

NUCLEAR PHYSICS 16



Tensor forces



RECAP

→ Summary of T-matrix

- 1) T-matrix $\rightarrow T = V + VG_0T$ (LSE)
- 2) T-matrix is easy to
solve analytically for a few selected
potentials $\langle \vec{p}' | V | \vec{p} \rangle = \lambda g(\vec{p}')g(\vec{p})$
(separable potentials)

$$2') \langle \vec{p}' | T(\epsilon) | \vec{p} \rangle = \tau(\epsilon) g(\vec{p}') g(\vec{p}) \quad \text{ansatz for separable potentials}$$

\downarrow
 $\tau(\epsilon) = \frac{1}{1 - I(\epsilon)} , \quad I(\epsilon) = \frac{\beta e}{(2\pi)^3} \frac{g^2(\epsilon)}{E - \frac{e^2}{2m}} \sum$

— \otimes —

$$3) \langle p' | v | p \rangle = \underbrace{\langle_0 \langle \lambda | S(\vec{x}') | \vec{p}(\vec{x}) \rangle}_{\downarrow}$$

contactorange
potential + separable
regulator

$$f(x) \xrightarrow[x \rightarrow 0]{} 1 \quad \text{and} \quad f(x) \xrightarrow[x \rightarrow \infty]{} 0$$

3') $\Lambda \rightarrow$ auxiliary parameter

(a parameter of the theory,

but not a physical parameter)

$$\frac{d}{d\gamma} \tau(\xi; \Lambda) \leq 0 \rightarrow C_0 = C_0(\Lambda)$$

Bound states w/ T-matrix

$\Leftrightarrow \tau(\epsilon)$ has poles at $\epsilon \rightarrow E_B$

$$\tau(\epsilon) \rightarrow \frac{\text{Res } \tau(\epsilon)}{\epsilon - E_B}$$

$$\begin{aligned} \text{Res } \tau(\epsilon) &= V |B\rangle \langle B| V \\ &= G^1(E_B) |B\rangle \langle B| G^1(E_B) \end{aligned}$$

$\epsilon |B\rangle = G_0(\epsilon) V |B\rangle \hookrightarrow \text{bound state equation}$
(BSE)

$$\Rightarrow \langle \vec{p}' | V | \vec{p} \rangle = g(\vec{p}) g(\vec{p}') \rightarrow \text{separable } V$$

$$\langle \vec{p}' | B \rangle = \mathcal{N} \frac{g(p)}{p^2 + \gamma^2}$$

$$E_B = -\frac{\vec{p}^2}{2m}$$

$$\left(\frac{d^3 \vec{p}'}{(2\pi)^3} |\langle \vec{p}' | B \rangle|^2 \right) = 1 \rightarrow \text{obtain } \mathcal{N}$$

B) Pro tip: use the vertex functions
(instead of the wave function)

$$|\psi_B\rangle = C_0(\epsilon) |\phi_B\rangle$$

8.a) $\tau(\epsilon) \rightarrow$ $\epsilon \rightarrow \epsilon_B$ $\frac{|\phi_B\rangle \langle \phi_B|}{E - E_B}$

8.b) $|\phi_B\rangle = V C_0 |\phi_B\rangle \rightarrow$ easier \approx

$$3') |B\rangle = G_0 |\beta\rangle \quad \xrightarrow{\text{real basis}}$$

$$\Leftrightarrow \phi_B(\vec{q}) = \frac{1}{(2\pi)^3} \int \frac{d^3 \vec{e}}{(2\pi)^3} \langle \vec{p} | V | \vec{e} \rangle \psi_B(\vec{e})$$

$E_B - \frac{e^2}{2\mu}$

$$|\phi_B\rangle = V G_0 |B\rangle$$

easier

$$\phi_B(\vec{p}) = \int \frac{d^3 \vec{e}}{(2\pi)^3} \langle \vec{p} | V | \vec{e} \rangle \psi_B(\vec{e})$$

$E_B - e^2/2\mu$

9) More about the T-matrix

9.a) Partial wave expansion

9.b) Solving it by means of
discretization methods

⇒ Advance topics (covered in lecture
notes)

→ This is enough about the Tundra
at the moment



WHAT WAS THE TENSOR FORCE? (previous
lessons)

- 1) Yukawa's idea \rightarrow exchange of
a scalar meson

$$V_Y(r) = -g^2 \frac{\bar{e}^{mr}}{4\pi r}$$

- 2) fit the deuteron plus a quadrupole
moment; $Q_D \approx 0.26 \text{ fm}^7 e$

3) Nuclear forces are not central at long distances

→ pseudoscalar particle

$$g_A \approx 1.26$$

(pion)

$$\rho_\pi \approx 92.4 \text{ MeV}$$

$$m_\pi \approx 138$$

MeV

$$V_{\text{ext}}(\vec{q}) = -\frac{g_A}{4\rho_\pi} \bar{\psi}_L \bar{\psi}_R \frac{\bar{\sigma}_\mu \cdot \vec{q} \bar{\sigma}_\nu \cdot \vec{q}}{\vec{q}^2 + m_\pi^2}$$

$$\bar{s}_1 \bar{q} \bar{s}_2 \bar{q} \rightarrow \frac{1}{3} \bar{s}_1 \bar{s}_2 q^2 + (\bar{s}_1 \bar{q} \bar{s}_2 \bar{q} - \frac{1}{3} \bar{s}_1 \bar{q})$$

Deuteron will have
a small D-wave
component

Tensor force

b
mixes different
partial waves

→ We can Fourier-transform the previous potential & obtain:

$$V_{\text{DFT}}(\vec{r}) = \frac{g_A^2 m_\pi^3}{48\pi e_\pi^2} \bar{\epsilon}_1 \bar{\epsilon}_2 \times [\bar{\sigma}_1 \cdot \bar{\sigma}_2 W_c(m_\pi r) + S_1(r) W_7(m_\pi r)]$$

