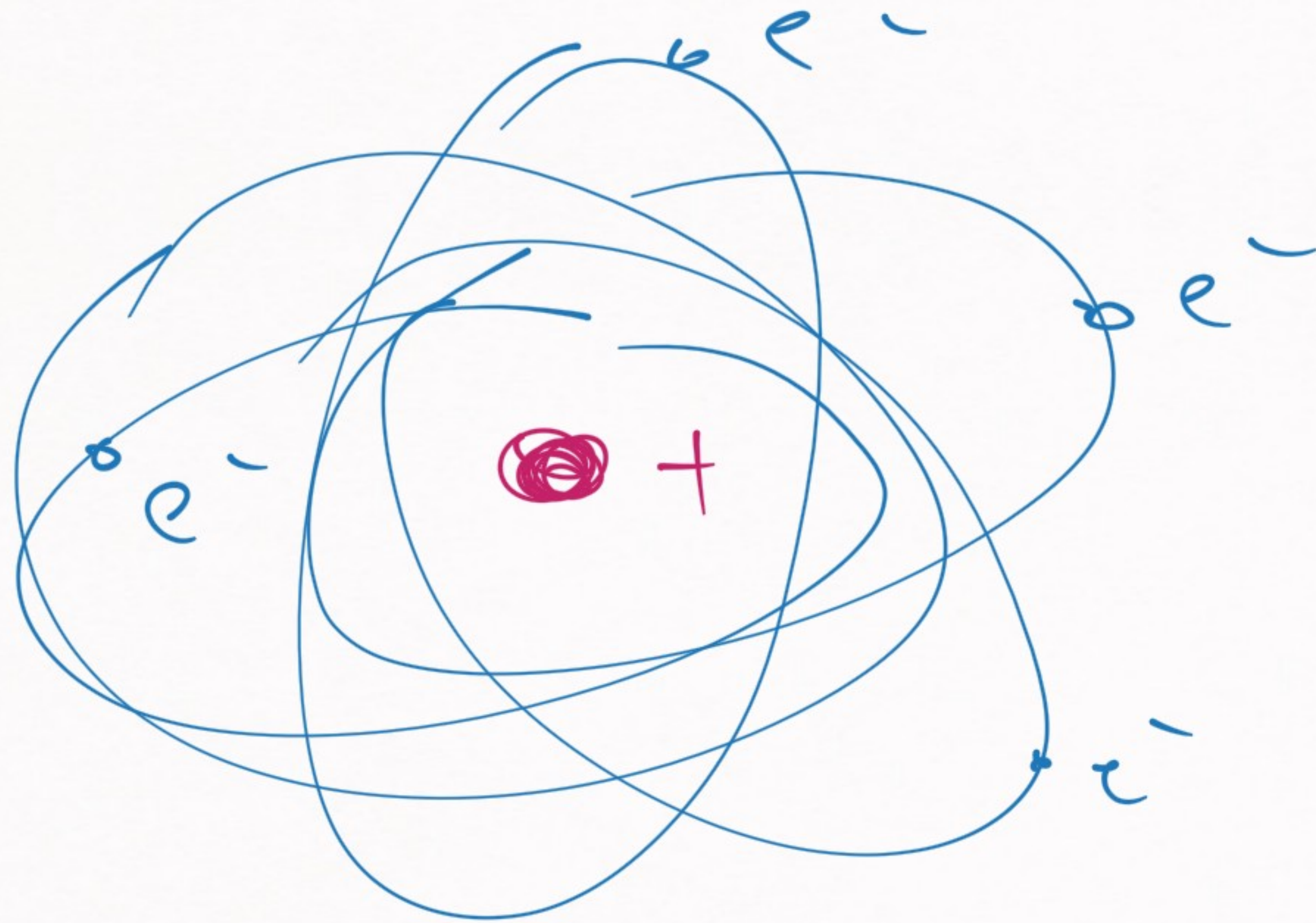


$10^{-9} \text{ m}$   
 $\rightarrow$  molecules  
(CHEMISTRY)