$W_c(x), W_t(x) \rightarrow$ certain functions of $x = m\pi r$

Bottom-line is : $V_{\sigma_1 \sigma_2}(\vec{r}) \propto \overline{\sigma}_1 \cdot \overline{\sigma}_2 W_c + S_n(r) W_t$

$$\textcircled{1} \quad \overline{\sigma}_1 \cdot \overline{\sigma}_2 = \begin{cases} 1 & S=1 \\ -3 & S=0 \end{cases} \quad \begin{matrix} \text{spin-spin} \\ \text{term} \end{matrix}$$

(1)

$\underbrace{\qquad}_{\text{tensor term}}$

$\underbrace{\qquad}_{\text{term}}$

(2)

(S is the total spin)

$$\textcircled{2} \quad S_{12}(\vec{r}) = 3\vec{\sigma}_1 \cdot \hat{r} \vec{\sigma}_2 \cdot \hat{r} - \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

$\overbrace{\quad}$ tensor operator will generate
some complications

$[S_{12}(\vec{r}), \vec{r}^2] \neq 0 \rightarrow$ it does not
conserve
angular momentum

Luckily we have the following:

$$\left[\vec{S}_{12}(\vec{r}), \vec{S}^2 \right] = 0 \quad \left[\vec{S}_{17}(\vec{r}), \vec{J}^2 \right] = 0$$

} total spin S
Total angular momentum
are conserved

$$\vec{J} = \vec{L} + \vec{S}'$$

\rightsquigarrow

only true for $S \leq 2$

Simplification \rightarrow the tensor force only acts
on triplets ($S=1$)

\Rightarrow Singlet wave ($S=0$)

$$\langle 001 | S_{12}(\vec{r}) | 100 \rangle = 0 \rightarrow \text{Why?}$$

$S_{12}(\vec{r}) \rightarrow$ symmetric w/ respect to
changing particles 1 & 2

$$S_{12}(\vec{r}) = 3\sigma_1 \cdot \vec{\hat{r}} \cdot \vec{\sigma}_2 \cdot \vec{\hat{r}} - \vec{\sigma}_1 \cdot \vec{\sigma}_2 = S_{21}(\vec{r})$$

$$|00\rangle = \frac{1}{\sqrt{2}} [|+\downarrow\rangle - |-+\rangle]$$

\downarrow \downarrow
 1 2

$S_{12}(\vec{r}) |00\rangle_s = 0$

1 ↲ > 2

$$|00\rangle_s \rightarrow -|00\rangle_s$$

↑

$$S_{12}(\vec{r}) |00\rangle_s = -S_{12}(\vec{r}) |100\rangle_s$$

3') For singlets ($S=0$) the tensor force vanishes

$$\text{---} \otimes \text{---}$$

2) Triplet waves \rightarrow no simplification possible

$[S_{12}(\vec{r}), \vec{\gamma}'] = 0 \rightarrow$ construct state

w/ good $\vec{\gamma}'$

2.a) $\sum_{e\text{me}}(r)|sm\rangle$ (w/o tensor force)

↓
good less (but not good J)

2.b) $|jm\rangle = \sum_{\substack{sm \\ e\text{me}}} \sum_{e\text{me}}(r)|sm\rangle$
 $\times \langle e\text{me } sm | jm \rangle$

Clebsch-Gordan coefficient

$$3) |j_m\rangle = |(s\ell)_{jm}\rangle$$

→ matrix element of $S_{12}(\vec{r})$ w/ these states

$$\langle (s'e')_{j'm'} | S_{12}(\vec{r}) | (s\ell)_{jm} \rangle = \delta_{s'se}^j \delta_{jj'} \delta_{mm'} \quad \text{if } \langle S_{12}(\vec{r}), j' \rangle = 0$$

For $NN \rightarrow S=S'=1$ (we can ignore the S' part)

$$\langle (s e') j' m' | s_n(r) | (1l) jm \rangle = \sum_s \delta_{e e'} \delta_{jj'} \delta_{ss'}$$

$$e' = e, e \pm 1$$

$$(e' \neq e \pm 1)$$

breaks
parity

EXAMPLE : DEUTERON

i) No tensor force $\rightarrow \psi_d(\vec{r}) = \frac{\phi(r)}{r} Y_{00}(\hat{r}) |S\text{md}\rangle$

angular momentum part

spin part

DECOUPLED

The diagram illustrates the decomposition of the deuteron wavefunction. It shows the wavefunction $\psi_d(\vec{r})$ as a product of a radial part $\frac{\phi(r)}{r}$, a spherical harmonic function $Y_{00}(\hat{r})$, and a state $|S\text{md}\rangle$. The angular momentum part, consisting of the radial part and the spherical harmonic, is grouped by a curly brace. This part is further divided into an 'angular momentum part' (the radial part) and a 'spin part' (the spherical harmonic). Below this, another curly brace groups the 'angular momentum part' and the 'spin part', with the word 'DECOUPLED' written below it, indicating that they are independent components.

2) Tensor Force i

$$\psi_d(\vec{r}) = \frac{u(r)}{\sqrt{r}} \sum_{l=0}^{\infty} |j_l m_l\rangle$$

S-wave component

$$+ \frac{w(r)}{\sqrt{r}} \sum_{ms} \sum_{l=0}^{\infty} |j_{lm}s\rangle$$

m_s
 m_l

$$\times \langle j_{lme} | j_m | j_{lmd} \rangle$$

D-wave component

→ we can calculate the matrix elements
of $S_{12}(\vec{r})$

$$|S(s_{md})\rangle = Y_{00}(\vec{r}) |s_{md}\rangle$$

$$|D(1_{md})\rangle = \sum T_{2ms}(\vec{r}) |1_{md}\rangle$$

$$\times \langle 2_{ms} 1_{ms} | s_{md} \rangle$$

$\mathcal{B} = \{ |S(1m) \rangle, |D(1m) \rangle \}$



$$S_2 = \begin{pmatrix} 0 & 2\sqrt{2} \\ 2\sqrt{2} & -2 \end{pmatrix} \rightarrow S_2(\hat{r}) \text{ evaluated in the deuteron}$$

Schrödinger equation:

$$1) \text{ w/o } S_{12}(r) \quad -\psi'' + 2\mu V_0(r) = -\gamma^2 \psi(r)$$

$$2) \text{ w / } S_{12}(r)$$

$$-\psi''(r) + 2\mu [V_C(r) + 2\sqrt{2}V_I w(r)] = -\gamma^2 \psi(r)$$

$$-\psi''(r) + \frac{6}{\sqrt{2}} w(r) + 2\mu [2\sqrt{2}V_I \psi(r)$$

$$+ (V_C - 2V_I) w(r)] = -\gamma^2 w(r)$$

Convention: $V = V_C \vec{\alpha}_1 \cdot \vec{\alpha}_2 + V_T S_{12}(\vec{r})$

TENSOR FORCE
S PARITY

$$\rightarrow L \rightarrow P = (-1)^L$$

$$S_n(\vec{r}) = 3 \vec{\alpha}_1 \cdot \vec{r} \vec{\alpha}_2 \cdot \vec{r} \cdot \vec{\alpha}_3 \vec{\alpha}_1$$

\swarrow
 (\vec{r}, \vec{r}) $\rightarrow P = (+)$

$$L=0 \rightarrow L=\downarrow$$

$$P=+ \rightarrow P=-1$$

$$\langle S_1 \hat{G} | P \rangle \neq 0$$

ϵ

$S_{12} \rightarrow$ two
powers
of r
(r_i, r_j)

$\hat{G} \rightarrow$ odd powers
of r

(r_1, r_2, r_3, \dots)

$$\langle S_1 S_{17}(\vec{r}) | P \rangle = 0$$

(one can check this explicitly)

$$\rightarrow \int d^2 \hat{r}$$

→ We go back to the deuteran

$$-\ddot{w} + 2\mu \begin{pmatrix} v_c \\ 2\sqrt{2}v_T \\ 2\sqrt{2}v_A \end{pmatrix} \begin{pmatrix} v \\ u \end{pmatrix}$$
$$+ \begin{pmatrix} 0 & 0 \\ 0 & \frac{6}{r^2} \end{pmatrix} \begin{pmatrix} v \\ u \end{pmatrix} = -\gamma^2 \begin{pmatrix} v \\ w \end{pmatrix}$$

Deuteron \rightarrow [S & D wave components]

Asymptotic behavior:

$$U(r) \rightarrow A_S e^{-\gamma r}$$

$$W(r) \rightarrow A_D e^{-\gamma r} \left(1 + \frac{3}{\gamma r} + \frac{3}{(\gamma r)^2} \right)$$

$$r \rightarrow \infty$$

Observables/quantities that depend on D-waves:

i) Normalization

$$\int_0^{\infty} dr u^*(r) = 1 \rightarrow \int_0^{\infty} dr (u^*(r) - w^*(r)) = 1$$

2) The mean square radius &
the quadrupole moment

$$r_m^2 = \frac{\langle r^2 \rangle}{4} = \frac{1}{4} \int_0^\infty r^2 (U - W) dr$$

(matter radius)

$$Q_d = \frac{1}{20} \int_0^\infty r^2 W (2\sqrt{2} U - W) dr$$

e (quadrupole moment)

$\sqrt{20}$

3) D-wave probability (probability of
the deuteron to be in D-wave)

$$P_D = \int_0^{\infty} dr w^2(r) \leq 0.03 - 0.05$$

(3-5% probability)

→ not an observable (actually)

Experimental values:

$$r_m = 1.9754(9) \text{ fm}$$
$$\alpha_s = 0.2859(3) \text{ fm}^2$$

} \rightarrow fitting a J_0(r) form
in an e.m. field

$D_D - (3-5)\%$ → not really observable

$$\eta = \frac{\Delta D}{D_s} = 0.6256(4) \rightarrow$$

From scattering
at low energies

TENSOR FORCE \mapsto CHANGES THE SCATTERING AMPLITUDES

We have to take spin & angular momentum
into account



[SCATTERING WAVE FUNCTION]

- 1) NO SPIN $\rightarrow \psi(\vec{r}) \rightarrow e^{i\vec{k} \cdot \vec{r}} + f(\omega) \frac{e^{i\vec{k} \cdot \vec{r}}}{r}$
 $v \rightarrow \infty$
- $$\frac{d\sigma}{d\Omega} = |f(\omega)|^2$$
- 2) SPIN (NO/TENSOR FORCE)
 $\psi(\vec{r}) \rightarrow e^{i\vec{k} \cdot \vec{r}} \left(\sum_{m_s} \langle s m_s | + \sum_{m'_s} \langle s' m'_s | \right) \frac{e^{i\vec{k} \cdot \vec{r}}}{r} \left(\sum_{m'_s} \langle s' m'_s | \right)$
- Same s

$$\frac{\partial \sigma}{\partial \omega} = |f_{m_1 m_2}(\omega)|^2$$

3) SPIN (W/TENSOR FORCE IN NN)

$$f(\vec{r}) \rightarrow e^{ik_r l s_m} + \sum_{m'} f_{m' m}(\omega) e^{-ik_r l s_{m'}}$$

$s=0, 1$ (no change in $f(\vec{r})$ asymptotic form)

2) & 3) different once we consider the partial wave expansion

No spin

$$f(\omega) = \sum_l (2l+1) f_l P_l(\cos\theta)$$

$$f_l = \frac{e^{iz_l \delta_\theta}}{2ik} - 1$$

$\Rightarrow S$ -matrix
(in partial waves)

SPIN (W/O TENSOR FORCE) |

$$f_{es}(k) = \frac{e^{2i\delta_{es}(k)} - 1}{2ik}$$



more coefficient here

diagonal in ℓ

SPIN w/ TENSOR FORCE |

$$f_e(k) = \frac{e^{2i\delta_{e-1}}}{z_1 k} \rightarrow f_{ee'}^{(S)} = \frac{1}{z_1 k} \underbrace{\left(S_{ee'}^{(S)}(k) - \delta_{ee'} \right)}$$