$10^{-10}$  m

→ atoms

→ ATOMIC PHYSICS



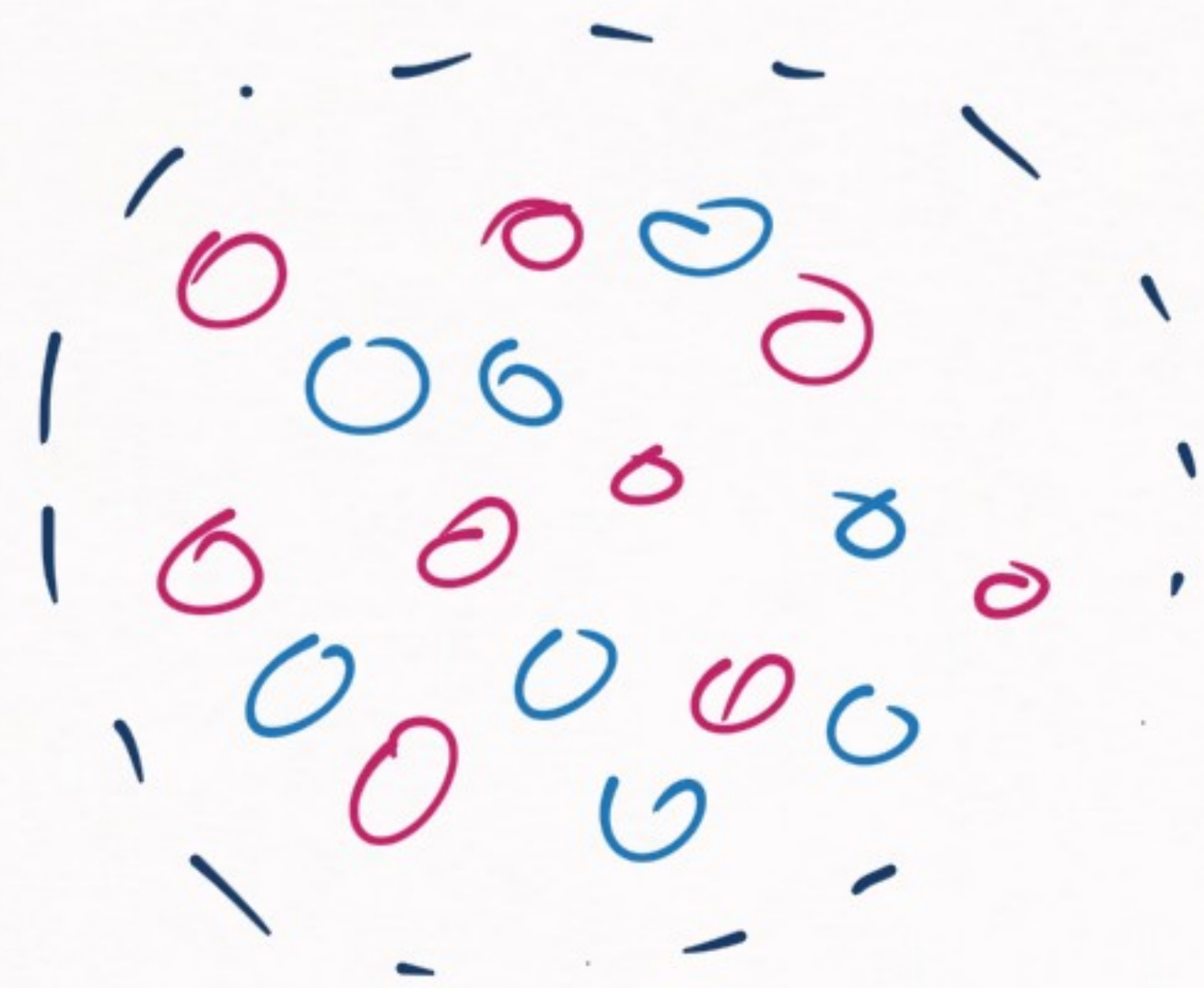
~~CLASSICAL  
MECHANICS~~

NO LONGER  
VALID

QUANTUM  
MECHANICS

$10^{-15} \text{ m}$

→ nuclei, protons, neutrons



→

NUCLEAR  
PHYSICS

↓  
here

$10^{-15} \text{ m} = 1 \text{ fm}$  (fermi)  
≡

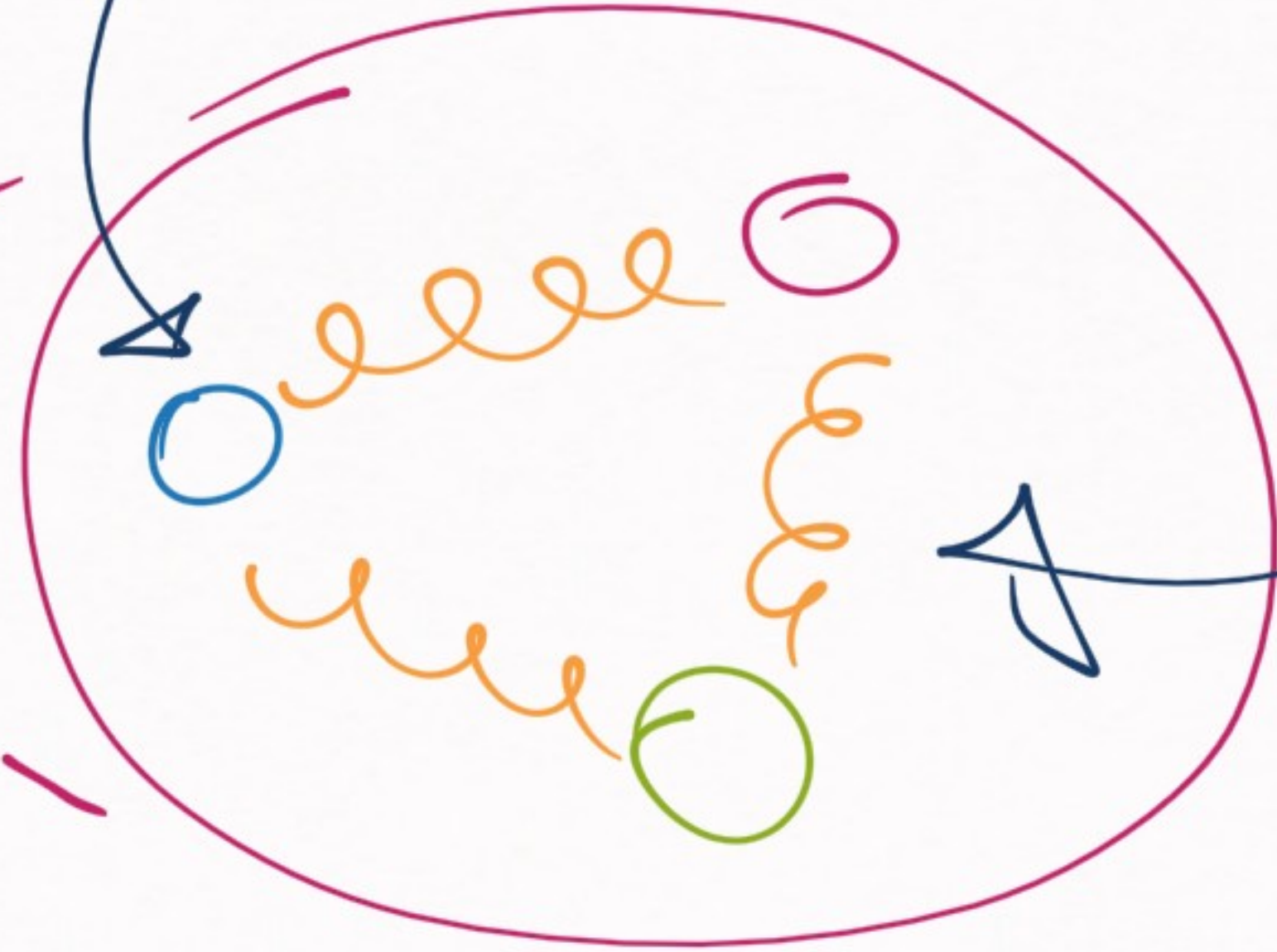
$10^{-16}$  m

quarks & gluons

proton



$10^{-15}$  m



$10^{-16}$  m

QUANTUM  
CHROMO-  
DYNAMICS

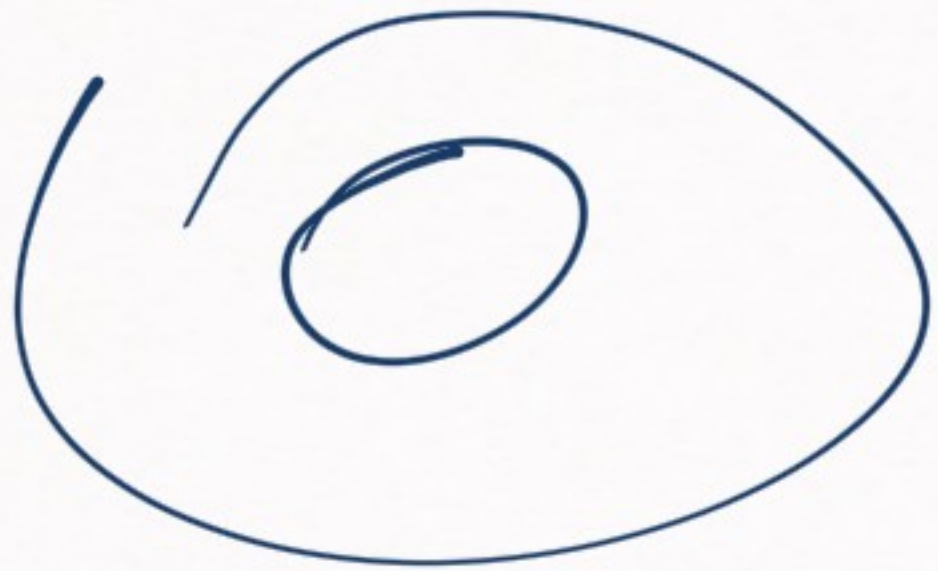
(QCD)

$10^{-35}$  m

→ Planck scale

→

QUANTUM GRAVITY



→

METAPHYSICS AGAIN



CHOOSE  
DISTANCE  
SCALE

⇒

EACH DISTANCE  
SCALE REQUIRES  
A DIFFERENT  
EXPLANATION

→ SOME INDEPENDENCE BETWEEN  
EXPLANATIONS @ DIFFERENT  
SCALE



→ meaning:

⊛ We don't need to know  
QUANTUM MECHANICS to  
understand CLASSICAL MECHANICS  
(but it helps)

⊛ → similar examples

→ [ INDEPENDENCE  
OF EXPLANATIONS ]

↳ (RENORMALIZATION)

↳ [ this isn't only about removing  
infinities ]

→ WE ARE BACK

→ IF I CONTINUE BEYOND  
THE TIME FOR THIS CLASS

TELL ME !!

# PRACTICAL EXAMPLE (SCALES IN PHYSICS)

↳ **ATOMIC PHYSICS**

Hydrogen atom (probably  
seen this when undergraduate)

# Atomic Physics



- 1) Standard approach  
(ground state & its properties)
- 2) Scales-based approach  
(exploit the  $\exists$  natural scale)

## 1) STANDARD APPROACH

1.a) Two-body problem (evident)

1.b) Potential (Coulomb)

1.c) Solve Schrödinger equation

$$1.b) V(r) = -\frac{\alpha}{r}$$

$$V(r) = -\frac{\alpha}{r} \rightarrow \text{Fine structure constant}$$

distance

$$\alpha = \frac{e^2}{4\pi} \approx \frac{1}{137}$$

units


→

[E]  
[L]

} interchangeably

$$\text{conversion factor} \rightarrow \hbar c = 197.3 \text{ MeV}\cdot\text{fm}$$

$$V(r) = - \frac{\alpha}{r}$$

$\underbrace{\hspace{1cm}}_{\text{MeV}}$ 


$\left. \begin{matrix} \text{MeV} \\ \text{fm} \end{matrix} \right\} \rightarrow \text{natural for nuclear physics}$

$$r = 1 \text{ fm} \Rightarrow V(r) = -\alpha (1 \text{ fm})^{-1}$$

$$\Rightarrow V(r) = -\alpha (1 \text{ fm})^{-1} E_c$$

$$= -\alpha (200 \text{ MeV})$$

$\left. \begin{matrix} \text{eV} \\ \text{\AA} \end{matrix} \right\} \rightarrow \text{also possible}$

$$E_c = 1973 \text{ \AA eV}$$

(check)