The S-matrix is now a matrix

In angular momentum space

→ We are not going to explain the details
(JUST THE OUTCOME)

$$f_{\text{rel}'}^{(s)}(\kappa) = \frac{s_{\text{rel}'}^{(s)}(\kappa) - s_{\text{rel}'}}{z\kappa} \rightarrow (s)^s = 1$$

|
is an extension
of circle

For the NN case :

$$\boxed{S=3, J=L \leq 1} \quad L = J \leq 1 \rightarrow \text{mixing of partial waves}$$

(Single channel $S(k) = e^{i\delta_k(k)}$)

$$S^S(\alpha) = \begin{pmatrix} (\cos\epsilon_j - \sin\epsilon_j) & e^{i\delta_1} \\ \sin\epsilon_j \cos\epsilon_j & 0 \end{pmatrix} \begin{pmatrix} 0 & \alpha \\ 0 & e^{i\delta_2} \end{pmatrix} \begin{pmatrix} (\cos\epsilon_j \sin\epsilon_j) \\ -\sin\epsilon_j \cos\epsilon_j \end{pmatrix}$$

Uncoupled \rightarrow Coupled channel case

$$\begin{pmatrix} e^{j\zeta_1} \\ e^{j\zeta_2} \\ \dots \\ e^{j\zeta_n} \end{pmatrix}$$

N phase shifters

$$R \begin{pmatrix} e^{j\zeta_1} \\ e^{j\zeta_2} \\ \dots \\ e^{j\zeta_n} \end{pmatrix} R'$$

rotation matrix

$\frac{n(n-1)}{2}$ angles $\frac{n(n+1)}{2}$ phases

NN scattering {

$S=0$ singlets \rightarrow normal
scattering
+ theory

$S=1$ triplet

a different definition of phase shift,

(that includes a series of
mixing angles)

→ It's complicated (you only need
to understand the general idea)

→ Even more complicated because

[
 └ **3 DIFFERENT DEFINITIONS
OF THE COUPLED-CHANNEL
PHASE SHIFTS** ┘

→ It's just how the field developed

Two phase shift definitions:

3) Eigen phase shifts (Blatt-Biedenharn
phase shift)

This idea of a rotation

$$S = \begin{pmatrix} \cos\epsilon_j & \sin\epsilon_j \\ \sin\epsilon_j & -\cos\epsilon_j \end{pmatrix} \begin{pmatrix} e^{i\delta_{j-1}} & 0 \\ 0 & e^{i\delta_{j+1}} \end{pmatrix} \begin{pmatrix} \cos\epsilon_j & \sin\epsilon_j \\ -\sin\epsilon_j & \cos\epsilon_j \end{pmatrix}$$

2) Nuclear-bar phase shift, (Stapp-Ypsilantis)