Units  $\rightarrow$  fm, MeV

use  $hc$  as a  
conversion factor



$$V(r) = -\frac{\alpha}{r} \rightarrow \text{Coulomb potential}$$

# 1.c) Schrödinger equation

$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E \psi(\vec{r})$$

reduced mass  $\left\{ \rightarrow \frac{1}{\mu} = \frac{1}{m_e} + \frac{1}{m_p} \approx \frac{1}{m_e} \right.$

Standard approach  $\rightarrow$  solve Schrödinger

$$\left[ -\frac{\hbar^2}{2\mu} \nabla^2 + V(\vec{r}) \right] \Psi(\vec{r}) = E \Psi(\vec{r})$$



$$\left[ -\nabla^2 + 2\mu V(\vec{r}) \right] \Psi(\vec{r}) = 2\mu E \Psi(\vec{r})$$

$$U(\vec{r}) = 2\mu V(\vec{r})$$

"reduced potential"

$$2\mu E = \gamma^2$$

wave number

$$[-\nabla^2 + 2\mu V(\vec{r})]\Psi(\vec{r}) = -\gamma^2 \Psi(\vec{r}) \Rightarrow \text{D} \oplus$$

$$2\mu V(\vec{r}) = -2\mu \frac{\alpha}{r} = -\frac{2}{a_B r}$$

$$a_B = \frac{1}{\mu \alpha} \approx \frac{1}{m_e \alpha}$$

length scale  
(Bohr radius)

$$a_B \approx \frac{137}{0.5 M_{\mu}} \approx 270 \text{ MeV}^{-1} \approx 54000 \text{ fm}$$

$\hbar c \approx 200 \text{ MeV fm}$

$$\textcircled{*} \Rightarrow \left[ -\nabla^2 - \frac{2}{a_B r} \right] \Psi(\vec{r}) = -\gamma^2 \Psi(\vec{r})$$

$$a_B = \frac{1}{\mu \alpha} \approx \underline{\underline{5.4 \cdot 10^4 \text{ fm}}}$$

$$\left[ \nabla^2 + \frac{2}{a_B r} \right] \Psi(\vec{r}) = \gamma^2 \Psi(\vec{r})$$

↳ we can solve it

$$\underline{\Psi}(\vec{r}) = \frac{1}{\sqrt{4\pi}} \frac{2}{a_B^{3/2}} e^{-r/a_B} \quad , \quad \gamma = \frac{1}{a_B}$$

→ ground state solution of  
the hydrogen atom

$$E_B = -\frac{\gamma^2}{2M} = -\frac{1}{2M} \left(\frac{1}{a_B}\right)^2 \approx \underline{\underline{-13.6 \text{ eV}}}$$

$\sqrt{\langle r^2 \rangle}$   $\rightarrow$  mean square radius of  
the hydrogen atom

$$\langle r^2 \rangle = \langle \psi | r^2 | \psi \rangle = \int d^3\vec{r} \ r^2 |\psi(\vec{r})|^2$$

$$= \underline{\underline{3a_0^2}}$$

$$\Rightarrow \underline{\underline{\sqrt{\langle r^2 \rangle} = \sqrt{3} a_0}}$$

# [STANDARD APPROACH]

Potential  $\rightarrow$  Schrödinger equation  $\rightarrow$  Solution

$$E_B = -\frac{1}{2\mu} \gamma^2, \quad \left[ \gamma = \frac{1}{a_B}, \quad \sqrt{\langle r^2 \rangle} = \sqrt{3} a_B \right]$$

non-relativistic  
theory

powers of  $a_B$



( SCALES -BASED APPROACH )