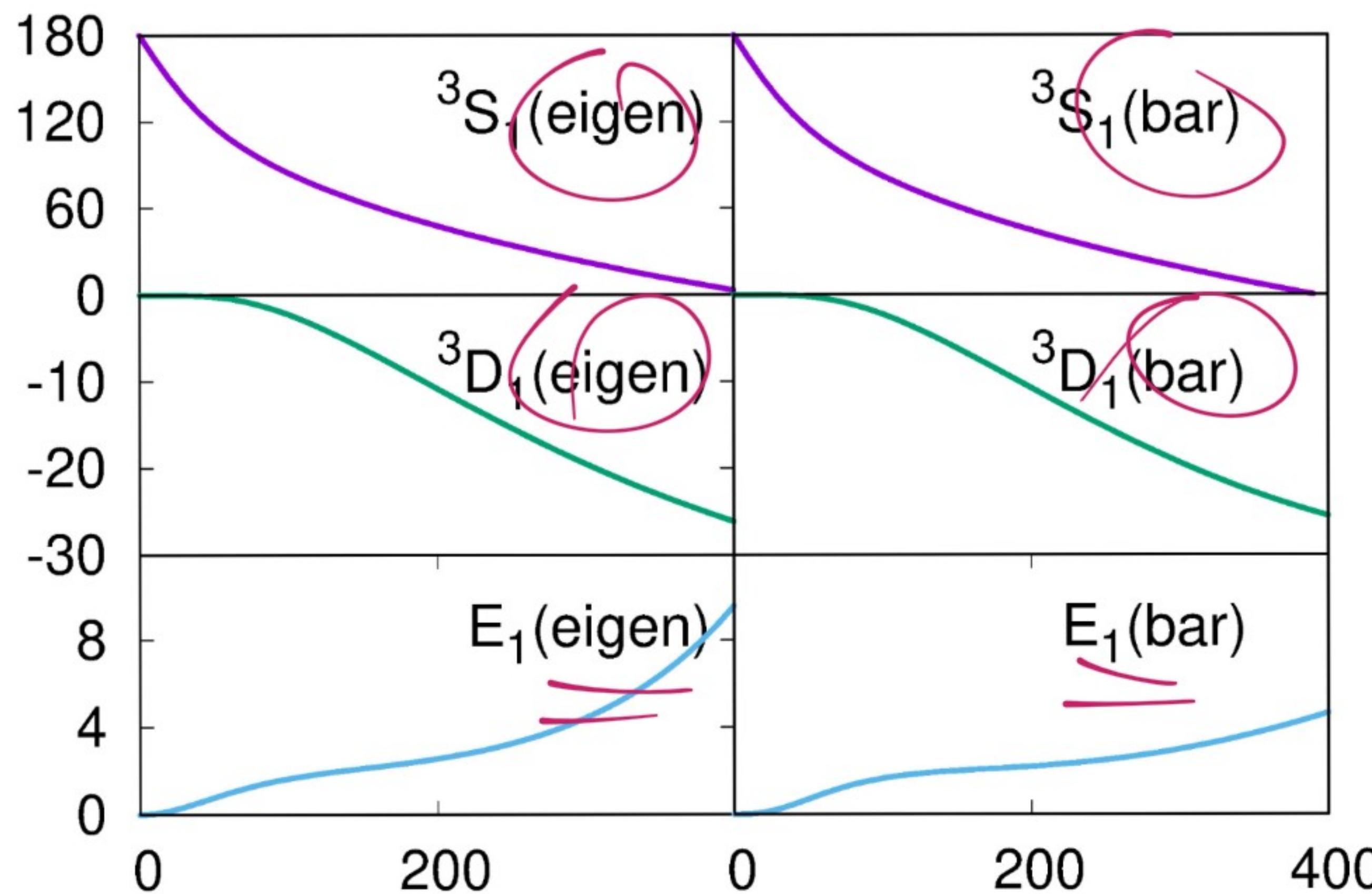
- Metropolis phase

$$\text{shift}(\Delta_s) \\ i \sin(2\bar{\epsilon}_j) e^{i(\bar{\delta}_{lj} + \bar{\delta}_{rj})}$$

$$S = \begin{pmatrix} \cos(2\bar{\epsilon}_j) e^{-i\bar{\delta}_{lj}} & \\ & \cos(2\bar{\epsilon}_j) e^{-i\bar{\delta}_{rj}} \\ i \sin(2\bar{\epsilon}_j) e^{i(\bar{\delta}_{lj} + \bar{\delta}_{rj})} & i \sin(2\bar{\epsilon}_j) e^{i(\bar{\delta}_{lj} + \bar{\delta}_{rj})} \end{pmatrix}$$

- two most common, use parametrizations
- don't worry if you don't understand this (this requires a heavy investment of time doing all the coupled-channel scattering mults) → Pictures nice!

EXAMPLE : TRIPLET $J=1$ S S D-WAVE PHASES



(Nijmegen II
potential)

Small II

mixing angles
(ϵ_J , $\bar{\epsilon}_J$)

Summary

- TENSOR FORCE IS IMPORTANT PART OF NUCLEAR FORCE
- IT CREATES SOME TECHNICAL COMPLICATIONS (IN CASE YOU DO NW, YOU WILL HAVE TO MASTER THESE COMPLICATIONS)

[ONE BOSON EXCHANGE (OBE) MODEL]

→ this is easy stuff

GBC MODEL → First qualitatively successful
description of the nuclear
force

problem addressed

WHAT IS THE ORIGIN OF NUCLEAR FORCES?



1) Before QCD:

1.a) Pion theories (SO's) \rightarrow failed

1.b) OBE model \rightarrow succeeded

2) Post QCD:

2.a) Quark-model related

2.b) Effective field theory methods

(ongoing)

Why did people propose the OBE model?

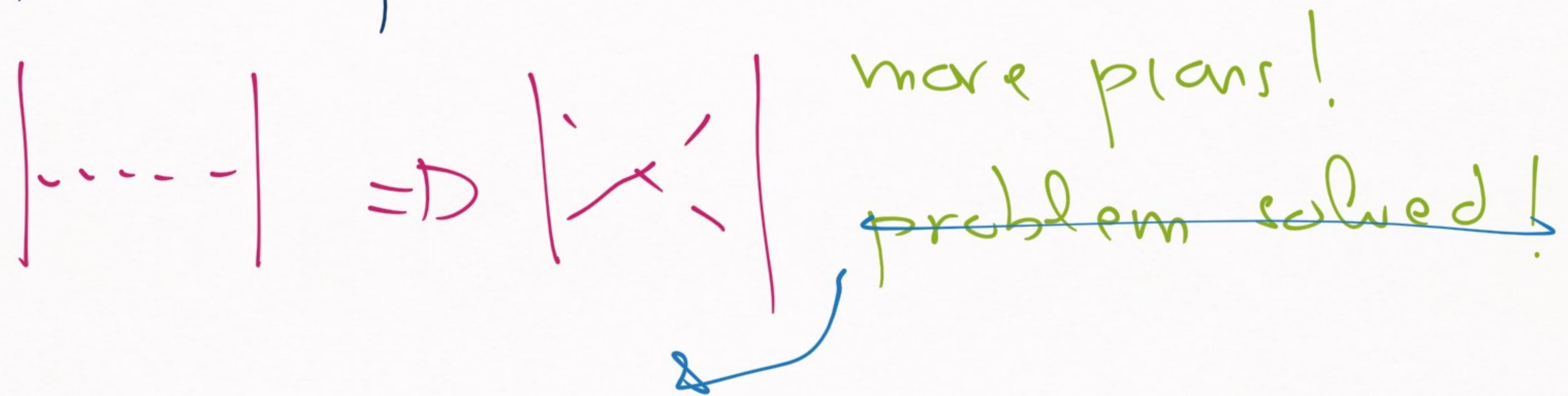
1) Yukawa's idea \rightarrow one pion exchange (OBE)

$| \sim \sim - | \pi \rightarrow$ first step in describing NN

2) How to extend OBE?

\hookrightarrow this will lead to problems

2.a) Multi-plan theories

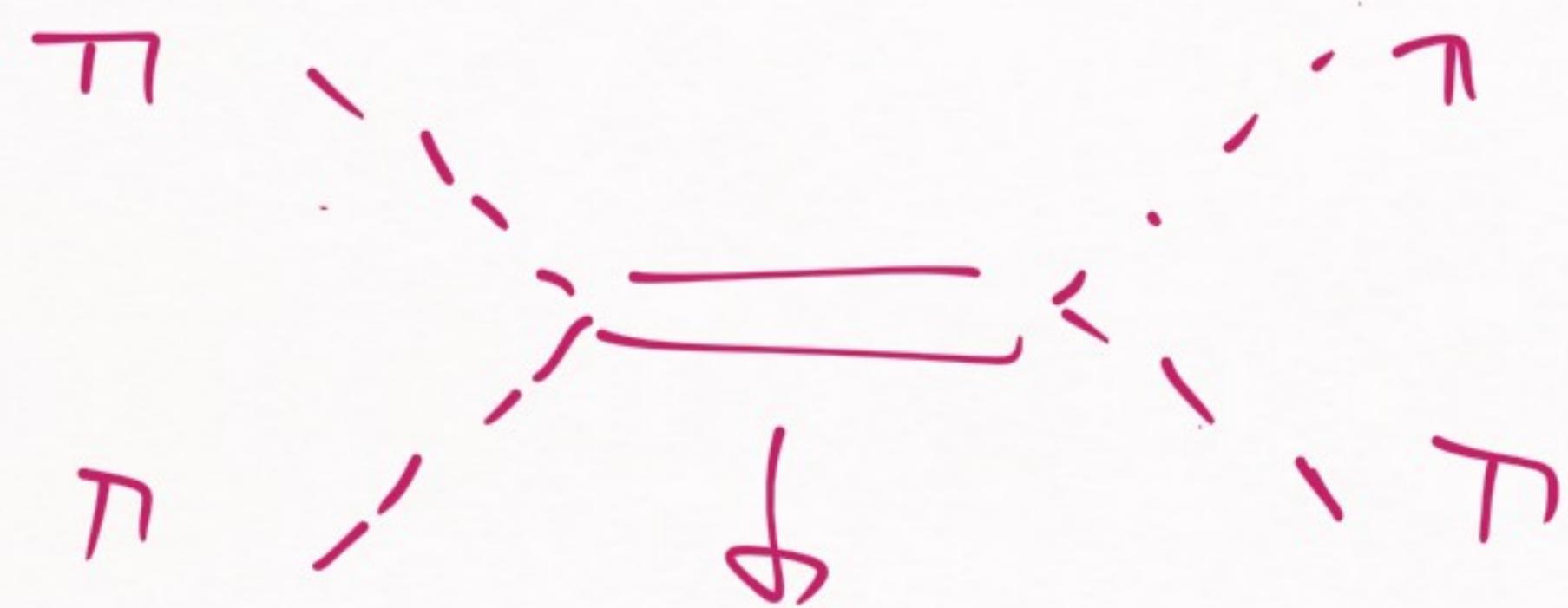


In the 50's we didn't understand

plan dynamics

(no chiral symmetry, no renormalization)

2.b) $\pi\pi$ scattering \rightarrow resonances



resonance

$(\sigma, \epsilon, \omega)$

→ inspired to
consider this
idea in NN



ΣD



ρ, ω, δ

FJDE Δ

→ substitute mult.-pion
exchange<
by the exchange
of resonance
appearing
in $\pi\pi$ scattering

3) THE FINAL RESULT IS VERY INTUITIVE

$$\left| \begin{array}{c} \dots \\ \vdots \\ \pi \end{array} \right| \rightarrow \left| \begin{array}{c} \dots \\ \pi, \delta, e \\ \omega, .. \end{array} \right|$$

Yukawa's
idea (π)

Natural extension of
Yukawa's idea

What are the important bosons in OBE?

1) The pion (π): $J^P = 0^-$, $J = 1$, $m_\pi = 140 \text{ MeV}$

(explains the deuteron's quadrupole moment)

2) The sigma (σ): $J^P = 0^+$, $J = 0$, $m_\sigma = 500 \text{ MeV}$

(strong mid-range attraction)

3) The rho (ρ) : $J^P = 1^-$, $I = 1$, $m_\rho \approx 770 \text{ MeV}$

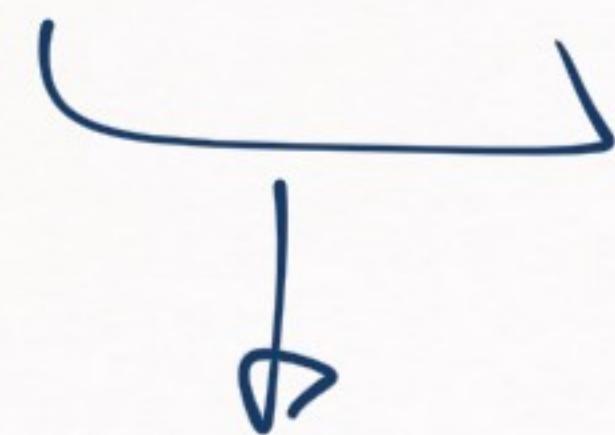
(counters the excessive strength of the pion's tensor force at short distance)

4) The omega (ω) : $J^P = 1^-$, $I = 0$, $m_\omega \approx 780 \text{ MeV}$

(provides short-range repulsion)

The OBE potential:

$$V_{OBE} = V_T + V_C - V_W + V_F + \dots$$



Really simple description
of the nuclear force



optionally
we can add
more bosons



$V_\pi(\vec{q}) \rightarrow$ you already know

$$V_\sigma(\vec{q}) = -\frac{g^2}{g^2 + m_\sigma^2}$$

$$V_\omega(\vec{q}) = \left(\frac{g_\omega^2}{g^2 + m_\omega^2} - \frac{(f_\omega + g_\omega)^2}{4M_\omega^2} \right) \frac{(\bar{\sigma}_1 \cdot \vec{q}) \cdot (\bar{\sigma}_2 \cdot \vec{q})}{g^2 + m_\omega^2}$$

$$V_\rho(\vec{q}) = \bar{\epsilon}_1 \cdot \bar{\epsilon}_2 \left[\begin{matrix} f_\omega \rightarrow g_C & m_\omega \rightarrow m_\rho \\ f_\omega \rightarrow f_\rho & \text{(same potential form)} \end{matrix} \right]$$

Traditional simplifications:

$$\left. \begin{array}{l} f_e \gg g_c \Rightarrow g_e \approx 0 \\ f_w \ll g_w \Rightarrow P_w \approx 0 \end{array} \right\}$$