→ method for lazy people



this can be  
a good thing

$\hat{O} \rightarrow$  observable

$$\langle \hat{O} \rangle = \langle \psi | \hat{O} | \psi \rangle$$

$|4\rangle \rightarrow$  Hydrogen atom wave function

dimensions of  $[L]^3$

Input  $\rightarrow$

$$\left[ \nabla^2 + \frac{2}{a_0 r} \right] \psi(\vec{r}) = E \psi(\vec{r})$$

$\Rightarrow$  [ THE INPUT ONLY CONTAINS  
ONE PARAMETER ]

$\searrow$   $a_B$

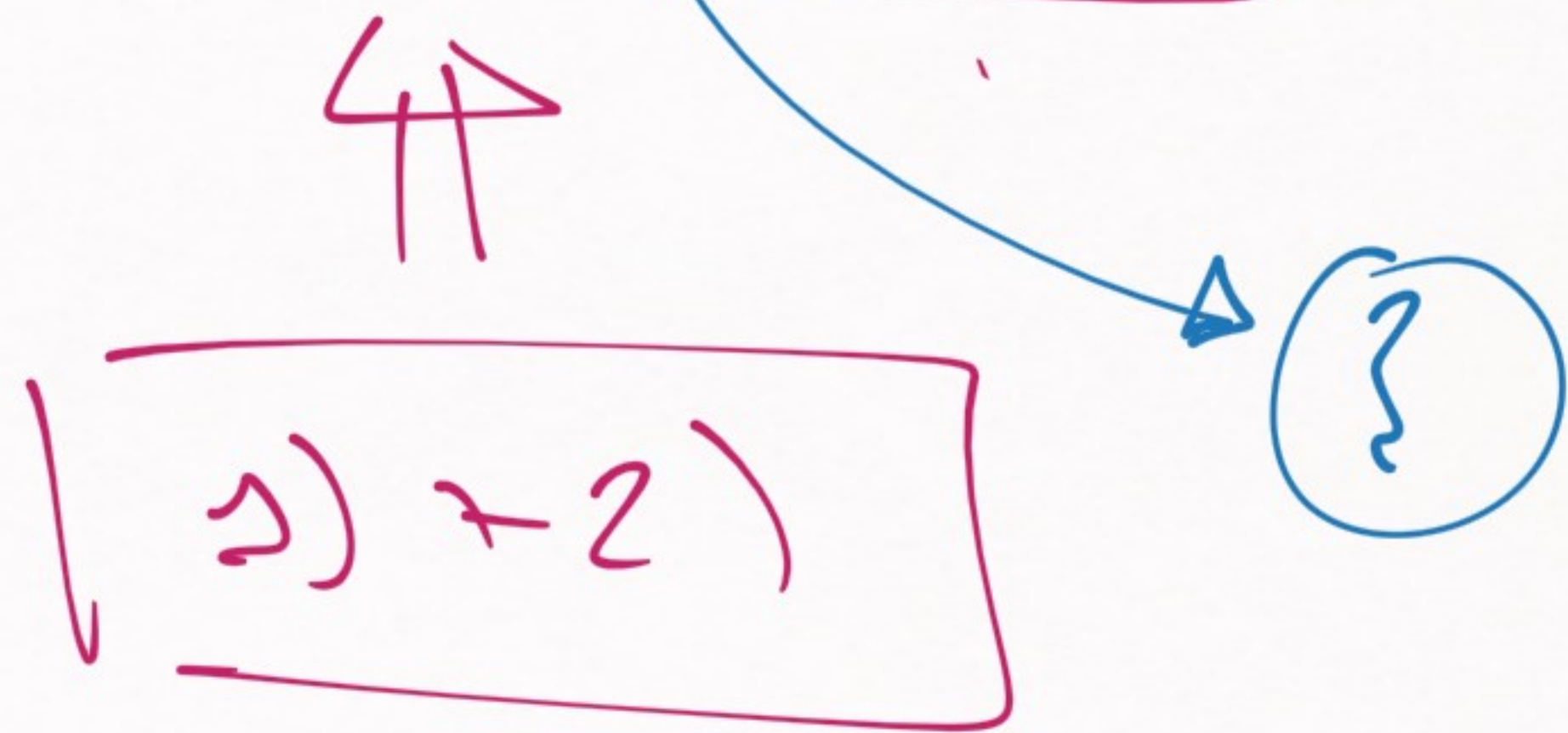
$$[\langle \hat{O} \rangle] = [L]^n \Rightarrow \boxed{\langle \hat{O} \rangle \sim \text{const} \times a_B^n}$$