Sometimes
people use
these

OBC Model | \rightarrow exhibits singular potentials

$(V_T, V_E, V_W \propto 1/r^3)$

Renormalizations was not yet
properly understood

\rightarrow Form Factors

Form factors | \rightarrow take into account
the finite size of hadrons

$$V_\pi, V_\rho, V_\omega \rightarrow \left. \frac{1}{r^3} \right|_{r \rightarrow 0} \rightarrow \text{only true if hadrons are point-like}$$

$$\rightarrow \text{finite } \left. \frac{1}{r^3} \right|_{r \rightarrow 0} \rightarrow \text{finite-sized hadrons}$$

→ NUCLEAR PHYSICISTS DECIDED TO INCLUDE
FORM FACTORS TO REGULARIZE V_{03F}

$$V_M(\vec{g}) \rightarrow V_M(\vec{g}) F_M^2(\vec{g})$$

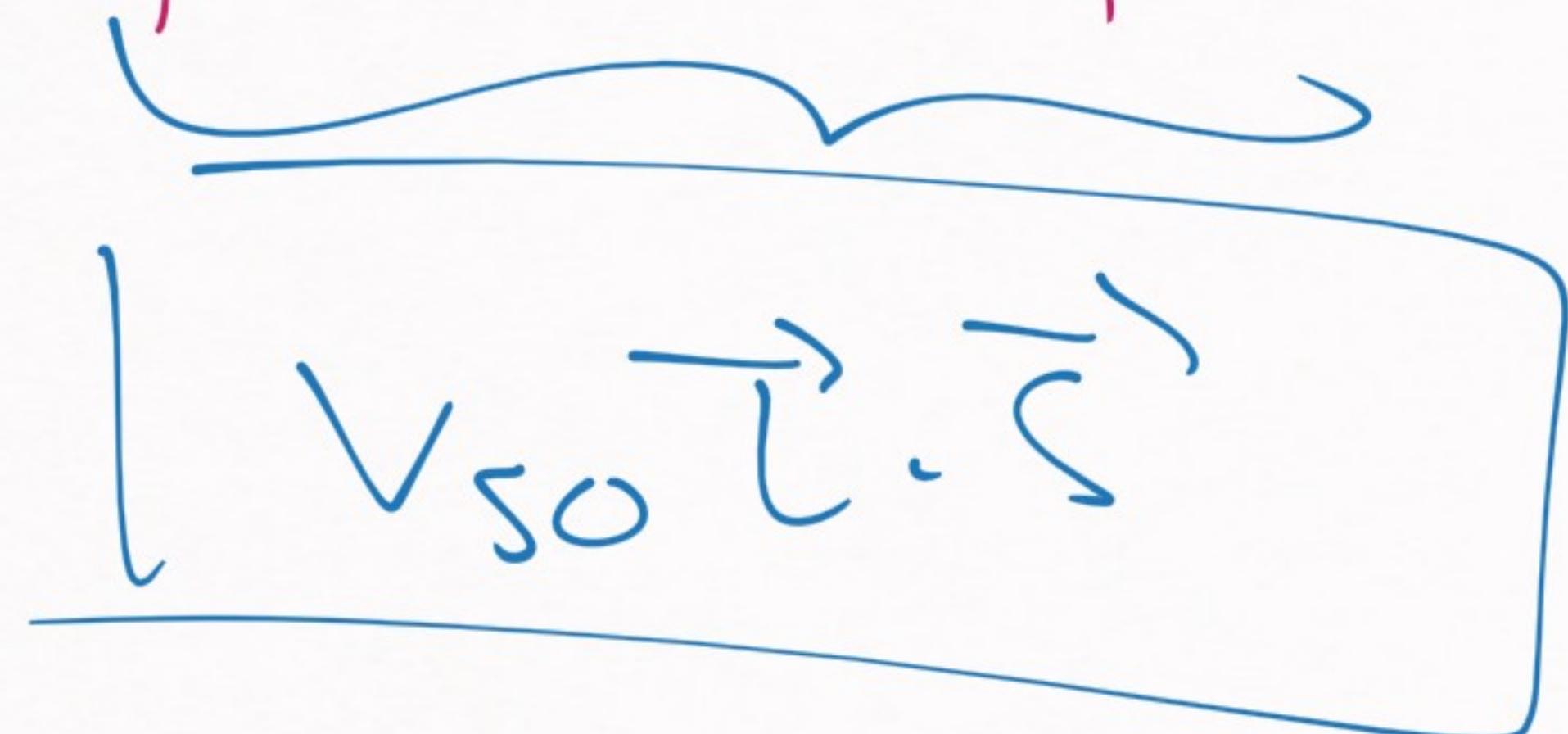
M = $\pi, \sigma, \rho, \omega, \dots$

$$F_M(\vec{g}; \Lambda) = \left(\frac{\Lambda^4 \cdot M^2}{\pi^2 + \vec{g}^2} \right)^{\alpha}$$

MULTIPOLEAR FF
(POWDER CHOICE)

→ [RELATIVISTIC EFFECTS]

Some particle waves require a strong spin-orbit potential



NON-RELATIVISTIC POTENTIALS

DO NOT HAVE THIS

Relativistic Corrections

$\rightarrow (\sigma, \omega)$ \rightarrow good $T\bar{T}$ force

RECD?

OBE MODEL

1) NATURAL EXTENSION OF Yukawa's IDEA

$$\left| \dots - \right| \rightarrow \left| \dots - \overset{+ \Theta, \rho, \omega}{\underline{\dots}} \right|$$

2) FORM FACTORS

Otherwise we end up w/ $1/r^3$ potentials

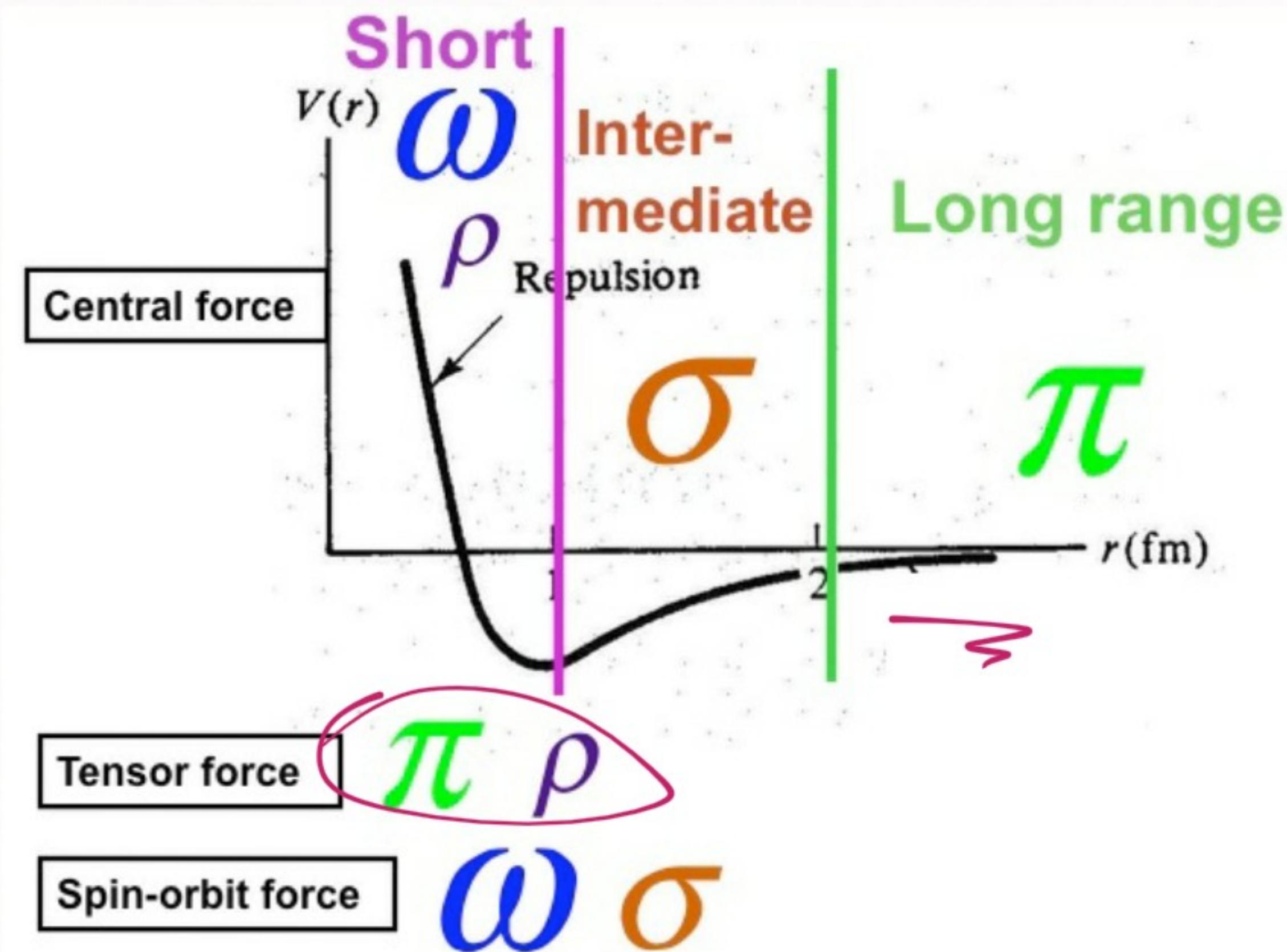
3) RELATIVISTIC CORRECTIONS

→ Spin-orbit Force

4) EACH MESON HAS A JOP

↳ Let's check this

Role of different mesons:



→ long/medium/short
range

SCHOLARSHIP ARTICLE ON NUCLEAR FORCE

Table 1: The four most important mesons and the main characteristics of their contributions to components of the nuclear force.

(BY
MKALEEM)

| Meson | Central | Spin-Spin | Tensor | Spin-Orbit |
|---------------|--|--|--|---|
| $\pi(138)$ | -- | weak, long-ranged | strong, long-ranged | -- |
| $\sigma(500)$ | strong, attractive, intermediate-ranged | -- | -- | moderate, intermediate- ranged |
| $\omega(782)$ | strong, repulsive, short- ranged | -- | -- | strong, short-ranged, coherent with σ |
| $\rho(770)$ | -- | weak, short-ranged, coherent with π | moderate, short-ranged, opposite to π | -- |

-> GOOD SUMMARY (64 & 65)

R. Machleidt et al., *The Bonn meson-exchange model for the nucleon-nucleon interaction*

Table 5
Meson parameters used in the relativistic (energy-independent) momentum space
one-boson-exchange potential (OBEPQ)

| | $g_a^2/4\pi; [f_a/g_a]$ | $g_a^2/4\pi(k^2=0)$ | m_a [MeV] | Λ_a [GeV] | n_a |
|----------|-------------------------|---------------------|-------------|-------------------|-------|
| π | 14.6 | 14.27 | 138.03 | 1.3 | 1 |
| ρ | 0.81; [6.1] | 0.43 | 769 | 2.0 | 2 |
| η | 5 | 3.75 | 548.8 | 1.5 | 1 |
| ω | 20; [0.0] | 10.6 | 782.6 | 1.5 | 1 |
| δ | 1.1075 | 0.64 | 983 | 2.0 | 1 |
| σ | 8.2797* | 7.07 | 550 * | 2.0 | 1 |

Nuclear mass: $m = 938.926$ MeV. For notation and empirical values see table 4.
Here, there are NN vertices only.

* The parameters for the σ -boson given in the table apply only to the $T=1$ NN potential. For $T=0$ we have: $m_\sigma = 720$ MeV, $g_\sigma^2/4\pi = 16.9822$ and $\Lambda_\sigma = 2$ GeV. The parameters for the other mesons in the table are the same for $T=0$ and $T=1$.

$a_0(980)$

→ USUAL
PARAMETERS
IN OBE
MODEL

THE END

(FOR TODAY)