$[L]$   $\rightarrow$  dimensions (length, energy)

1)  $[\langle \hat{O} \rangle] = [L]^n \Rightarrow$

$$\langle \hat{O} \rangle = c_0 \hat{O} a_n$$

2)  $a_B$  is the only input parameter in the Schrödinger equation



$\textcircled{?}$   $\rightarrow$  what is the size of  $\underline{c_0}$

SCALES



$$\langle \hat{O} \rangle = \langle \hat{C}_0 \rangle a_0^s$$



NATURALNESS

[size of this]



(assumption) →

$$C_0 \sim \mathcal{O}(1)$$

physical problems  
are not always  
natural

→ back to the STANDARD APPROACH

$$\gamma = \frac{1}{a_B}$$
$$\sqrt{\langle r^2 \rangle} = \sqrt{3} a_B$$

EXPLICIT  
CALCULATION

SCALE APPROACH

$$\gamma = \frac{c}{a_B}, \quad c \sim G(1)$$

$$\sqrt{\langle r^2 \rangle} = d a_B, \quad d \sim G(1)$$

$$\gamma = \frac{c}{a_B} \rightarrow c = \underline{1} \sim \mathcal{O}(1) \quad \checkmark$$

$$\sqrt{\langle r^2 \rangle} = d a_B \rightarrow d = \underline{\sqrt{3}} \sim \underline{1.7} \sim \mathcal{O}(1) \quad \checkmark$$



↓ SCALES + NATURALNESS →

VERY GOOD APPROX. TO ANY PROBLEM

1) Physics depends on a characteristic  
scale

+  
2) Physics is natural (in cases)

1) + 2)



good estimations to  
a lot of problems



→ PROBLEM →

NUCLEAR PHYSICS



GOOD: WORK  
FOR PHYSICISTS

ASSUMPTION OF  
NATURALNESS  
WILL FAIL

# EXAMPLE

$$\left[ \begin{array}{c} s \\ p \end{array} \right] \text{ system}$$
$$s = 1/2$$

hp

$\rightarrow \left[ \begin{array}{l} s=0 \\ s=1 \end{array} \right. \rightarrow \text{virtual state (quasi bound)}$   
 $\rightarrow \text{deuteron (bound state)}$

$$\left[ \gamma_0, \gamma_1 \right] (\gamma_s)$$

1) What does natural means  
in nuclear physics?

ATOM

$$V(r) = -\frac{q}{r}$$

$$q = \frac{e^2}{4\pi}$$



$$V_{np}(r) = -\frac{g^2}{4\pi} \frac{e}{r}$$

$$m$$

$$g = \frac{1}{m_p}$$

$$m$$

→ this one

→ usual solution is to use  $m$

(solution 2 →  $a_Y \approx \frac{1}{\Lambda_Y}$ ,  $\Lambda_Y = 300 \text{ MeV}$ )

$m \leq 140 \text{ MeV}$

→ convention

$$\gamma_0 \sim \mathcal{O}(1) \times m\pi$$

$$\gamma_1 \sim \mathcal{O}(1) \times m\pi$$

natural  
solution

$$\gamma_0 \sim 8MeV$$

$$\gamma_1 \sim 45MeV$$

$$\frac{m\pi}{17} \rightarrow \frac{1}{17} \ll 1$$

$$\frac{m\pi}{3} \rightarrow \frac{1}{3} \lesssim 1$$

FOR WATER

→ ATOMIC PHYSICS : NATURAL

→ NUCLEAR PHYSICS : NON NATURAL  
(FINE TUNED)

↓  
=

DIFFICULT

EASY

END FOR  
TODAY

→ FRIDAY (IS: SO)

